

SAO ISAGEX EXPERIENCE. I. DATA ACQUISITION

Edited by E. M. Gaposchkin

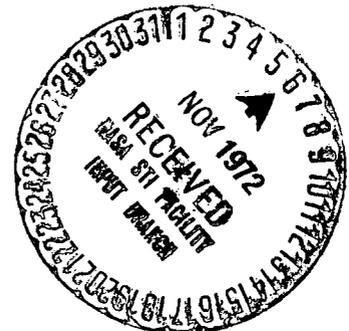
(NASA-CR-128396) ISAGEX (INTERNATIONAL  
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Smithsonian Institution  
Astrophysical Observatory  
Cambridge, Massachusetts 02138

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**ABSTRACT**

The International Satellite Geodesy Experiment (ISAGEX) has completed the data acquisition phase. This report describes the contributions and methods of the Smithsonian Astrophysical Observatory to the program. The report will provide users of the data with necessary supporting information. A sequel will be prepared when the analysis of the ISAGEX is completed.

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## INTRODUCTION

E. M. Gaposchkin

The ISAGEX program is the third in a series of cooperative satellite-tracking campaigns. The first in 1967 and the second in 1968, organized by SAO, were primarily camera tracking programs. There were, respectively, four and five laser tracking instruments operating during those intervals. Where possible, the tracking schedules were established to accommodate these systems. In 1967, there were five satellites and in 1968, there were six satellites suitably equipped with corner reflectors. However, three of them were in almost identical orbital configurations, so for some purposes the number was, in reality, three and four.

ISAGEX was initiated in 1969 by the French CNES with its "Proposition for an International Laser and Photographic Observation Campaign on Satellites Equipped with Laser Reflectors." ISAGEX took on added importance with the increased number of laser systems (10) and of precision satellite-tracking cameras (30) and the launch of a seventh retroreflector satellite (Peole), by France.

There are many purposes of a tracking campaign, and the archives of data will be useful for applications not envisaged at the inception of the program. The cooperating groups proposed the following three broad objectives, which gave shape to the program:

A. To organize a well-coordinated tracking campaign of the seven satellites equipped with laser retroreflectors in such a way that its contribution to our knowledge of the gravity field of the earth and other geodetic parameters will be significant.

B. To collect the set of observations made by the different participating agencies and to make those data available to the scientific community with all information necessary for use in computations.

C. To further research and development of instrumentation and operations of high-precision tracking systems for future space experiments.

ISAGEX is primarily a program of coordinated observations and data exchange. The data are to be distributed to all participants as per the operations plan, and subsequent analysis is largely at the option of the individual agencies. The analysis objectives of the participants and others are given in the International Satellite Geodesy Experiment Plan, published by CNES on November 10, 1970.

The planning and execution of the program has been documented in several CNES reports. The program ran from January through August 1971 and, broadly speaking, all the objectives were met. The data reduction has been completed for the laser data, which have been forwarded to the CNES data bank. The reduction of photographic observations is now under way. Therefore, the first objective and part of the second have been achieved.

The purpose of this report is to describe SAO's experience and methods. Included are the information necessary to use the laser data and descriptions of the observing system, the calibration methods, the reduction methods, and the process of data validation. In addition, there is a discussion of the various aspects of data acquisition. We hope that with such a document, improvement of the data systems will be furthered. We have concentrated on the technological and operations aspects of ISAGEX and on laser tracking in general. Within the next year, a sequel will be prepared, on the scientific results to come from the ISAGEX data. Some results already in hand are reported here and elsewhere (Gaposchkin, Kozai, Veis, and Weiffenbach, 1971).

There were substantial objectives for camera observations during the campaign. Since the SAO Baker-Nunn data have already been discussed in considerable detail, this report is restricted to the SAO laser systems.

In addition to the above objectives, ISAGEX was a test bed. We have seen how successfully a multinational observing program can be carried out. This success, in the absence of any more than informal agreements, is due to the good faith and mutual interest of all parties concerned. This sort of cooperation is enormously

important for the future. With scientific objectives becoming ever more ambitious, the requirements for tracking data become more demanding. It is apparent that several groups pooling resources can achieve much more than can individuals alone. The future programs of all groups will be materially advanced by such cooperation, and it can even be argued that some programs are not feasible without it.

SAO agreed to participate with laser units that were in the process of construction. The fielding of these units and their subsequent operation taxed the resourcefulness of the whole organization. Indeed, it had to be considered an experiment to see whether such a program of fabrication, field installation, and immediate data acquisition in amounts and with a necessary precision could even be accomplished. The statistics attest to the increased volume of data as the program progressed.

ISAGEX was used as a period for improvement of the accuracy and reliability of the laser network under routine field operation. We faced the difficulties of repairing malfunctions and detecting operating problems while the observing program was continuing. This was a completely different situation from operating one or two systems, under essentially laboratory conditions, with all SAO technical personnel available at a domestic site. We had a mixed record as a result. Some problems slipped through the system until the validation process. Needless to say, the system has since been modified. In addition, studies were begun to improve the system's accuracy. Photographing the oscilloscope waveform for centroid detection was attempted on a routine basis. Analysis of these data is in progress, and preliminary results are reported here. This experiment will lead to improved signal detection and analysis in future operation.

ISAGEX was intended to provide a framework for individual agencies developing laser tracking to take data in an organized program. They could have routine predictions and an immediate evaluation of their data. This situation is very helpful for new systems. In the final accounting, only a few such new systems participated. However, they did have the benefits described, although on the whole, the data taken were too few to be geodetically significant. These systems have been in operation, and we hope they will be able to participate in a more substantial way during future programs such as EPSOC, currently being conducted by SAO.

The ISAGEX program has ushered in a new era of cooperative tracking programs. We have every indication that the laser data taken are of 1-m accuracy with 60-cm noise. We have good confirmation of the 10-m accuracy of our current geodetic tools, as well as the very real opportunity to obtain 1-m geodesy using these and other data.

Everyone at SAO and many individuals at CNES, NASA, and other organizations contributed in a substantial way. It is impossible to acknowledge them all. The contributors of this report join me in expressing our gratitude to these people. The program has achieved what it has only through such cooperation.

#### REFERENCE

GAPOSCHKIN, E. M., KOZAI, Y., VEIS, G., and WEIFFENBACH, G.

1971. Geodetic studies at the Smithsonian Astrophysical Observatory. Presented at the XVth IUGG General Assembly, Moscow, August.

## SAO NETWORK DESCRIPTION

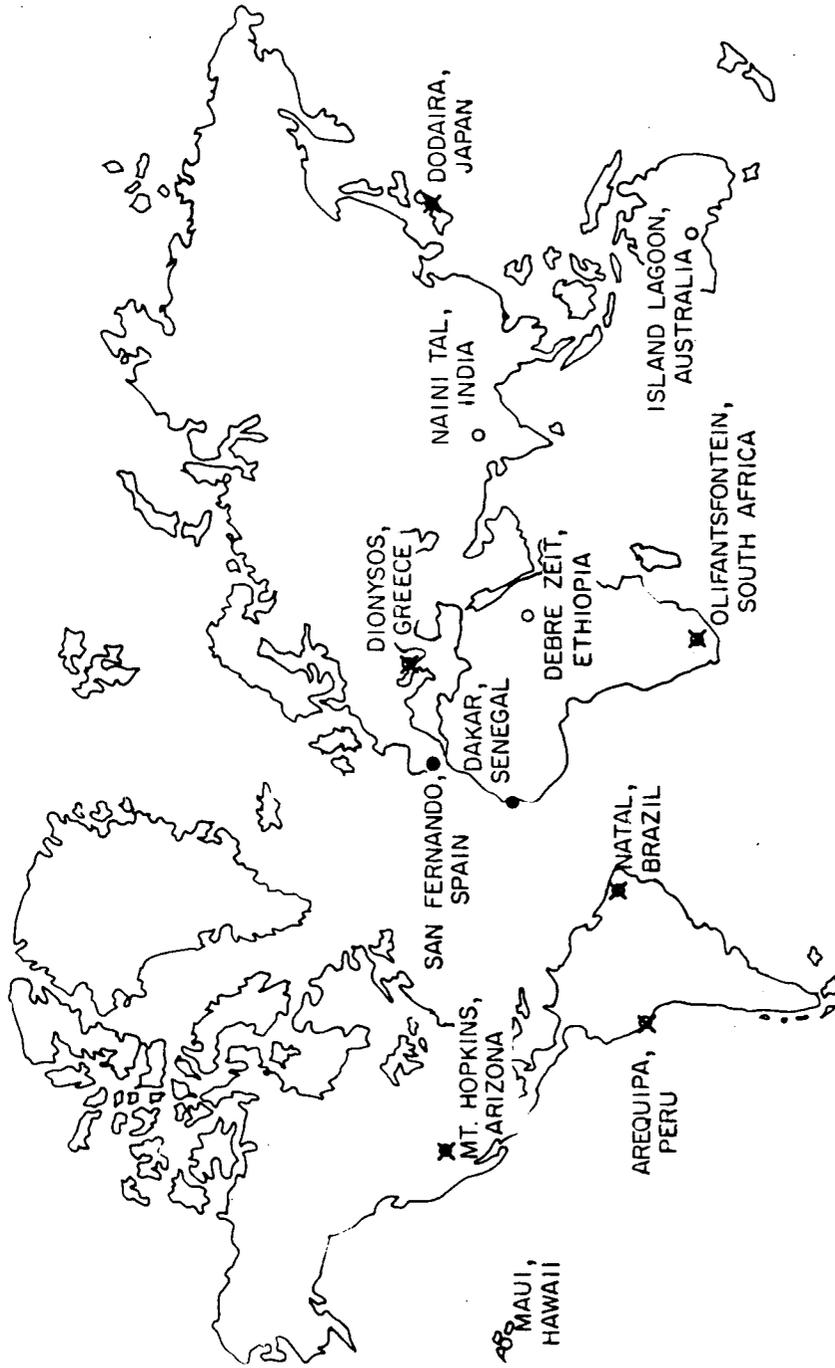
J. M. Thorp and M. A. Bush

Fifteen years ago, SAO conceived the idea of a worldwide network of photographic observing stations to track the artificial satellites proposed for the International Geophysical Year (IGY). Since the United States planned to launch its IGY satellites from Cape Kennedy into orbits with low inclinations, the original locations of the astrophysical observing stations were selected to obtain the best practical coverage of such orbits. Later, the low-latitude network configuration was modified to recognize the existence and importance of high-inclination satellites, which allow analysis of atmospheric and geodetic conditions at high latitude.

The ISAGEX network configuration is depicted in Figure 1, showing the locations of 11 astrophysical observing stations and the station in Dakar, Senegal, which was operated in cooperation with CNES. Each site is equipped with a Baker-Nunn tracking camera and a highly precise timing system. In addition, five stations have been augmented with laser ranging systems. Table 1 lists the COSPAR number and the location of the sites used in the ISAGEX program.

Figure 2 shows the Baker-Nunn camera at San Fernando, Spain, and Figure 3, the new SAO laser ranging system at Natal, Brazil.

The Baker-Nunn camera is a modified Super-Schmidt f/1, of 500-mm focal length (20 inches) and 500-mm aperture. A pyrex spherical mirror 760 mm (30 inches) in diameter and three corrector elements, two positive and one negative, constitute the optics, designed by J. G. Baker. The focal surface is approximately spherical, and the film is stretched under tension on a specially designed pyrex spherical surface. The field is 30° along the tracking axis and 5° along the perpendicular one. When the satellite position is fairly well known, the field along the tracking axis can be reduced to 15°, resulting in a considerable savings in film usage.



- SMITHSONIAN ASTROPHYSICAL OBSERVATORY SITES AND COOPERATING AGENCIES
- × SAO AND COOPERATING AGENCY LASER INSTALLATIONS
- FRENCH LASER

Figure 1. Configuration of the Baker-Nunn ISAGEX network.

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Figure 2. The Baker-Nunn camera at San Fernando, Spain.



Figure 3. The new SAO laser ranging system at Natal, Brazil.

Table 1. Sites used in the ISAGEX program.

Station		
Location	COSPAR number	Equipment
San Fernando, Spain	9004	Baker-Nunn
Naini Tal, India	9006	Baker-Nunn
Maui, Hawaii	9012	Baker-Nunn
Mt. Hopkins, Arizona	9021	Baker-Nunn
Mt. Hopkins, Arizona	7921	Laser
Olifantsfontein, South Africa	9022	Baker-Nunn
Olifantsfontein, South Africa	7902	Laser
Island Lagoon, Australia	9023	Baker-Nunn
Dodaira, Japan	9025	Baker-Nunn
Arequipa, Peru	9027	Baker-Nunn
Arequipa, Peru	7907	Laser
Debre Zeit, Ethiopia	9028	Baker-Nunn
Dionysos, Greece	9030	Baker-Nunn
Dionysos, Greece	7930	Laser
Natal, Brazil	9039	Baker-Nunn
Natal, Brazil	7929	Laser
Dakar, Senegal	9020	Baker-Nunn
Dakar, Senegal	7820	Laser

The standard film is Kodak Royal-X pan-recording 2475 (extended red) emulsion on a 4-mm estar base. The scale on the film is  $2.46 \mu \text{ arcsec}^{-1}$ , and 80% of the light is placed on a 20- $\mu$ -diameter disk. The camera can photograph stars of 14th mag with a 20-sec exposure. A barrell-type shutter, rotating in front of the focal surface

at a precise angular velocity, chops the trails of the stars (or of the satellites, if the camera is stationary) and provides the breaks that are used as references for the reduction of the film. The shutter (or chopper) rotates five times, making five breaks per exposure. When the shutter is in the middle of the central break, an electrical contact strobes a flashing tube that records the time from a slave clock in the camera. Time is thus recorded on the same film on which the satellite and stars appear. With the use of a phase shifter, the shutter can be synchronized to time signals so as to produce the middle of the central break at a predetermined time. This method can be used to perform simultaneous observations from several stations whose shutters are in full synchronization.

The Baker-Nunn can be operated in two basic modes, stationary and tracking. In the first, more simple mode, the camera is held stationary while the images of the satellite trail along the film. In the tracking mode, the body of the camera is driven at the same rate as the apparent angular velocity of the satellite, holding the image of the satellite to a point of light on the film.

The precision timing system is composed of an EECo clock that utilizes a 5-MHz Sulzer crystal oscillator as its frequency standard, a high-frequency receiver to monitor the WWV signal, and a VLF receiver. The VLF receiver monitors the very accurate frequency tones transmitted by various VLF stations around the world. The frequency of the crystal oscillator is continuously compared with these frequency tones, and the difference is displayed on both a chart recorder and an accumulated time-deviation counter. The oscillator can therefore be adjusted periodically by means of a tuning capacitor. As a further aid to more accurate timekeeping, a portable clock is carried from station to station to measure the relative settings of the clocks. Timing to within 100  $\mu$ sec is routinely achieved at all camera stations. At the laser sites, timing is maintained to within 50  $\mu$ sec, an accuracy necessary for laser ranging.

Since 1966, SAO has been improving the accuracy of its tracking technique by installing laser tracking systems at several of its camera locations. The Baker-Nunn provides very accurate directional data, and the laser provides the added dimension of range or the accurate determination of a satellite's height above the earth. The

increased accuracy of the Smithsonian tracking program as a result of the addition of lasers has innumerable research benefits.

SAO currently operates five laser tracking systems collocated with Baker-Nunn cameras at tracking stations in Mt. Hopkins, Arizona; Natal, Brazil; Arequipa, Peru; Olifantsfontein, South Africa; and Athens, Greece. With the exception of the last, which was assembled and operated in cooperation with the NTU of Athens, all systems were designed and built to SAO specifications and represent near state-of-the-art ruby-laser technology.

The following characteristics apply to the four systems operated by SAO: The type is ruby Q-switched; the peak power is 400 Mw; the pulse width is 18 nsec; the energy output is 7 J; the maximum pulse repetition rate is 4 pulses  $\text{min}^{-1}$ ; the beam divergence can be varied from 0.5 to 6 mrad (or from 2 to 20 arcmin); the pointing is automatic-static, which permits day and night ranging and does not require that the satellite be sunlit; and the range resolution of the counter is 1 nsec. The laser transmitter utilizes two ruby rods, one for the oscillator and one from the amplifier stage. The rods are stimulated by the discharge of xenon high-voltage lamps. At lasing, some of the output energy is sampled by a photodetector that triggers the range measuring counter. The pulse, after traveling in space from the transmitter to the satellite, is reflected back from the retroreflectors mounted on the spacecraft and is focused by a 20-inch Cassegrain telescope onto a photomultiplier tube. The signal generated by the photomultiplier stops the counter, which then displays the elapsed time.

One of the problems encountered in laser ranging is the variation in signal strength of the pulse reflected from the satellite. This variation is due in part to the fact that the signal varies inversely as the fourth power of the satellite range and in part to an observed "scintillation," or random effect. The variations in signal strength affect the range measurements. As we said, satellite range is obtained from a time-interval counter that is started by the transmitted pulse and stopped by the receiving pulse. The resolution of the counter is 1 nsec, but the duration of the pulse is 18 nsec. Hence, the counter reading changes significantly if it stops at different points on the pulse's leading edge. The counter stops when it reaches a threshold

that has been set near the half-amplitude point of a weak return pulse. Since the system is calibrated for such a pulse and setting, errors are introduced when the return pulse varies from its average value. These errors, however, can be corrected if a photograph of each return pulse displayed on an oscilloscope is obtained. An automatic recording system capable of doing this for every pulse has been devised and is currently being field-tested. This correction can reduce the error in range measurements to 1 ft or less.

DESCRIPTION OF THE SAO LASER SYSTEM CURRENTLY DEPLOYED  
IN BRAZIL, PERU, AND SOUTH AFRICA

P. W. Sozanski

The main components of SAO's laser system are the laser transmitter, the static-pointing pedestal, the telescope photoreceiver, the data system, and the epoch timing system (see Figure 1).

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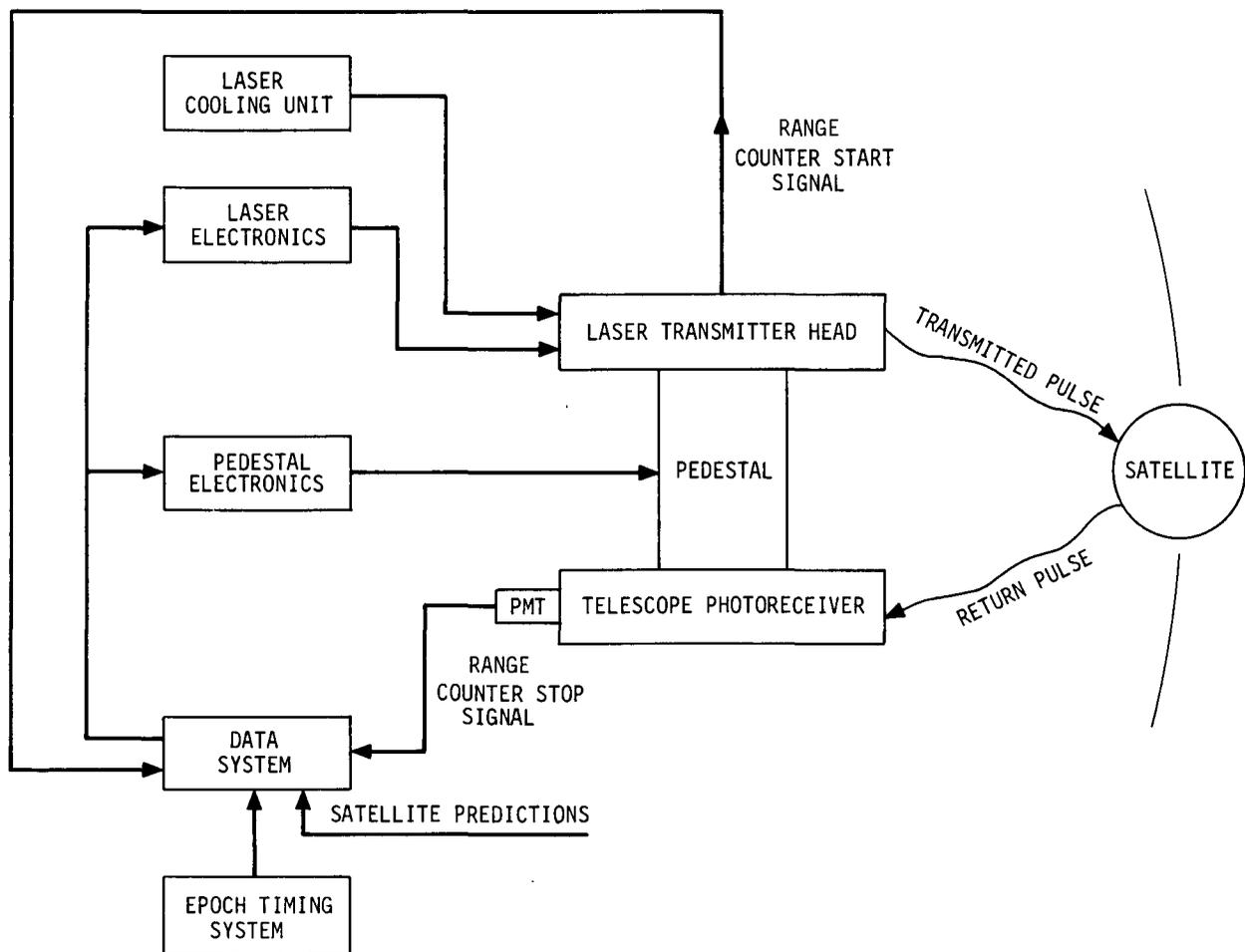


Figure 1. Laser ranging system.

To describe the operation of the laser system in simple terms, the laser transmitter head and the telescope photoreceiver are pointed by the static-pointing pedestal (see Figure 2) to the altitude and azimuth coordinates in accordance with predictions generated in Cambridge. The laser is then pulsed, under electronic or manual control, at the appropriate epoch, and a very short pulse of monochromatic light in a narrow beam is projected from the laser transmitter head toward the satellite. The transmitted pulse is detected at the transmitter by a photodiode whose output is an electrical pulse that starts the range interval counter and reads out the station clock to mark epoch. The light is reflected back from the satellite by its cube-corner reflectors and is detected photoelectrically by the telescope photoreceiver whose output is an electrical pulse that stops the range interval counter. The range from the laser system to the satellite is then calculated from the elapsed time, with due corrections for atmospheric and other effects.

Procurement started in early 1969, and the systems were fielded in late 1970.

A detailed description of the SAO laser system components follows:

#### 1. LASER TRANSMITTER

The laser transmitter was purchased in March 1969 from:

Spacerays, Inc.  
Northwest Industrial Park  
Burlington, Massachusetts 01803  
(617)272-6220

The system is a flash-pumped, Q-switched ruby system with an oscillator and one amplifier stage. The output is a 4- to 5-J pulse 18 nsec wide. The beam is collimated with a Galilean telescope with an aperture of 12.7-cm diameter, the beam divergence is variable from 0.3 to 6.0 mrad (measured at full width, half-power points), the repetition rate is 4 ppm, and the wavelength of the output is 694 nm.

The laser transmitter system consists of three major units: the laser transmitter head, the power supply and control electronic units, and the cooling unit.

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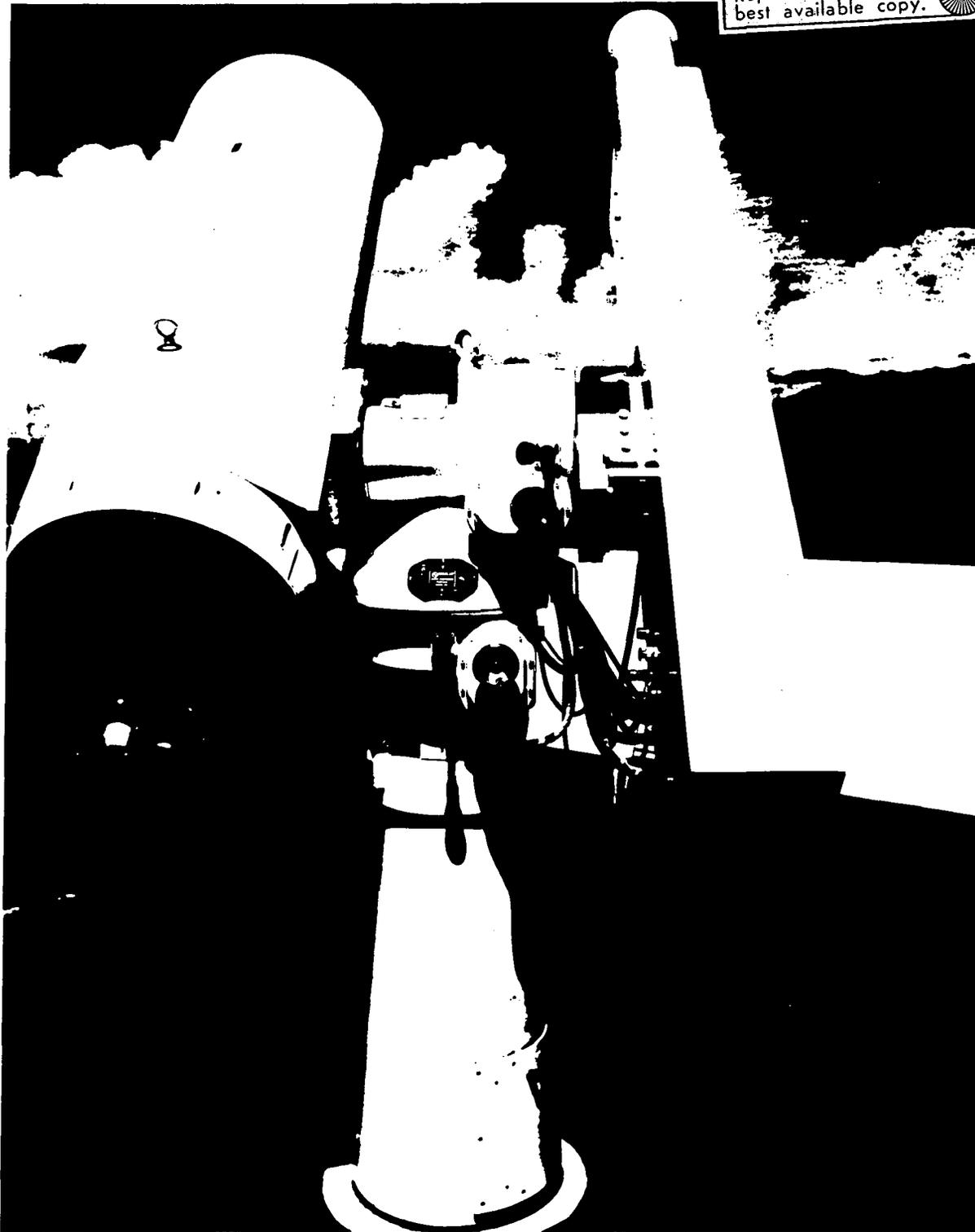


Figure 2. Laser transmitter head (right) and telescope photoreceiver (left) mounted on static-pointing pedestal.

## 2. PEDESTAL

The pedestal was purchased in May 1969 from:

Tinsley Laboratories, Inc.  
2448 Sixth Street  
Berkeley, California 94710  
(415)843-6836

This pedestal is a static-pointing, open-loop unit of an altitude-over-azimuth biaxial configuration. Its overall accuracy is within  $0^{\circ}.008$  (great circle error of 0.5 arcmin or less).

The pedestal is a static-pointing unit, i. e., it moves to a given pointing direction, waits for the satellite to pass through that direction, and then moves on to the next such static point.

The unit is of the open-loop type, i. e., it does not operate as a servomechanism and does not require a feedback error signal. It relies instead on starting at a known pointing direction of two orthogonal axes and on simple addition and subtraction of known increments of motion about those axes to arrive at a new predetermined pointing direction. The known increments of motion are provided by reliable, precision, incremental-stepping motors fed by precomputed number-of-steps input data. The initial starting position is established by optical goniometers, and the continuous addition and subtraction is maintained by solid-state arithmetic units, counting registers, comparison logic circuitry, and visual displays.

The pedestal is positioned by manual decade-switch selection or by using a pre-punched paper-tape input.

## 3. TELESCOPE PHOTORECEIVER

The telescope photoreceiver was purchased in May 1969 from several vendors (see below). It contains three subsystems (see Figure 3). The first, the main subsystem, contains two components, a 53-cm-diameter f/4 paraboloidal primary and a

14.6-cm-diameter flat secondary. The primary has at least a 50-cm-diameter clear aperture and when combined with the secondary produces a field greater than 20 arcmin in diameter, an overall accuracy of better than 1/4 wave; after aluminizing and SiO overcoating, the primary has a minimum combined reflectance greater than 60% between 400 and 700 nm.

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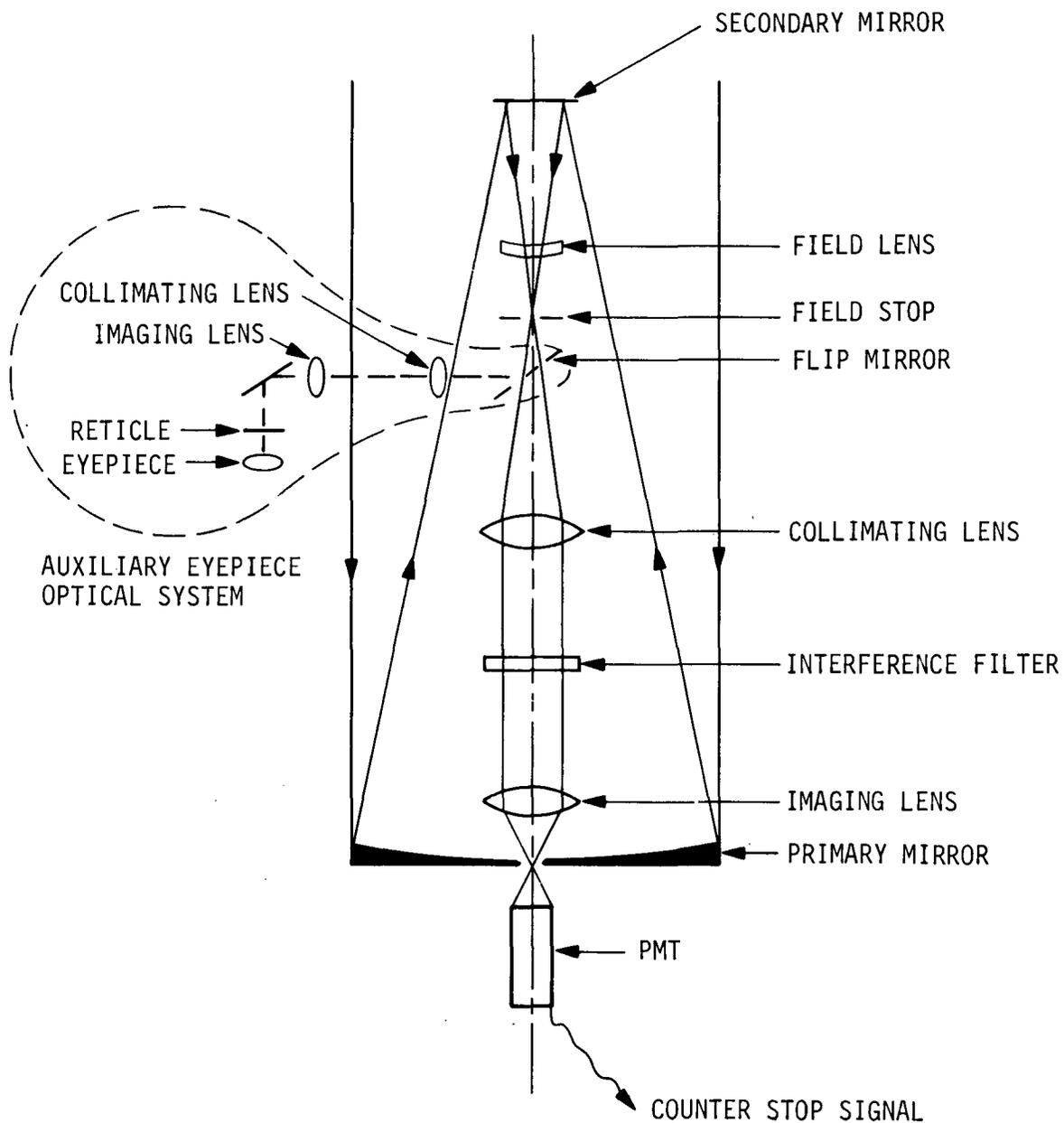


Figure 3. Telescope photoreceiver.

The second subsystem, the photomultiplier tube (PMT) optical subsystem, transfers the beam reflected by the main optical system through its components and onto the face of the PMT. The beam passes through a holder for a stack of gelatin filters, then through a field lens that directs the beam through the field stop wheel containing six apertures. These apertures are sized to produce fields of 2, 4, 8, 12, 16, and 20 arcmin in diameter. The diverging beam passes next through a collimating lens and then through a tiltable interference filter, a final imaging lens, and onto the face of the PMT. This system is designed to produce a 3.8-cm-diameter spot on the face of the PMT independent of field stop settings.

The third subsystem, the auxiliary viewing subsystem, consists of 1/4-wave flat flip mirror that, when inserted into the beam, directs the beam through the exit lenses onto a front surface diagonal mirror that reflects the beam through an illuminated reticle and out to the eyepiece for visual viewing.

A detailed breakdown of the photoreceiver is as follows:

Telescope. A 50-cm telescope is used to detect the laser return and was bought from:

Tinsley Laboratories, Inc.  
2448 Sixth Street  
Berkeley, California 94710  
(415)843-6836

Photomultiplier Tube. A RCA Model 7265, selected for a quantum efficiency of 4.5% or greater at 694 nm and a gain of  $2 \times 10^7$  or greater at 2400 v, is used to trigger the counter. It was purchased from:

Radio Corporation of America  
Industrial Tube Division  
New Holland Pike  
Lancaster, Pennsylvania 17604  
(717)397-7661

Photomultiplier Tube Housing. A Products for Research Model PR 2100 (modified) housing, used to hold the PMT, was purchased from:

Products for Research  
78 Holten Street  
Danvers, Massachusetts 01923  
(617)774-3250

Interference Filters. A 20 Å interference filter and a 7 Å interference filter were purchased from:

Thin Films Products Division  
Infra-Red Industries  
80 4th Avenue  
Waltham, Massachusetts 02154  
(617)894-8410

#### 4. DATA SYSTEM

The data system was purchased during the interval from March 1969 through September 1970 from several vendors (see below). The system (Figure 4) consists of the measurement instrumentation as well as the digital-control and data-handling systems for the laser transmitter. The four functional subsystems are described below.

Counter. An Eldorado ElectroData Model 796 (modified) counter with a 1-nsec resolution is used to obtain the satellite range times. This unit was purchased from:

Eldorado ElectroData Corporation  
601 Chalomar Road  
Concord, California 94520  
(415)686-4200

Oscilloscope. A Tektronix Type R454 oscilloscope with modification 163D and a Tektronix Model C-40 oscilloscope camera are used to provide a means for visual monitoring and photographic recording of the laser transmitter output and the laser return pulses. The oscilloscope and camera were purchased from:

Tektronix, Inc.  
P. O. Box 500  
Beaverton, Oregon 97005  
(503)644-0161

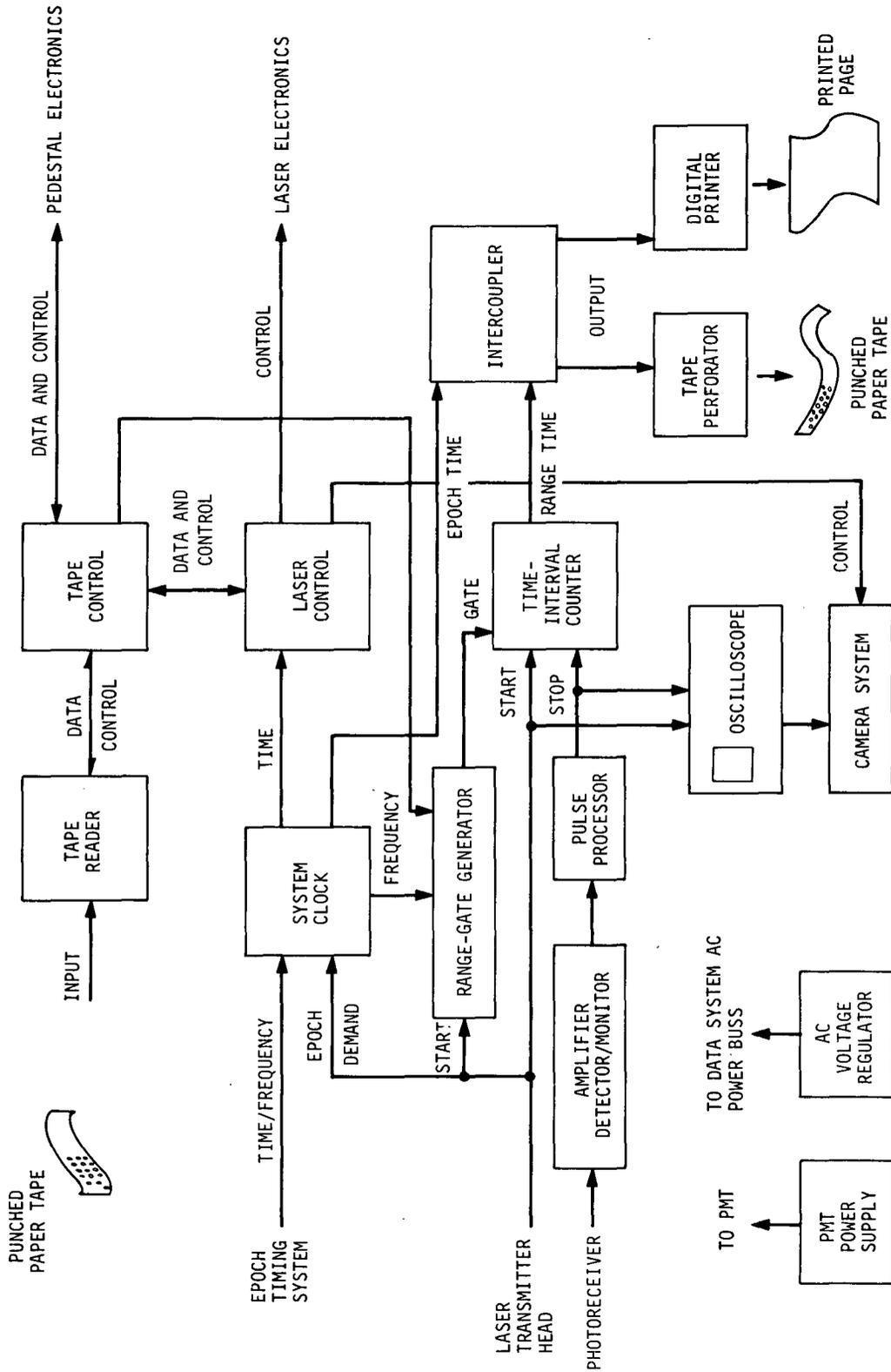


Figure 4. Laser data system.

Control System. In addition to providing the basic time-interval measurement for satellite ranging, the laser data system must also record the observation epoch time (system clock); program the operating sequence of the laser transmitter unit, the pedestal, and the data system itself (tape reader, tape control, and laser control); condition the stop channel and the return pulse (range-gate generator and amplifier detector/monitor); and print out the digital data (intercoupler, digital printer, and tape perforator). All the control-system components with the exception of the digital printer and tape perforator were built by SAO. The digital printer and tape perforator (Model ASR-32, modified by SAO) were purchased from:

The Teletype Corporation  
5555 Touhy Avenue  
Skokie, Illinois 60076  
(312)982-2000

Racks, Power, and Cabling. A Western Devices rack and blower unit, used to hold most of the data system, was purchased from:

Zero Manufacturing Company  
1121 Chestnut Street  
Burbank, California 91503  
(213)849-5521

A Bud Radio Company Model 2707 Series 60 rack is used to hold the tape-reader system and was purchased from:

Gerber Electronics  
852 Providence Highway  
Dedham, Massachusetts 02026  
(617)329-2400

A General Radio Model 1581-ALR2 voltage regulator is used to supply regulated AC power to the data system. It was purchased from:

General Radio Company  
300 Baker Avenue  
West Concord, Massachusetts 01781  
(617)369-4400

A Northeast Scientific regulated high-voltage power supply Model RQE-3001-21230 provides high voltage to the photomultiplier tube. It was purchased from:

Northeast Scientific Corporation  
30 Wetherbee Street  
Acton, Massachusetts 01720  
(617)263-7706

## 5. EPOCH TIMING SYSTEM

The epoch timing systems, a Model ZA 34675 single-channel unit and a Model ZA 34685 dual-channel unit, were purchased in March 1965 from:

Electronic Engineering Company of California  
1601 East Chestnut Avenue  
Santa Ana, California 92700  
(714)547-5651

The EECo timing system (see Figure 5) is used to provide epoch. It has a display resolution of 10  $\mu$ sec and an electrical resolution of 1  $\mu$ sec. The single-channel unit consists of a crystal oscillator, accumulator, oscilloscope, VLF receiver, chart recorder, WWV receiver, and a battery backup system. The dual-channel unit consists of the above plus an additional crystal oscillator, accumulator, and VLF receiver.

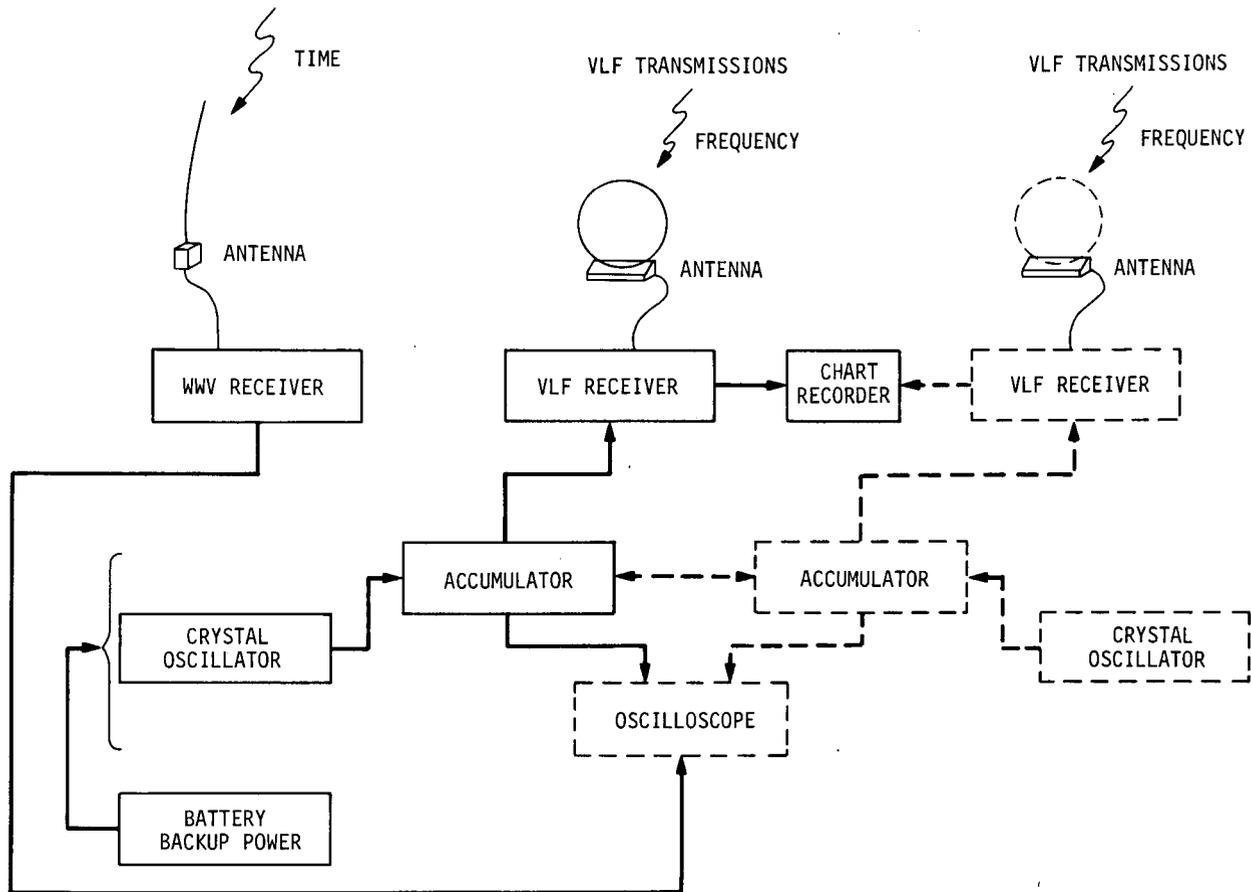


Figure 5. Epoch timing system (EECo). Solid lines show single-channel unit, while the dual-channel unit is represented by the solid lines and the dotted lines.

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## CALIBRATION

C. R. H. Tsiang

The ranging accuracy of the laser system is calibrated through a procedure of ranging on a fixed land-based target at a surveyed distance, generally on the order of 0.25 to 1 mi. A calculation can be made to obtain the expected range time based on the surveyed distance and the atmospheric refractivity. Once an average range time is obtained from a series of target measurements, it is possible to compute a calibration number  $\tau_c$ , which can be reported along with the satellite range times. This number covers delays in the range counter, cabling, telescope, output detector, photomultiplier tube, and signal amplifier. It does not provide a means for obtaining a calibration factor for atmospheric delays, but otherwise accounts for all components in the range measuring path to and from the satellite.

The following formula can be used in computing the atmospheric refractivity N:

$$N = 80.29 \frac{P}{T} - 11.9 \frac{e}{T} ,$$

where P is the measured barometric pressure (in millibars), e the partial pressure of water vapor, and T the temperature (K). The calculation for two-way range time has been based on

$$\tau_s = \frac{R_s}{0.15} (1 + N \times 10^{-6} + 6.917 \times 10^{-4}) ,$$

where  $\tau_s$  is the calculated range time (nsec) for a surveyed distance of  $R_s$  (m). After noting that local temperature and pressure variations at any one location never change N by more than  $\pm 10\%$  and that the range equation gives subnanosecond variations in  $\tau_s$  for such changes, we decided that fixed values of N could be determined for each station. Rather than each station calculating its mean barometric pressure, we prepared a chart, which gives a direct conversion from the station altitude above the geoid in kilometers to values of N (see Figure 1).

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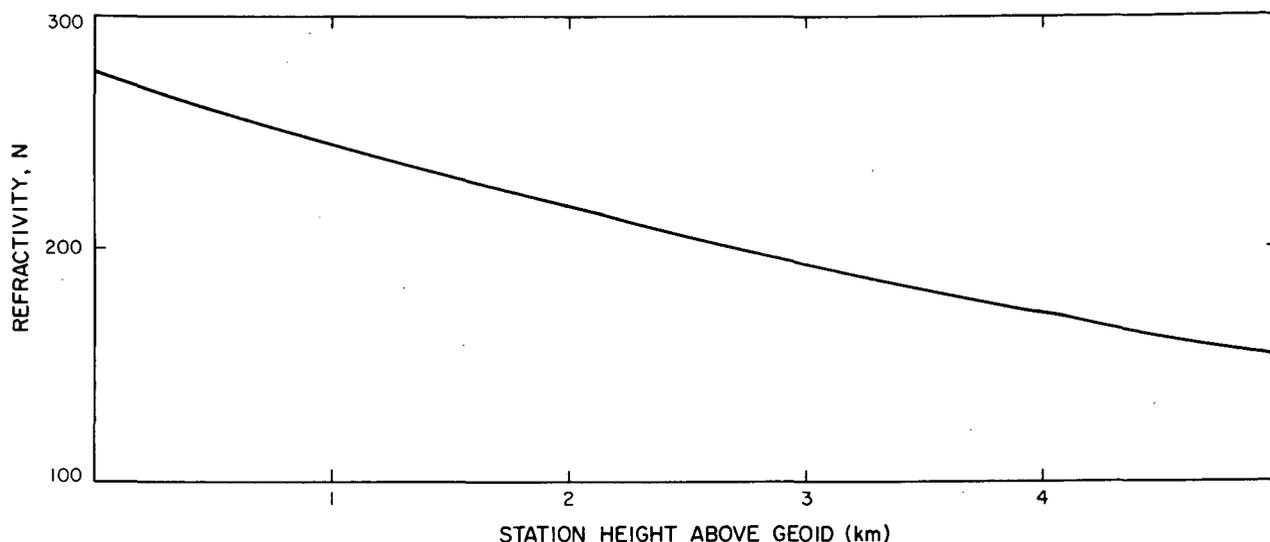


Figure 1. Atmospheric refractivity as a function of station height for an ambient temperature of 15°C at 6943 Å.

The system-calibration number reported in word eight of the 33333 observation message is obtained by subtracting the observed range time  $\tau_m$  from  $\tau_s$ , the range time calculated from the surveyed laser-to-target range:

$$\tau_s - \tau_m = \tau_c$$

The resulting system-calibration number  $\tau_c$  is reported as a signed quantity, which is added to all range measurements in the 33333 message. Generally,  $\tau_c$  is negative in the SAO laser systems.

In theory, the calibration of the instrument should change only if its components are changed, moved, or affected by environmental fluctuations or aging. By using the whole system to range on a fixed land-based target, we hoped that all such factors could be covered by the single system-calibration number. Attempts were made to simulate real operational conditions by regulating the pulse-repetition rate, photomultiplier-tube voltage, counter thresholds, amplifier gain, return signal level, output power level,

etc. Several unavoidable differences existed between the ranging measurements on the satellites and those on the target – viz., corner cubes vs. nonspecular reflecting surface of the target; small solid angle subtended by the satellite vs. full-beam reflection by the target (8-ft  $\times$  8-ft wooden surface painted flat white); point-source satellite image vs. off-axis, near-field reflection by the target; short air-path length to ground-based target vs. full atmosphere to satellite, etc. None of these differences is trivial, but for operation of the laser at the original design levels (0.5-m resolution and 1-m accuracy), the calibration procedure and results appear to be satisfactory.

For the three stations with the new lasers, the variations in the reported calibration number over the course of most of ISAGEX did not exceed 15 nsec. All the new stations did experience some problems with calibration ranging when first set up, but they were confined to the first week of operations and apparently were cleared up by the beginning of the second ISAGEX period. For the most part, the variations in the reported calibrations were due to the following:

- A. Replacement or relocation of components in the laser head, photoreceiver, or data system.
- B. Readjustment of signal operating levels in the target-ranging procedure.
- C. Readjusted survey figures for the laser-to-target distance.

Attempts at standardizing and improving the reliability of the calibration tests were made throughout ISAGEX as more experience and knowledge of the instrumentation was gained. During investigations into the error-reducing capabilities of photographically determined range-time corrections, certain observations led to the establishment of procedures that would minimize the effects of variations in signal level during target observations. These effects – the most detrimental ones for satellite ranging – were reduced to a level so that the most significant systematic error component of the calibration lay in the accuracy of the ground-survey information. Therefore, where survey information was questionable, additional measurements were made in an attempt to allow no more than a 4-cm error.

Unfortunately, this has led to the establishment of more than one survey target distance at two of the South American stations. Table 1 shows the results obtained by local surveyors using conventional techniques. Note that certain sites have been measured to greater resolution than others. Only where special comments are included should this be considered significant. Ultimately, corrections will be applied to all ISAGEX calibration data when the stations are resurveyed with a laser geodimeter.

Apart from the systematic error contributed by the differences in survey results, the net range-time uncertainty introduced by the calibration should generally be better than  $\pm 2$  nsec. Photographic reduction of pulse images offers the possibility of decreasing the error to less than  $\pm 1$  nsec for many of the periods after May 1971. Reduction of these photographic data and distribution of the results will be made in 1972.

Table 1. ISAGEX target range history: Effective dates of change to new values of survey distance and refractivity constant N.

Station				
Location	COSPAR number	Date (1971)	Survey distance (m)	Refractivity constant N
Arequipa, Peru	7907	January 5	313.135	250
		October 1	312.844	250
Mt. Hopkins, Arizona	7921	January 5	776.329	172
Olifantsfontein, South Africa	7902	February 12	404.48	290
Natal, Brazil	7929	January 5	316.08	248
		April 28	316.08	340
		April 29	314.67	340
		September 23	315.62	340
Dionysos, Greece	7930	January 5	327.25	0
		April 9	327.973*	0
		April 28	327.973	340

\* Measurement made by laser geodimeter.

## TIMING (EPOCH)

D. A. Arnold and J. M. Thorp

Epoch time is maintained at each station by use of the EEC<sub>o</sub> precision time system (see Thorp and Bush, this volume) and by reference to UTC(USNO). A portable clock is used to set the station clock, which is then maintained by referencing the frequency of a 5-MHz Sulzer oscillator to a known frequency, broadcast by one of the various VLF stations. Each observing station maintains an estimate of its timing uncertainty in two ways: First, the accuracy of the original clock set from a portable clock is expressed as an uncertainty (usually  $\pm 5 \mu\text{sec}$ ). Time is maintained at each station on one main channel, with one or more alternate channels keeping time independently for backup. If the main channel has to be reset to one of the backup channels, an additional uncertainty is added (usually  $\pm 5 \mu\text{sec}$ ). The second uncertainty is the deviation of the oscillator caused by its drift in frequency. The oscillator drift is determined by comparing the phase of the VLF station with that of the oscillator at a particular time each day. Each station steers or guides its oscillator to keep its time-drift uncertainty as small as possible (usually  $\pm 50 \mu\text{sec}$ ).

In addition to the above uncertainties, two sets of time corrections are added in order to have time equivalent to UTC(USNO). One set consists of corrections of hours, minutes, seconds, or parts of seconds when a failure has occurred in the main time-keeping channel. These corrections are confirmed by referring to the alternate time-keeping channel and the WWV time signals. If all channels fail, time reference is lost and a reset is necessary. The second set of corrections is added in Cambridge; this consists of the computed phase differences between the average VLF phase for a period (usually a month) and the phase of the VLF at the time the clock is set. These corrections, determined from data published in USNO time-service bulletins, are generally on the order of  $< 20 \mu\text{sec}$ .

Two files of time corrections are maintained by the Data Services Division at SAO. The first gives the difference between A.S and UT1, and the second, the difference between A.S and the clocks at the observing stations. The time system A.S is related to UTC(USNO) by the expression

$$A.S - UTC(USNO) = 6^S.140768 + 0.002592000(T - 39856.0)$$

for the period February 1, 1968, to January 1, 1972; T is the time in Modified Julian Days; 39856.0 is January 1.0, 1968; and the difference is given in seconds. The A.S - A.1 difference is about 0.8983 msec.

UT1 data are obtained from "Circular D," published monthly by the BIH. Values of UT1 - UTC(BIH) and AT - UTC(BIH) are listed at 5-day intervals. The difference A.S - AT is currently 35.3 msec. A.S - UT1 is calculated by the relation

$$A.S - UT1 = (A.S - AT) + [AT - UTC(BIH)] - [UT1 - UTC(BIH)] \quad .$$

A second-order polynomial is fitted to the A.S - UT1 values, and the coefficients are punched on cards. Usually, each polynomial covers a 50-day period. If the values change too rapidly, the interval can be reduced to 25 days.

The difference between the station clocks and UTC(USNO) is recorded by STADAD as described. The corrections are added to the A.S - UTC(USNO) difference to obtain the correction from the station clock to A.S time. Cards are punched giving these corrections as a series of straight-line segments specifying the values of the corrections at the beginning and end of each interval. A new card must be used whenever there is a gap, discontinuity, or change of slope in the time correction.

## ATMOSPHERIC REDUCTION OF LASER DATA

C. G. Lehr

Laser ranges determined by using the value of the velocity of light in a vacuum must be corrected for the fact that the laser pulse travels at a lower velocity during its passage through the earth's atmosphere. The correction is currently made by means of the following formula (obtained in a personal communication from Gordon D. Thayer):

$$r_m = r_v - \frac{2.238 + 0.0414 PT^{-1} - 0.238 h_s}{\sin \alpha + 10^{-3} \cot \alpha},$$

where  $r_v$  is the uncorrected range (m),  $r_m$  is the corrected range (m),  $P$  is the atmospheric pressure (mb) at the laser station,  $T$  is the temperature (K) at the laser station,  $h_s$  is the laser's elevation above mean sea level (km), and  $\alpha$  is the altitude angle of the satellite. The formula holds for a ruby laser, which operates at 694 nm. It should be used only when  $\theta_0 > 5^\circ$ , where  $\theta_0$  is the apparent altitude angle (i. e., the altitude angle uncorrected for atmospheric bending).

PULSE ANALYSIS

C. R. H. Tsiang and C. G. Lehr

The effects of variations in pulse amplitude and shape must be carefully considered in attempting to reduce the noise and bias in laser range measurements. The simplest mode of operation in making time-interval measurements employs only a fixed-voltage threshold discriminator. Range times obtained this way are susceptible to errors caused by phenomena such as leading-edge walk and leading-edge pulse distortion. Attempts were made during operations to keep these effects to a minimum, and further work was done to record some of the laser passes on film. Errors on the order of  $\pm 5$  nsec can be expected when point-to-point amplitude changes affect the fixed threshold counter triggering circuit. Reduction of the photographic images of the outgoing- and return-pulse oscilloscope traces may produce range-correction figures to decrease the net range-counter errors to about  $\pm 1.7$  nsec. No work has yet been done to evaluate fully the accuracy of the photographic data collected during ISAGEX, and further study is necessary to substantiate empirically the estimated accuracy of  $\pm 1.7$  nsec. The process of photoreduction of the data and discussion of the system-accuracy potential is given in Lehr, Pearlman, and Scott (1970a). Early attempts at testing the effectiveness of this technique are presented in Lehr, Pearlman, and Scott (1970b).

Table 1 lists the satellite passes covered by photography. Most returns were obtained on Polaroid film, rather than on 35-mm film, because of the former's quick developing process. This rapid feedback was advantageous to the laser operator for adjusting the amplification level in the return-signal circuitry. Saturation of the amplifier, the limited resolution range of the oscilloscope at a fixed gain, and the minimum voltage imposed by the counter threshold were the constraints that had to be satisfied. Returns were photographed successfully at all three new laser stations in spite of early problems, such as signal adjustment and operation of the laser with new procedures and with an already heavily burdened crew of observers.

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The listing of satellite passes gives some information of the quality and quantity of the data, even though little has been done so far to reduce the images for range correction factors. The column marked "Total reported observation points" refers to the number of range measurements reported by the station for each pass, and the next column gives the number of attempted photographic images. The actual number of traces of return or noise pulses is given in the column labeled "Images." Of these frames, only a few are of sufficient quality that they could be measured reliably. Those of reduction quality should produce measurements of at least  $\pm 3$ -nsec consistency, and probably no worse than 1.5 nsec rms. Those images that do not qualify to be counted in the "Reduction quality" column are usually very irregularly shaped, low-level returns or extremely strong returns that go off the oscilloscope screen or are distorted by the saturation of the amplifier. In some cases, there are images of noise pulses that have falsely stopped the range counter. Even in these rejected images, there is useful information about the behavior of the return-signal circuitry under extreme amplitude conditions. Aside from the measurement of the images for range correction figures, the most important additional data come from the cataloging of return-pulse amplitudes and shapes photographed under routine laser tracking procedures. These data can be applied to studies on the scintillation of returns and to the calculation or prediction of return amplitudes. The amount of photographic data amassed during the last part of ISAGEX is insufficient in itself to have much value. Some reduction work is planned, however, so that the results can be used to evaluate the effectiveness of the photographic technique and to improve the operational procedures of the data-recording system.

#### REFERENCES

LEHR, C. G., PEARLMAN, M. R., and SCOTT, J. L.

- 1970a. A photographic technique for improved laser-ranging accuracy. In Laser and Radar Investigations, ed. by Computer Sciences Corp., NASA, Washington, vol. III, pp. 51-56.
- 1970b. Range corrections from oscilloscopic displays of laser returns. Smithsonian Astrophys. Obs. Laser Rep. No. 4, 27 pp.

Table 1. Satellite returns.

Date	Time (UT)	Satellite	Total reported observation points	Number of frames		
				Total	Images	Reduction quality
<u>Brazil</u>						
June 17	22 <sup>h</sup> 53 <sup>m</sup>	7010901	7	4	4	0
June 18	08 23	6800201	9	2	2	0
June 19	06 06	7010901	3	2	2	0
August 5	22 34	6508901	39	45	17	0
August 11	20 53	6508901	31	24	12	0
August 11	21 22	6508901	11	15	7	0
August 13	23 14	7010901	1	3	1	1
August 14	08 03	6508901	20	15	10	4
August 14	22 20	6508901	13	3	3	3
August 22	02 14	6503201	5	3	1	0
August 22	06 32	6508901	19	9	6	0
August 23	21 33	6800201	10	15	3	0
August 24	03 28	6503201	1	2	1	0
August 31	03 53	6503201	7	6	4	1
September 1	05 56	7010901	9	4	3	1
September 23	22 27	6508901	25	22	11	3
September 27	23 31	6800201	18	18	15	8
October 1	22 58	6800201	20	21	15	3
October 1	22 59	6508901	20	9	2	1
October 3	22 24	6503201	15	9	7	0
October 4	21 11	6508901	15	3	3	0
October 16	06 58	6508901	14	3	3	3
October 24	05 22	6508901	15	7	6	2

Table 1 (Cont.)

Date	Time (UT)	Satellite	Total reported observation points	Number of frames		
				Total	Images	Reduction quality
<u>Peru</u>						
May 19	17 <sup>h</sup> 59 <sup>m</sup>	6508901	41	26	25	9
May 19	22 09	6800201	33	26	26	8
May 20	17 55	6508901	16	28	28	6
May 22	02 53	6508901	8	2	2	0
May 28	03 23	6508901	17	1	1	1
June 3	01 36	6508901	31	20	20	0
June 8	13 01	6508901	60	35	27	5
June 8	23 01	6800201	17	16	14	0
June 8	23 58	6508901	34	28	21	1
June 9	00 13	6701101	16	5	5	0
June 9	13 04	6508901	49	14	14	7
June 10	00 01	6508901	43	29	20	7
June 10	11 20	6800201	21	14	12	0
June 10	13 10	6508901	45	13	9	2
June 11	00 06	6508901	31	15	13	0
June 11	13 16	6508901	55	48	37	22
June 11	22 08	6800201	4	3	2	0
June 12	13 20	6508901	51	39	33	20
June 13	13 24	6508901	29	17	12	0
June 13	22 16	6508901	22	15	15	0
June 13	22 47	6800201	32	20	19	0
June 15	22 22	6508901	26	12	9	0
July 6	17 35	6508901	20	18	16	7
August 17	21 21	6508901	14	6	9	0

Table 1 (Cont.)

Date	Time (UT)	Satellite	Total reported observation points	Number of frames		
				Total	Images	Reduction quality
<u>South Africa</u>						
August 19	00 <sup>h</sup> 08 <sup>m</sup>	6508901	45	60	33	5
August 20	00 11	6508901	28	44	27	4
August 23	00 24	6508901	42	54	22	10
August 24	00 26	6508901	31	56	12	1
August 24	07 16	6800201	5	31	8	1
August 24	17 52	6800201	12	25	3	0
August 25	06 07	6800201	9	36	16	0
August 26	00 37	6508901	28	49	20	7
August 26	22 36	6508901	12	41	8	0
August 26	22 39	6508901	14	4	3	0
August 27	17 19	6508901	15	24	14	2
August 27	22 40	6508901	29	49	16	0
August 28	00 47	6508901	10	32	20	0
August 29	22 48	6508901	40	57	15	0
August 30	18 15	6800201	14	23	14	0
August 30	22 53	6508901	37	55	36	0
August 31	06 10	6800201	3	32	17	0
August 31	22 56	6508901	41	56	35	0
September 1	17 06	6800201	10	16	7	0
September 4	21 90	6508901	35	47	24	0

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**STATISTICS: RETURNS AND FAILURES**

B. R. Miller

The following statistics represent the SAO and Air Force Baker-Nunn optical and the SAO laser returns during the months of intensive tracking on the individual satellites.

The first 2 months' predictions were generated with a paucity of observations, which may account for the sparse number of returns. During the rest of ISAGEX, as more observations became available for use in predicting orbits, the number of returns increased also.

The numbers here reflect the data collected during ISAGEX after gross errors were removed in the orbit computations for predictions.

The data were processed for validation purposes before being used in analysis. Therefore, there is a discrepancy in these figures and the final data used in the analysis.

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## BE-B (6406401)

Station					
Location	COSPAR number	January	March	April	Total
<u>Optical Returns</u>					
San Fernando, Spain	9004		10	5	15
Naini Tal, India	9006		6	6	12
Maui, Hawaii	9012		5	7	12
Mt. Hopkins, Arizona	9021	6	3	13	22
Olifantsfontein, South Africa	9022	5			5
Island Lagoon, Australia	9023	13	8	5	26
Dodaira, Japan	9025	8	1	9	18
Arequipa, Peru	9027			1	1
Debre Zeit, Ethiopia	9028			2	2
Dionysos, Greece	9030		7	13	20
Natal, Brazil	9039	1		5	6
Rosamund, California	9113	5	1		6
Cold Lake, Canada	9114	2	16		18
Johnston Island	9117	1			1
Mt. John, New Zealand	9119	9	21	1	31
San Vito, Italy	9120	—	<u>2</u>	<u>13</u>	<u>15</u>
Total		50	80	80	210
<u>Laser Returns*</u>					
Arequipa, Peru	7907			6 (18)	6 (18)
Mt. Hopkins, Arizona	7921		1 (3)	6 (21)	7 (24)
Natal, Brazil	7929	1 (1)		3 (8)	4 (9)
Dionysos, Greece	7930	—	<u>3 (5)</u>	—	<u>3 (5)</u>
Total		1 (1)	4 (8)	15 (47)	20 (56)

\* The first number is the number of passes of the satellite; the number in parentheses is the total number of points in those passes.

## BE-C (6503201)

Station					
Location	COSPAR number	February	March	August	Total
<u>Optical Returns</u>					
San Fernando, Spain	9004	6	1	30	37
Naini Tal, India	9006	14	3		17
Maui, Hawaii	9012	12	2	42	56
Mt. Hopkins, Arizona	9021	8	5		13
Olifantsfontein, South Africa	9022	20	3	21	44
Island Lagoon, Australia	9023	47	6	11	64
Dodaira, Japan	9025		8	7	15
Arequipa, Peru	9027	2		8	10
Debre Zeit, Ethiopia	9028	3		1	4
Dionysos, Greece	9030	2	2	44	48
Natal, Brazil	9039	5		5	10
Rosamund, California	9113	2	3		5
Cold Lake, Canada	9114	5	3	20	28
Johnston Island	9117	10		3	13
Mt. John, New Zealand	9119	34	4	26	64
San Vito, Italy	9120	—	—	68	68
Total		170	40	286	496
<u>Laser Returns</u>					
Olifantsfontein, South Africa	7902	4 (6)	1 (2)		5 (8)
Arequipa, Peru	7907	2 (9)	11 (30)	25 (181)	38 (220)
Mt. Hopkins, Arizona	7921	2 (4)	2 (4)	2 (3)	6 (11)
Natal, Brazil	7929	7 (20)	2 (6)	13 (46)	22 (72)
Dionysos, Greece	7930	—	1 (3)	10 (37)	11 (40)
Total		15 (39)	17 (45)	50 (267)	82 (351)

Station		COSPAR number												Total
Location		January	February	March	April	May	June	July	August				Total	
<u>Optical Returns</u>														
San Fernando, Spain	9004	6	45	48	5	12	33	49	3			201	201	
Naini Tal, India	9006	13	23	16	13	4		5				74	74	
Maui, Hawaii	9012			9	12	17	21	27	22			108	108	
Mt. Hopkins, Arizona	9021	19	26	28	11	6	32	15	1			138	138	
Olifantsfontein, South Africa	9022	1	6	17	40	3	9	3	8			87	87	
Island Lagoon, Australia	9023	8	10	15	22	6	11	14	45			131	131	
Dodaira, Japan	9025	27	21	18	3	1	39	1				110	110	
Arequipa, Peru	9027	4		25	36	1	9	13	75			163	163	
Debre Zeit, Ethiopia	9028	8	15	11	12		15	9	2			72	72	
Dionysos, Greece	9030	10	19	24	7	24	14	36	3			137	137	
Natal, Brazil	9039	4	7	8	13		18	3	24			77	77	
Rosamund, California	9113	12	17	15		9						53	53	
Cold Lake, Canada	9114	3	27	14			17	5				66	66	
Johnston Island	9117	3	5				41	10	4			63	63	
Mt. John, New Zealand	9119	39	23		30	108	28	167	162			557	557	
San Vito, Italy	9120		33	12	1	38		63				147	147	
Total		158	276	260	205	229	287	420	349			2184	2184	
<u>Laser Returns</u>														
Olifantsfontein, South Africa	7902		7 (33)	13 (220)	26 (349)	32 (390)	36 (658)	49 (1362)	46 (1152)			209 (4164)	209 (4164)	
Arequipa, Peru	7907	5 (10)	2 (12)	13 (113)	23 (187)	41 (461)	58 (1442)	12 (404)	38 (742)			192 (3371)	192 (3371)	
Mt. Hopkins, Arizona	7921	2 (3)	4 (9)	7 (63)	13 (74)	16 (53)	2 (11)					44 (213)	44 (213)	
Natal, Brazil	7929	5 (13)	16 (68)	1 (7)	19 (79)	22 (110)	22 (108)	12 (74)	19 (231)			116 (690)	116 (690)	
Dionysos, Greece	7930		4 (8)	3 (8)	4 (8)	6 (26)	14 (66)	4 (17)				35 (133)	35 (133)	
Total		12 (26)	33 (130)	37 (411)	85 (697)	117 (1040)	132 (2285)	77 (1857)	103 (2125)			596 (8571)	596 (8571)	

## D1C (6701101)

Station					
Location	COSPAR number	March	April	June	Total
<u>Optical Returns</u>					
San Fernando, Spain	9004	6	17	51	74
Naini Tal, India	9006	6	14		20
Maui, Hawaii	9012	15	18	47	80
Mt. Hopkins, Arizona	9021	8	27	35	70
Olifantsfontein, South Africa	9022	11	7	5	23
Island Lagoon, Australia	9023	20	10	5	35
Dodaira, Japan	9025		16	1	17
Arequipa, Peru	9027		3	6	9
Debre Zeit, Ethiopia	9028	3	4	2	9
Dionysos, Greece	9030	2	19	33	54
Natal, Brazil	9039	1	2	4	7
Rosamund, California	9113		10	14	24
Cold Lake, Canada	9114				
Johnston Island	9117			12	12
Mt. John, New Zealand	9119	1	7		8
San Vito, Italy	9120	—	<u>32</u>	<u>11</u>	<u>43</u>
Total		73	189	226	488
<u>Laser Returns</u>					
Arequipa, Peru	7907			11 (134)	11 (134)
Mt. Hopkins, Arizona	7921	1 (6)		6 (19)	7 (25)
Natal, Brazil	7929			2 (11)	2 (11)
Dionysos, Greece	7930	—		<u>36 (187)</u>	<u>36 (187)</u>
Total		1 (6)		55 (351)	56 (357)

DID (6701401)

Station		COSPAR number							Total
Location		January	February	March	April	May	July	Total	
<u>Optical Returns</u>									
San Fernando, Spain	9004	7	5	9	3	28	76	128	
Naini Tal, India	9006	15	8	19	3	7		52	
Maui, Hawaii	9012		8	25	4	50	55	142	
Mt. Hopkins, Arizona	9021	12	8	13	5	12	18	68	
Olifantsfontein, South Africa	9022	8	11		14	2	5	40	
Island Lagoon, Australia	9023	24	18	9	14	7	5	77	
Dodaira, Japan	9025	4	1	6	4	2	5	22	
Arequipa, Peru	9027	1			4	4	5	14	
Debre Zeit, Ethiopia	9028	2	5		2	3	6	18	
Dionysos, Greece	9030	1	2	7	5	21	46	82	
Natal, Brazil	9039		12	1			4	17	
Rosamund, California	9113	10	1	4	2	13		30	
Cold Lake, Canada	9114	2					5	7	
Johnston Island	9117						5	5	
Mt. John, New Zealand	9119	27	22		29	12	10	100	
San Vito, Italy	9120		<u>10</u>	<u>6</u>	<u>15</u>	<u>6</u>	<u>94</u>	<u>131</u>	
Total		113	111	99	104	167	339	933	
<u>Laser Returns</u>									
Arequipa, Peru	7907	4 (5)			2 (15)	23 (272)		29 (292)	
Mt. Hopkins, Arizona	7921	1 (2)	2 (3)	4 (17)		56 (439)		63 (461)	
Natal, Brazil	7929				2 (8)	5 (26)	3 (19)	10 (53)	
Dionysos, Greece	7930				<u>1 (2)</u>	<u>21 (57)</u>	<u>14 (75)</u>	<u>36 (134)</u>	
Total		5 (7)	2 (3)	4 (17)	5 (25)	105 (794)	17 (94)	138 (940)	

Station		COSPAR number										Total
Location		March	April	May	June	July	August					Total
<u>Optical Returns</u>												
San Fernando, Spain	9004	3	10	2		11	32					58
Naini Tal, India	9006	1	18	1			2					22
Maui, Hawaii	9012		19	18	12	19	25					93
Mt. Hopkins, Arizona	9021	5	30	5	1	3	1					45
Olifantsfontein, South Africa	9022	25	15	5	1	3	8					57
Island Lagoon, Australia	9023	23	23	9	7	12	17					91
Dodaira, Japan	9025	3	16	1			3					23
Arequipa, Peru	9027	24	21	14	3	18	60					140
Debre Zeit, Ethiopia	9028	3	16	5	1	10	23					58
Dionysos, Greece	9030	4	22	6		4	37					73
Natal, Brazil	9039	4	8	1	3	6	35					57
Rosamund, California	9113	3	10	4	4							21
Cold Lake, Canada	9114	10	10			3	26					49
Johnston Island	9117		2		1	8	8					19
Mt. John, New Zealand	9119	19	40	43	40	58	74					274
San Vito, Italy	9120	<u>1</u>	<u>15</u>	<u>—</u>	<u>—</u>	<u>1</u>	<u>22</u>					<u>39</u>
Total		128	275	114	73	156	373					1119
<u>Laser Returns</u>												
Olifantsfontein, South Africa	7902	7 (51)	18 (108)	20 (181)	29 (328)	20 (257)	23 (239)					117 (1164)
Arequipa, Peru	7907	2 (8)	16 (106)	23 (243)	31 (594)	13 (233)	15 (173)					100 (1357)
Mt. Hopkins, Arizona	7921	3 (17)	10 (44)	15 (95)	1 (6)		1 (1)					30 (163)
Natal, Brazil	7929		9 (49)	10 (66)	12 (51)	9 (77)	14 (102)					54 (345)
Dionysos, Greece	7930	<u>2 (3)</u>	<u>2 (2)</u>	<u>2 (2)</u>	<u>—</u>	<u>3 (7)</u>	<u>9 (40)</u>					<u>18 (54)</u>
Total		14 (79)	55 (309)	70 (587)	73 (979)	45 (574)	62 (555)					319 (3083)

People (7010901)

Station		People (7010901)											
Location	COSPAR number	January	February	March	April	May	June	July	August	Total			
<u>Optical Returns</u>													
Maui, Hawaii	9012		8	8	3	7	2	9	9	46			
Olifantsfontein, South Africa	9022			1	2		2			5			
Arequipa, Peru	9027			4	10	10	11	8	18	61			
Debre Zeit, Ethiopia	9028	10	8	6	10	14	2	6	7	63			
Natal, Brazil	9039	11		4	7	12	9	6	16	65			
Johnston Island	9117	<u>1</u>	<u>14</u>	<u>2</u>	<u>1</u>	<u>1</u>	<u>2</u>	<u>7</u>	<u>—</u>	<u>28</u>			
Total		22	30	25	33	44	28	36	50	268			
<u>Laser Returns</u>													
Arequipa, Peru	7907			6 (30)	5 (14)	2 (22)	27 (208)		2 (14)	42 (288)			
Natal, Brazil	7929	<u>1 (1)</u>				<u>8 (21)</u>	<u>37 (192)</u>	<u>2 (20)</u>	<u>15 (58)</u>	<u>63 (292)</u>			
Total		1 (1)		6 (30)	5 (14)	10 (43)	64 (400)	2 (20)	17 (72)	105 (580)			

## EVALUATION OF LASER OPERATIONS

J. Thorp

We selected a 3-month period during ISAGEX for collecting laser data in order to evaluate the potential of the SAO laser system. We chose June, July, and August for its good weather and the South Africa station since it was ranging to only the Geos 1 and Geos 2 satellites. The evaluation, therefore, did not have to consider the questionable orbits or the bad aspect angles of the magnetically stabilized satellites.

Statistics were generated by using laser passage information compiled at each station. We considered both individual points (Figure 1) and total arcs (Figure 2). Two to 65 points were predicted per arc during this period, with an average of 35 points per arc. Arcs started and ended at 20° above the horizon.

For simplicity in preparing the graphs, each attempted point was considered an attempted arc and each successful point was considered a successful arc. This, however, did create some apparent discrepancies between the two graphs. The category "not attempted (N.A.) other," which includes pass conflicts, observer errors, and attempts not made owing to safety, shows a decided decrease when points are compared to arcs. The main reason for this is that points at the beginning and end of many passes were not attempted because of the hazards of operating below 30°.

Another discrepancy is apparent when successful points are compared to successful arcs. However, as the following statistics show, 44 to 64% of the arcs have less than 16 successful points:

	Percentage of arcs				Total predictions (points/arcs)	Total successes (points/arcs)
	Number of points					
	1-3	4-6	7-15	16 or more		
June	16%	20%	28%	36%	4834/141	1013/67
July	8%	10%	16%	66%	4001/114	1588/69
August	5%	11%	34%	50%	5572/152	1391/69

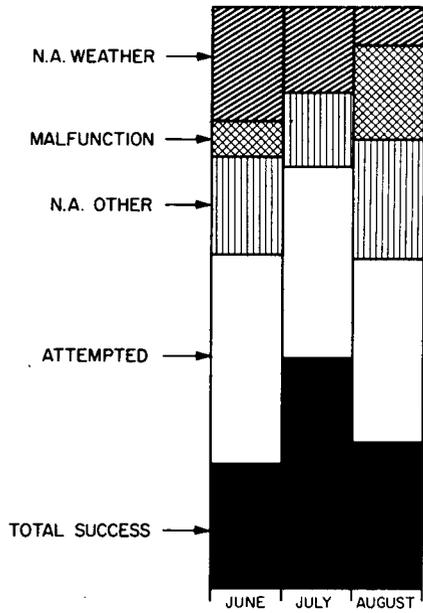


Figure 1. Percentage of total predicted points from South Africa.

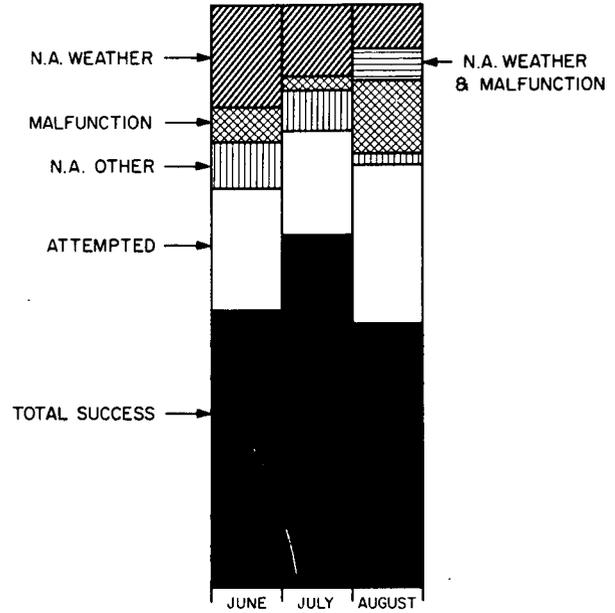


Figure 2. Percentage of total predicted arcs from South Africa.

The percentage of successful points per pass seems to be related to weather (July had the best weather). During August, a period was reported as not attempted owing to a malfunction, but, in fact, the maintenance was scheduled for and performed during cloudy weather.

From the statistics, we can expect that, in general, with the laser system operating properly, about 30% of the predicted points will be successful. When dealing with arcs, we see a higher percentage of overall success, in that about half the predicted arcs are successful. If we discount weather completely, the success ratio is higher, around 70%. The cases studied are ideal, and we feel that this is the best the system can do under present circumstances. When bad weather occurs and when satellites with less stable orbits are added, the success ratio drops substantially.

PREDICTION PROBLEMS

J. Latimer

Generating accurate predictions is crucial to the successful use of static-pointing laser systems. The prediction-observation-orbit-determination cycle is a self-sustaining process when it works properly. In general, the process functioned reasonably well during ISAGEX, except for the Peole satellite, which, with a perigee of 500 km, is subject to a great deal of atmospheric drag.

Section 1 discusses the prediction accuracies obtained, and Section 2 presents a technique for improving poor predictions. Section 3 deals with the drag problem, especially as it relates to Peole.

1. PREDICTION ACCURACIES

In Table 1 we have estimated the accuracy of predictions for the ISAGEX satellites during the saturation observing periods. The best way to express accuracies seems to be in topocentric arcminutes, since this bears most directly on the static-pointing laser system. In general, because of the beamwidth, prediction accuracies to 10 arcmin are desirable; a smaller accuracy presents problems for acquisition (although the problem is not completely insurmountable, as discussed in Section 2).

Predictions degrade exponentially in time. We give accuracies for the first and last day of each week's predictions. Interpolation will provide estimates for the other days. The figure for the first day is actually the rms of the orbit determination. Since we never found any discontinuity between the rms orbit fit and the residuals of the first day's observations using the extrapolated orbit, this seems an appropriate measure of the starting value.

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Table 1. Prediction accuracies (arcmin) from the first to the last day of each week.

ISAGEX period	Week	Peole	Geos 1	Geos 2	DID	BE-B	BE-C	DIC
I	Jan. 6-13	No predictions	1-27			1.3-42		
	Jan. 13-20	Non-SAO orbit (25-min time difference)	1.8-54			2.3-20		
	Jan. 20-27	4-X	1.3-3.5			11.5-X		
II	Feb. 10-18	4.5-1155	1-2				1-3.5	
	Feb. 18-24	4-215	1.8-2.5				1-3.8	
	Feb. 24-Mar. 3	4-90	0.8-2				1-7	
	Mar. 3-10	4.8-350	0.8-4.8				1-X	
III	Mar. 24-31	4-63	0.8-13	1-10		10.8-37		
	Mar. 31-Apr. 7	5-167	0.8-2.5	0.8-6.8		0.8-2.8		
	Apr. 7-14	7.3-253	0.5-2.8	0.5-7.5		0.8-X		
	Apr. 28-May 5	0.8-850, 115	0.7-1.3	0.8-1.5	0.8-10			
IV	May 5-12	5-232, 29	0.7-9	0.8-1.5	0.5-9			
	May 12-19	4-54, 87	0.7-6	0.5-0.8	0.4-21			
	June 5-9	1.5-220, 54	0.4-12	0.4-12				1-31
	June 9-16	2.4-53, 38	0.4-5	0.5-2				1-22
V	June 16-23	3.5-38, 22	0.4-7	0.3-11				0.7-78
	June 23-30	0.3-69, 55	0.3-6	0.3-2				0.5-X
	July 14-21	2-490, 81	0.5-1.3	0.5-2.5	0.8-2			
	July 21-28	4.8-X, 59	0.5-4	0.5-13	0.7-14			
VI	Aug. 11-18	4.8-X, 116	0.5-5	0.5-4.5				1-2.5
	Aug. 18-25	1.3-53, 87	0.5-3.4	0.5-1.4				1-9
	Aug. 25-Sept. 1	1.5-132, 86	0.5-2	0.5-2				0.8-1.3

The final day's accuracy is determined by a direct comparison of pointing angles from the expiring and the fresh predictions. This overlap day was generated to ensure operation even in the case of communications delays. Occasionally, there was no overlap day, and a comparison could not be made. This is indicated by an "X" in Table 1. Clearly, in any week, the final day's figure is uncertain by the amount of error in the next first day's prediction, which is almost always relatively small. The error is predominantly in the along-track direction; the across-track component is relatively insignificant.

Notice in the table the large errors in the Peole predictions. In an attempt to improve the accuracy, we generated Peole predictions twice a week beginning with period IV. From period IV on, we give two overlap-day comparisons in addition to the first orbit fit. Although there was noticeable improvement, Peole predictions remained troublesome (owing to atmospheric drag, see Section 3).

Finally, since we intended for these values to represent worst case situations, we chose to compare pointing angles at the culmination of each pass. Culmination, or the point of closest approach, is more sensitive to orbital errors than are lower elevation points.

## 2. FIELD UPDATING OF PREDICTIONS

One of the useful properties of laser ranging systems is their ability to operate at low elevation angles. This permits errors in the predicted satellite mean anomaly to be quickly detected and corrected. Generally, the error in satellite mean anomaly is the only significant one in satellite ephemerides, so that when this error has been determined by field personnel, they can update predictions from the Computations Center in Cambridge by applying a simple correction to the firing time of their look angles.

Figure 1 demonstrates how the error in mean anomaly is determined. The satellite is predicted to be at position 1, azimuth, altitude, and range from the station, at epoch 1, and at position 2 at epoch 2. Suppose that at epoch 1 the observed and the predicted range differ by  $\Delta r$ . (It is feasible to obtain returns in this case, provided the major component of the satellite's motion is toward the observer; this is true very

early in the pass for passes with high culmination.) We can assume that the predicted time interval between epochs 1 and 2 is correct, although the epochs themselves are in error by  $\Delta t$ ; and similarly  $s$ , the range interval between positions 1 and 2, is correct, although the range to position 1 is wrong. Then,  $\Delta t = (d/s) \Delta r$ , approximately.

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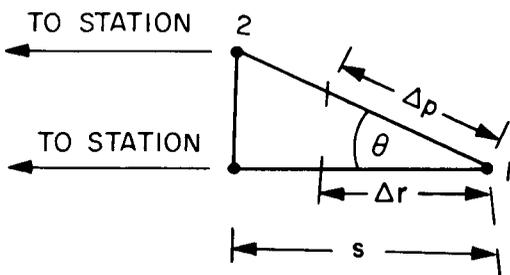


Figure 1. Satellite-station geometry at a low elevation angle.

Actually, for convenience,  $\Delta r$  and  $s$  are expressed in terms of propagation times rather than of distances. The time interval between successive predicted points,  $d$ , is almost always 15 sec;  $s$  varies, but it is typically around 0.5 msec for low elevation angles. In order to determine  $\Delta t$  quickly, station personnel use graphs of the linear relationships between  $\Delta t$  and  $\Delta r$  for the various values of  $s$  frequently encountered. The value  $\Delta r$  can be resolved to 1 nsec, so that, for example, when  $s = 0.5$  msec and  $d = 15$  sec,

$$\Delta t = \frac{30 \text{ msec mean-anomaly correction}}{1 \text{ nsec observed range error}} .$$

Figure 2 is the station graph for determining the error in mean anomaly. Given  $\Delta r$  and  $s$ ,  $\Delta t$  can be quickly found.

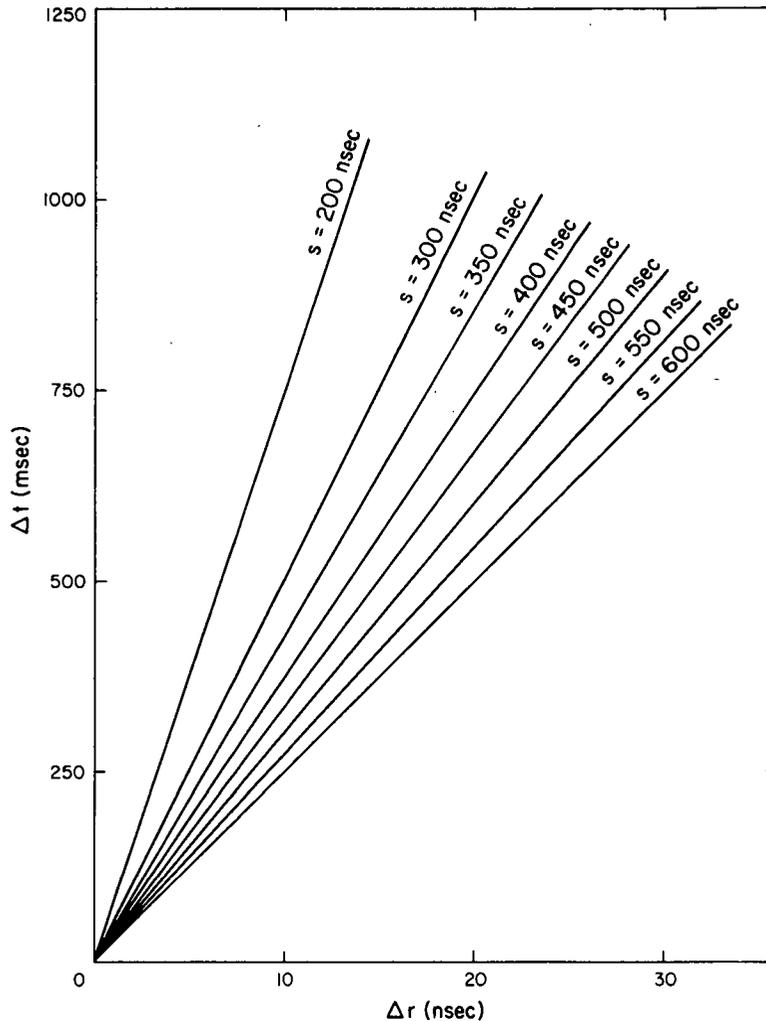


Figure 2. Typical station graph for determination of mean-anomaly error:  $s$  = propagation-time differences between successive predicted points;  $\Delta r$  = (observed - predicted) range propagation time;  $\Delta t$  = mean-anomaly correction;  $d = 15$  sec.

### 3. ATMOSPHERIC-DRAG EFFECTS

The Peole satellite (7010901) has an apogee of 730 km and a perigee of 500 km. Predictions of it are difficult because not only is it subject to considerable atmospheric drag but also the drag is highly variable. Figure 3 shows the effects of drag varying by about an order of magnitude.



For predictions to be useful for laser ranging, we estimate that the acceleration in mean anomaly (the empirically determined term that represents the effect of drag) ought to be correct within 5 or 10 parts per million. As can be readily seen from Figure 3, the term frequently changes by several times this amount in just a few days. The problem amounts to predicting the future state of the atmosphere, for which there is no satisfactory procedure.

An additional problem is that of obtaining an adequate orbital determination in the first place. Orbit determination was frequently weak because of insufficient observations.

We attempted to correlate the orbital acceleration of Peole with two parameters that are readily obtainable. The first is  $\psi$ , the geocentric angle between the satellite perigee and the center of the atmospheric "hot spot," i. e., the subsolar "bulge" in the atmosphere. We set the hot-spot center at the subsolar point delayed by  $30^\circ$  in longitude. The angle  $\psi$  is plotted in Figure 3, along with preliminary values of our second parameter, the 10.7-cm solar flux, which gives a rough indication of the solar influence on atmospheric activity.

It seems apparent that the shapes of both the  $\psi$  and the flux curves are reflected in the acceleration of Peole, although not in any quantitatively consistent manner.

We conclude that the only feasible way to solve the prediction problem caused by Peole's high drag is to increase the frequency of the prediction-observation-orbit-determination cycle. Yet, a cycle faster than the present twice-weekly one is impractical. Possibly, an orbit-computation capability on location at the laser sites would permit a rapid and accurate enough iteration of the cycle for the drag problem to be overcome.

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## SIGNAL STRENGTH AND OBSERVABILITY

J. Latimer

One problem that the laser-ranging technique continually presents is that of observability; that is, how can the observer know if a particular satellite and laser geometrical configuration is such that he is likely to receive a sufficiently strong return signal?

Although this question applies to all satellites, an especially striking situation is that of the four magnetically stabilized satellites (BE-B, BE-C, D1C, and D1D) when they are observed from stations in the southern hemisphere. It can be very difficult to observe them successfully because the retroreflector arrays, on the north-seeking ends of these satellites, tend to face away from southern stations.

To see the effect of this problem, we selected at random 26 passes (342 observations) of BE-C observed from station 7907 (Arequipa, Peru), at latitude  $-17^\circ$ . The observations were made between August 31 and November 5, 1971, and all give acceptably small orbit-fit residuals. Figure 1 is a plot of the passes, in the station's altitude-azimuth coordinate system. Indeed, when the satellite is in the northern half of the sky, it cannot be observed from this station. (Attempts were made to observe all passes.)

If we assume that the satellite is always oriented along the lines of the earth's magnetic field, we can calculate the aspect angle at the satellite between the symmetry axis (North) and the line of sight to the station. We used the spherical-harmonic representation (up to degree and order 4<sup>\*</sup>) of Cain, Hendricks, Langel, and Hudson (1967) as the geomagnetic model and derived aspect angles (see Appendix A) for each

---

\* Although Cain's field is represented to degree and order 10, we found that the truncated field yielded aspect angles differing only by about  $1^\circ$ . This is sufficient, considering that the satellite is likely to oscillate about the field direction with an amplitude of a degree or more.

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of the 342 observations. Figure 2 is a plot of the reflective area versus the angle of incidence for the BE-C satellite. Cases 1 and 2 are measured from different radial angles. We used mean values, since there is no way to determine the radial angle, and the difference is slight. The histogram in Figure 3 shows the observed frequency of occurrence of each aspect angle. The significance of the aspect angle is its relation to the effective area of the satellite retroreflector array.

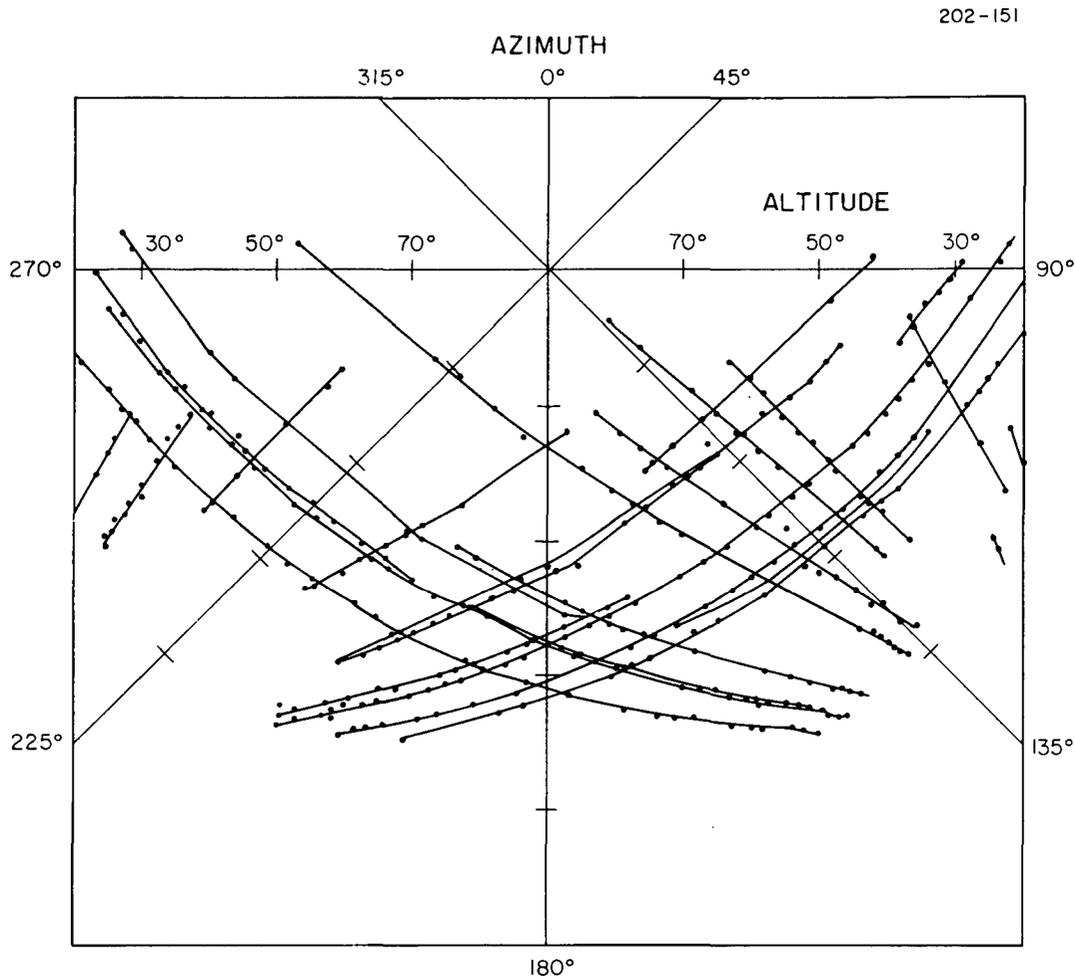


Figure 1. 26 passes of BE-C in the altitude-azimuth coordinate system of the Peru station.

We could avoid generating laser predictions for magnetically stabilized satellites that are impossible to observe by calculating the aspect angle, choosing a suitable limit for the angle, and suppressing all predicted points exceeding that limit. Although the

histogram of Figure 3 serves to confirm Figure 2 (there were no returns from aspect angles greater than 110°, the Figure 3 cutoff), the slow fall-off of returns for large aspect angles suggests that we ought to consider the range equation if we wish to determine whether particular satellite-station geometries are likely to be observable.

202-151

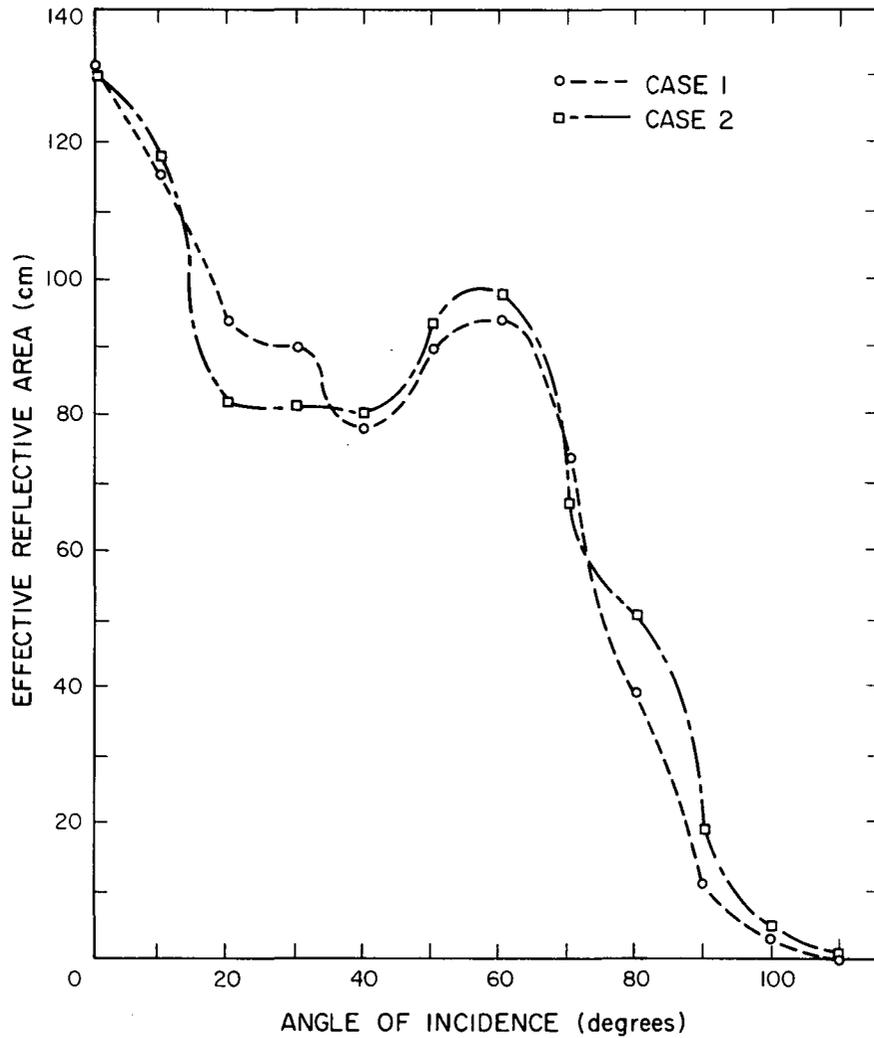


Figure 2. Reflective area vs. angle of incidence of BE-C (from Minott, 1963).

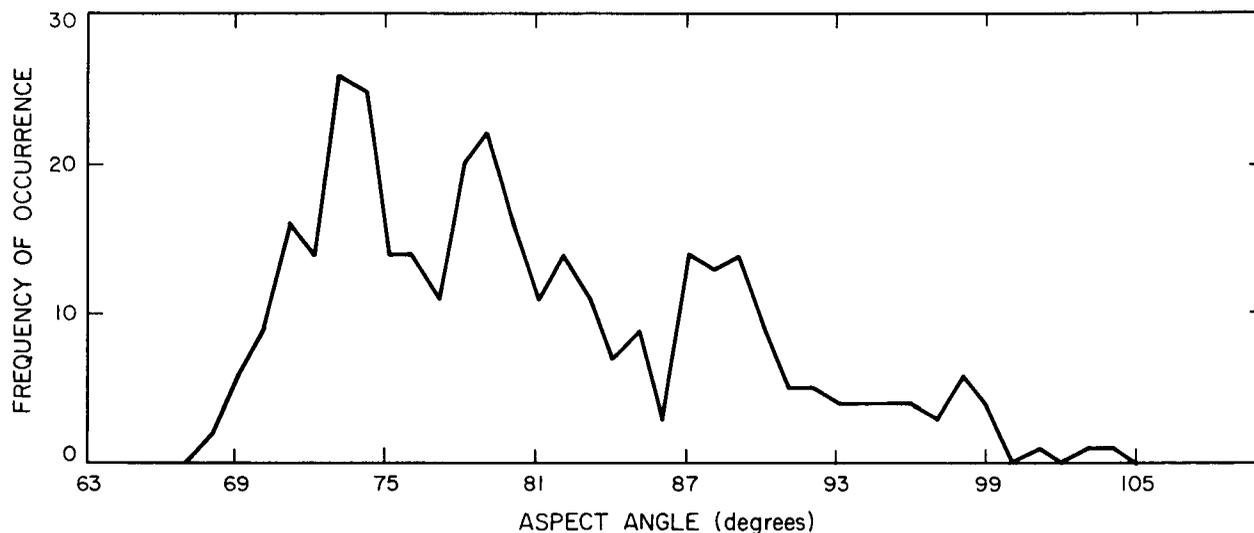


Figure 3. Histogram of the aspect angles of satellite BE-C.

We estimated average signal strengths for all the data by using the range equation of Appendix B and the retroreflector area function of Figure 2. The results are displayed in Figure 4. The large population at low signal strengths is disturbing, but we must consider the following factors:

A. Error in the assumed beam divergence. Owing to the method of recording the transmitted beam divergence, some estimates are surely too high; therefore, some signals are greater than indicated.

B. Scintillation. We estimated only the average signal – the actual signal may vary by more than an order of magnitude (Jaffe, 1971), and some of the low average signals probably yielded high actual signals.

C. Weak-signal conditions. There are many more opportunities to attempt observation under weak-signal conditions than under strong-signal conditions. Thus, although the probability of success diminishes, the number of opportunities greatly increases, yielding a significant number of weak-signal returns.

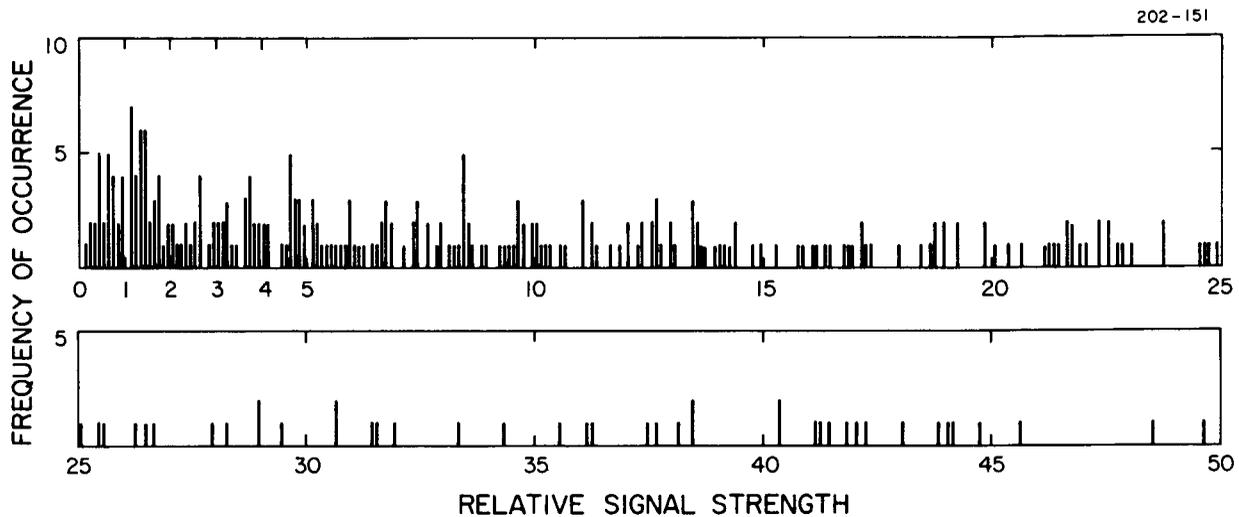


Figure 4. Histogram of relative signal strengths for Peru BE-C data.

Point C is easily verified. We constructed a uniform distribution of 706 satellite positions at a height of 1150 km (typical for BE-C) over Peru such that all points were above the elevation angle limit of  $10^\circ$ . By using a typical beam divergence (3.5 mrad) and the same values for the magnetic field that we used in Figure 3, we computed the expected average signal strengths. From the histogram in Figure 5, it is clear that the population of weak-signal situations is large.

In conclusion, we can see that satellite configurations yielding aspect angles greater than  $110^\circ$  can be suppressed without loss of possible observations. We will attempt to measure signal strengths directly in order to have a better idea of both the minimum useful signal and the statistical behavior of scintillation. In addition, we intend to extend this analysis to the gravitationally stabilized satellites Geos 1, Geos 2, and Peole. For these, the principal problem is that of obtaining returns at low station-elevation angles.

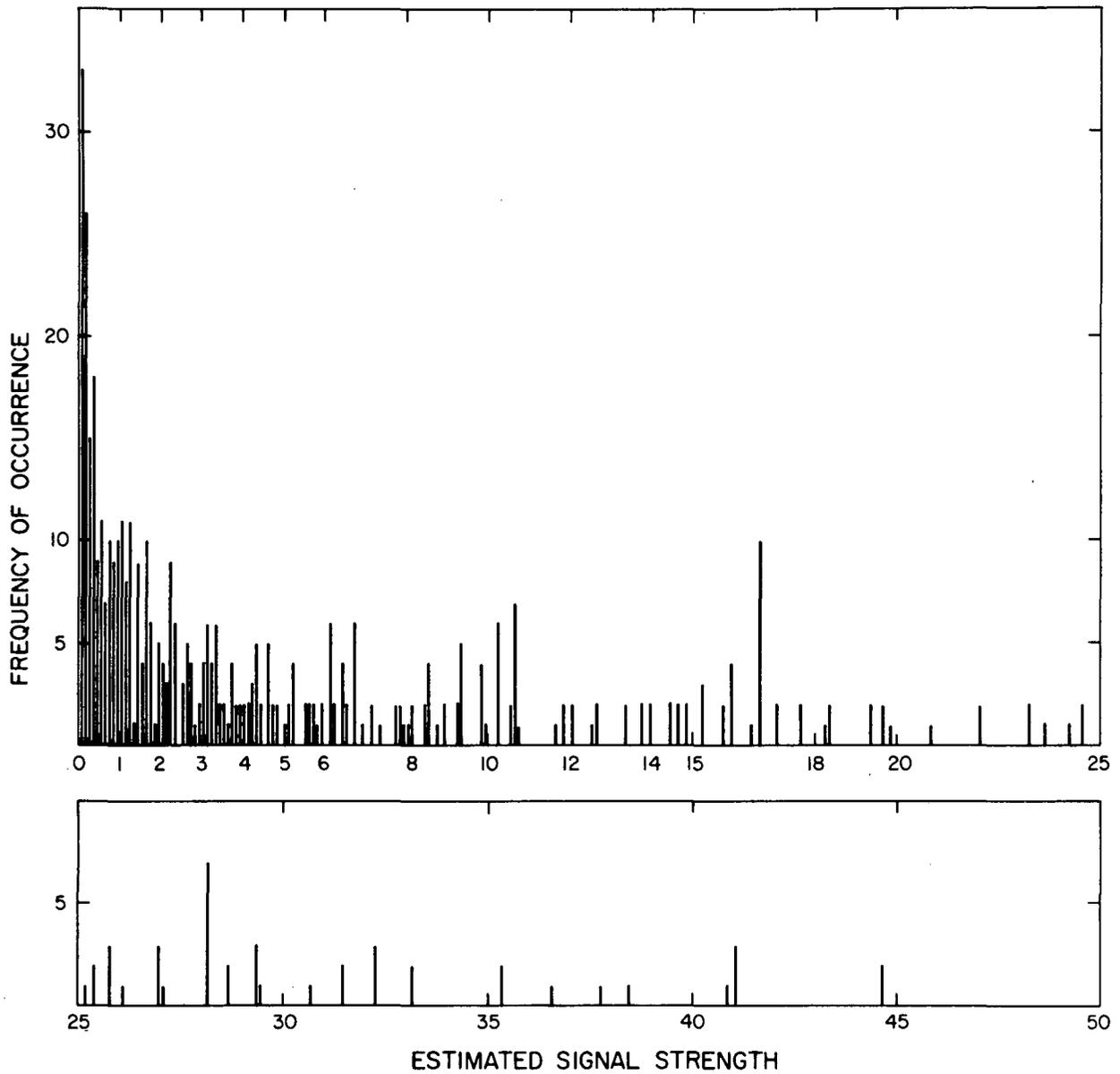


Figure 5. Histogram of estimated signal strengths for a uniform distribution of positions of BE-C from Peru.

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## APPENDIX A

### COMPUTATION OF ASPECT ANGLE $\alpha$

Station position:  $\vec{S}$

Observation vector (line-of-sight direction and range):  $\vec{O}$

Satellite position:  $\vec{P}$

Magnetic field at  $\vec{P}$ :  $\vec{F}$

Clearly,

$$\vec{S} + \vec{O} = \vec{P}$$

and

$$\cos \alpha = - \frac{\vec{O} \cdot \vec{F}}{|\vec{O}| |\vec{F}|} ,$$

where  $\vec{F} = \vec{F}(\vec{P})$ ; i. e.,

$$\vec{F} = -\nabla V ,$$

where

$$V = a \sum_{n=1}^{n_{\max}} \left(\frac{a}{r}\right)^{n+1} \sum_{m=0}^n \left( g_n^m \cos m\phi + h_n^m \sin m\phi \right) P_n^m(\theta) ;$$

$r$ ,  $\theta$ , and  $\phi$  are polar coordinates for  $\vec{P}$ ,  $a$  is the earth's mean radius,  $P_n^m$  are Schmidt normalized spherical functions, and  $g_n^m$  and  $h_n^m$  are the coefficients of the magnetic-field representation.

## APPENDIX B

### RANGE EQUATION

We use the range equation from Lehr (1966):

$$S = \frac{E}{2.86 \times 10^{-19}} \left( \frac{1}{R^4} \right) \left( \frac{A_s}{\Omega_t} \right) \left( \frac{A_r}{\Omega_s} \right) T^2 \text{ photons} ,$$

where  $E$  is the output laser energy ( $\approx 7.2$  J),  $R$  is the range to the satellite in megameters,  $T$  is the atmospheric transmission for Peru ( $\approx e^{-0.071/\sin \alpha}$ , where  $\alpha$  is the elevation angle),  $A_s$  is the effective area of the satellite reflective surface,  $\Omega_s$  is the solid angle of the reflected laser beam ( $= 2.83 \times 10^{-9}$  sr),  $A_r$  is the effective area of the receiver aperture ( $= 0.21 \text{ m}^2$ ), and  $\Omega_t$  is the solid angle of the transmitted laser beam. The numerical factor converts joules to photons at a wavelength of  $6943 \text{ \AA}$ .

In addition, we use  $N = S/58$  to represent the photodetector conversion efficiency, where  $N$  is the number of photoelectrons generated.

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## DATA FORMATS

B. R. Miller

### 1. INTRODUCTION

The observational data formats described are the ones used by SAO in various computer programs.

The optical observation format is the same as that used in the past. It has been reproduced here to facilitate use by ISAGEX members.

The laser format has been revised to provide room for time designation to the nearest nanosecond and range to 0.01 m. Temperature is now given in degrees Celsius, pressure in millibars, and humidity in percent.

### 2. SAO OPTICAL OBSERVATION CARD FORMAT AND EXPLANATION

<u>Field</u>	<u>Column</u>	<u>Description</u>
1	<u>1-7</u>	<u>Satellite identification</u>
	1-2	year of launch from 1900
	3-5	number of launch in that year
	6-7	particle number
		Satellite 1959 a1, for example, would be designated 5900101.
2	<u>8-12</u>	<u>Observation number</u> - Each observation of a satellite in a given year is designated by a different number. The source of an observation is also indicated by the observation number.
	1-9999	miscellaneous
	10000-19999	Baker-Nunn, field-reduced
	30000-39999	Moonwatch
	50000-59999	miscellaneous
	70000-79999	photoreduced Baker-Nunn

<u>Field</u>	<u>Column</u>	<u>Description</u>
3	<u>13</u>	<u>Blank</u>
4	<u>14-17</u>	<u>Station number</u> – In the COSPAR numbering format, e.g., 9039 is Natal, Brazil.
5	<u>18-23</u>	<u>Date of observation</u>
	18-19	year, from 1900
	20-21	month
	22-23	day
6	<u>24-33</u>	<u>Time designation</u> – Different types of observations are made using different time systems. Different times used in reporting SAO observations are as follows:
		a. Field-reduced Baker-Nunn observations – generally WWV received before 1966, UTC(USNO) after.
		b. Photoreduced Baker-Nunn observations – A.S
		Note: A.S is a time scale with a fixed relation to NBS(A) before April 1968 and to A.1 after then. Values of (A.S-WWV emitted) are available in tabular form.
	24-25	hour
	26-27	minute
	28-29	second
	30-33	fraction of seconds, to 0.1 msec
7	<u>34-52</u>	The interpretation of the following field depends on the code in column 56. If column 56 is 0, then the observation is right ascension and declination ( $\alpha$ , $\delta$ ).
	34	blank
	35-36	hours of $\alpha$
	37-38	minutes of $\alpha$
	39-40	seconds of $\alpha$
	41-43	fractions of seconds to 0.001 sec
	44	sign of $\delta$
	45-46	degrees of $\delta$
	47-48	minutes of $\delta$
	49-50	seconds of $\delta$

<u>Field</u>	<u>Column</u>	<u>Description</u>
	51-52	fractions of seconds to 0.01 sec If column 56 is 1, the observation is altitude and azimuth corrected for atmospheric refraction. Altitude and azimuth observations not corrected for atmospheric refraction have 3 in column 56.
	34-36	degrees of azimuth; 999 indicates azimuth is in mils
	37-38	minutes of azimuth
	39-40	seconds of azimuth
	41-43	fraction of seconds to 0.001 sec
	37-41	mils to nearest tenth if azimuth is in mils; decimal point assumed before column 41
	44	blank
	45-46	degrees of altitude; 999 indicates altitude is in mils
	47-48	minutes of altitude
	49-50	seconds of altitude
	51-52	fractions of seconds to 0.01 sec
	45-51	mils to nearest tenth if altitude is in mils; decimal assumed before column 51 If column 56 is 4, the observation is direction cosines ( $\ell$ , $m$ ), corrected for refraction; a 5 in column 56 indicates the observation is in direction cosines uncorrected for refraction.
	34	sign of $\ell$ (blank or minus)
	35-42	$\ell$ to 8 decimal places (decimal point implied before column 35)
	43	blank
	44	sign of $m$ (blank or minus)
	45-52	$m$ to 8 decimal places (decimal point implied before column 45) $n = \ell^2 + m^2$

8

53-58Index codes

53

time-precision index

<u>Code</u>	<u>Standard error in timing <math>\sigma_t</math></u>
0	No estimate
1	$\sigma_t \leq 0.0003$ sec
2	$0.0003 < \sigma_t \leq 0.002$
3	$0.002 < \sigma_t \leq 0.005$

<u>Field</u>	<u>Column</u>	<u>Description</u>
--------------	---------------	--------------------

<u>Code</u>	<u>Standard error in timing <math>\sigma_t</math></u>
-------------	---

4	0.005 < $\sigma_t$ $\leq$ 0.02
5	0.02 < $\sigma_t$ $\leq$ 0.05
6	0.05 < $\sigma_t$ $\leq$ 0.2
7	0.2 < $\sigma_t$ $\leq$ 0.5
8	0.5 < $\sigma_t$ $\leq$ 2.0
9	$\sigma_t$ > 2.0

54-55

position precision index

<u>Code</u>	<u>Standard error in direction <math>\sigma_D</math></u>
-------------	--

00	No estimate
01	$\sigma_D \leq 1''5$
02	$1''5 < \sigma_D \leq 2''5$
03	$2''5 < \sigma_D \leq 3''5$
04	$3''5 < \sigma_D \leq 4''5$
05	$4''5 < \sigma_D \leq 5''5$
06	$5''5 < \sigma_D \leq 6''5$
07	$6''5 < \sigma_D \leq 7''5$
08	$7''5 < \sigma_D \leq 8''5$
09	$8''5 < \sigma_D \leq 9''5$
10	$9''5 < \sigma_D \leq 10''5$
11	$10''5 < \sigma_D \leq 11''5$
12	$11''5 < \sigma_D \leq 12''5$
13	$12''5 < \sigma_D \leq 13''5$
14	$13''5 < \sigma_D \leq 14''5$
15	$14''5 < \sigma_D \leq 15''5$
16	$15''5 < \sigma_D \leq 16''5$
17	$16''5 < \sigma_D \leq 17''5$
18	$17''5 < \sigma_D \leq 18''5$
19	$18''5 < \sigma_D \leq 19''5$
20	$19''5 < \sigma_D \leq 20''5$
21	$20''5 < \sigma_D \leq 22''$
22	$22'' < \sigma_D \leq 23''5$
23	$23''5 < \sigma_D \leq 26''$
24	$26'' < \sigma_D \leq 29''$
25	$29'' < \sigma_D \leq 33''$

<u>Field</u>	<u>Column</u>	<u>Description</u>
	<u>Code</u>	<u>Standard error in direction <math>\sigma_D</math></u>
	26	33" < $\sigma_D$ ≤ 38"
	27	38" < $\sigma_D$ ≤ 45"
	28	45" < $\sigma_D$ ≤ 54"
	29	54" < $\sigma_D$ ≤ 1!1
	30	1!1 < $\sigma_D$ ≤ 1!3
	31	1!3 < $\sigma_D$ ≤ 1!7
	32	1!7 < $\sigma_D$ ≤ 2!1
	33	2!1 < $\sigma_D$ ≤ 2!7
	34	2!7 < $\sigma_D$ ≤ 3!5
	35	3!5 < $\sigma_D$ ≤ 4!4
	36	4!4 < $\sigma_D$ ≤ 5!8
	37	5!8 < $\sigma_D$ ≤ 7!5
	38	7!5 < $\sigma_D$ ≤ 9!7
	39	9!7 < $\sigma_D$ ≤ 13'
	40	13' < $\sigma_D$ ≤ 17'
	41	17' < $\sigma_D$ ≤ 22'
	42	22' < $\sigma_D$ ≤ 28'
	43	28' < $\sigma_D$ ≤ 37'
	44	37' < $\sigma_D$ ≤ 49'
	45	49' < $\sigma_D$ ≤ 1°1
	46	1°1 < $\sigma_D$ ≤ 1°4
	47	1°4 < $\sigma_D$ ≤ 1°8
	48	1°8 < $\sigma_D$ ≤ 2°4
	49	2°4 < $\sigma_D$

56 observation type index

<u>Code</u>	<u>Explanation</u>
0	right ascension, declination
1	altitude, azimuth (corrected for refraction)
2	not used
3	altitude, azimuth (uncorrected for refraction)
4	$\ell, m$ (direction cosines, corrected for refraction)
5	$\ell, m$ (direction cosines, uncorrected for refraction)

<u>Field</u>	<u>Column</u>	<u>Description</u>																						
57		This index refers to the date of equator and equinox to which the observation is referred. (Meaningful for right ascension and declination only.)																						
		<table border="1"> <thead> <tr> <th><u>Index</u></th> <th><u>Date</u></th> </tr> </thead> <tbody> <tr> <td>0</td> <td>Date of observation</td> </tr> <tr> <td>1</td> <td>1855.0</td> </tr> <tr> <td>2</td> <td>1875.0</td> </tr> <tr> <td>3</td> <td>1900.0</td> </tr> <tr> <td>4</td> <td>1950.0</td> </tr> </tbody> </table>	<u>Index</u>	<u>Date</u>	0	Date of observation	1	1855.0	2	1875.0	3	1900.0	4	1950.0										
<u>Index</u>	<u>Date</u>																							
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1	1855.0																							
2	1875.0																							
3	1900.0																							
4	1950.0																							
58		instrument description index																						
		<table border="1"> <thead> <tr> <th><u>Code</u></th> <th><u>Optical observations</u></th> </tr> </thead> <tbody> <tr> <td>0</td> <td>naked eye and binoculars, visual</td> </tr> <tr> <td>1</td> <td>telescope, aperture less than 5 inches</td> </tr> <tr> <td>2</td> <td>apogee telescope, astronomical refractor or reflector, theodolite, visual</td> </tr> <tr> <td>3</td> <td>Baker-Nunn camera, photographic</td> </tr> <tr> <td>4</td> <td>small missile tele-camera, tracking cameras with focal length 20 inches or greater, photographic</td> </tr> <tr> <td>5</td> <td>cinetheodolite, tracking cameras with focal length less than 20 inches, photographic</td> </tr> <tr> <td>6</td> <td>Harvard meteor camera (Super-Schmidt), photographic</td> </tr> <tr> <td>7</td> <td>stationary telescope or camera with focal length equal to or less than 10 inches, photographic</td> </tr> <tr> <td>8</td> <td>direction observation associated with a laser instrument</td> </tr> <tr> <td>9</td> <td>other instruments</td> </tr> </tbody> </table>	<u>Code</u>	<u>Optical observations</u>	0	naked eye and binoculars, visual	1	telescope, aperture less than 5 inches	2	apogee telescope, astronomical refractor or reflector, theodolite, visual	3	Baker-Nunn camera, photographic	4	small missile tele-camera, tracking cameras with focal length 20 inches or greater, photographic	5	cinetheodolite, tracking cameras with focal length less than 20 inches, photographic	6	Harvard meteor camera (Super-Schmidt), photographic	7	stationary telescope or camera with focal length equal to or less than 10 inches, photographic	8	direction observation associated with a laser instrument	9	other instruments
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7	stationary telescope or camera with focal length equal to or less than 10 inches, photographic																							
8	direction observation associated with a laser instrument																							
9	other instruments																							
9	<u>59-64</u>	<u>Blank</u>																						
10	<u>65-70</u>	<u>Conversion from the UT1 to the A. 1 time system, i. e., A. 1 - UT1</u>																						
	65	minus if A. 1 - UT1 is negative, or tens digit if positive and necessary																						

<u>Field</u>	<u>Column</u>	<u>Description</u>
	66	units digit of A. 1 - UT1 in seconds
	67-70	decimal fraction A. 1 - UT1
11	<u>71-80</u>	<u>Identification information</u>
	71-75	film number
	76	contains an S if observation is simultaneous
12	<u>77-78</u>	<u>Passive or flash information</u>
		a. If the satellite is a flashing one, column 77 will contain an F and column 78 will contain the number of the flash as it actually occurred. (This does not apply to ANNA flashes.)
		b. If the satellite is passive, columns 77 and 78 will contain the frame number.
13	<u>79</u>	Contains the letter associated with the film number if any; otherwise it will be blank.
14	<u>80</u>	Used for balloon satellites to indicate a precision reduction correction for satellite size has been added; otherwise blank.
11	<u>71-80</u>	<u>Moonwatch</u> - used for apparent magnitude information.

Precisely reduced Baker-Nunn observations are given in the coordinate system of the SAO Star Catalog (equator and equinox of 1950.0). The positions have been corrected for annual aberration, and the star positions, for proper motion to the year of observation. No corrections have been applied for diurnal aberration or parallactic refraction.

The time of the observation is given in A. S (Smithsonian Atomic Time), defined by the expression

$$A. S - UTC(USNO) = 6^S.140768 + 0.002592000 (T - 39856.0)$$

for the time period February 1, 1968, to the present; T is the Universal Time in Modified Julian Days (MJD), and 39856 is January 1, 1968:

$$MJD = \text{Julian Day} - 2400000.5$$

### 3. SAO LASER OBSERVATION FORMATS AND EXPLANATION

<u>Field</u>	<u>Column</u>	<u>Description</u>
1	<u>1-7</u>	<u>Satellite identification</u>
	1-2	year of launch from 1900
	3-5	number of launch in that year
	6-7	particle number
		Satellite 1964 64A, for example, would be designated 6406401
2	<u>8-12</u>	<u>Observation number</u>
		20000-29999 uncorrected observation
		70000-79999 corrected observation
		90000-99999 GOCC laser and direction observation
3	<u>13</u>	<u>Blank</u>
4	<u>14-17</u>	<u>Station number</u> – In the COSPAR numbering format, e. g., 7921 is SAO laser site at Mt. Hopkins, Arizona. Station designations in the 7000 series include laser sites.
5	<u>18-23</u>	<u>Date of observation</u>
	18-19	year from 1900
	20-21	month
	22-23	day
6	<u>24-35</u>	<u>Time designation</u> – Different types of observations are made using different time systems. Time systems used are indicated by the code in column 57.
	24-25	hour
	26-27	minute
	28-29	second
	30-35	fraction of seconds to 1 $\mu$ sec
7	<u>36-52</u>	<u>Interpretation</u> of the following field depends on the codes in columns 56 and 57.
	36	blank

<u>Field</u>	<u>Column</u>	<u>Description</u>										
	37-46	range in meters (decimal implied before column 45 allows range observations to be specified to 0.01 m)										
	47-48	blank										
	49-52	value of refractivity correction to 0.01 m - code 1 in column 57										
8	<u>53-58</u>	<u>Index codes</u>										
	53	time precision index										
		<table border="1"> <thead> <tr> <th><u>Code</u></th> <th><u>Standard error in timing <math>\sigma_t</math></u></th> </tr> </thead> <tbody> <tr> <td>0</td> <td><math>\sigma_t \leq 0.000005</math> sec</td> </tr> <tr> <td>1</td> <td><math>\sigma_t \leq 0.0003</math> sec</td> </tr> </tbody> </table>	<u>Code</u>	<u>Standard error in timing <math>\sigma_t</math></u>	0	$\sigma_t \leq 0.000005$ sec	1	$\sigma_t \leq 0.0003$ sec				
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0	$\sigma_t \leq 0.000005$ sec											
1	$\sigma_t \leq 0.0003$ sec											
	54-55	standard deviation of the range $\sigma_r$ in meters and tenths of meters										
	56	observation type index										
		<table border="1"> <thead> <tr> <th><u>Code</u></th> <th><u>Explanation</u></th> </tr> </thead> <tbody> <tr> <td>1</td> <td>altitude, azimuth on laser instrument</td> </tr> <tr> <td>8</td> <td>laser range</td> </tr> </tbody> </table>	<u>Code</u>	<u>Explanation</u>	1	altitude, azimuth on laser instrument	8	laser range				
<u>Code</u>	<u>Explanation</u>											
1	altitude, azimuth on laser instrument											
8	laser range											
	57	code to indicate time system and corrections applied										
		<table border="1"> <thead> <tr> <th><u>Code</u></th> <th><u>Explanation</u></th> </tr> </thead> <tbody> <tr> <td>0</td> <td>UTC emitted at transmission of laser pulse - no corrections applied to range</td> </tr> <tr> <td>1</td> <td>A. S time at reception of laser pulse - refractivity correction given in columns 49-52 but not applied to range</td> </tr> <tr> <td>2</td> <td>A. S time at reception of laser pulse - refractivity correction applied to range</td> </tr> <tr> <td>3</td> <td>UTC time at the satellite (GOCC observations - refractivity correction applied to range)</td> </tr> </tbody> </table>	<u>Code</u>	<u>Explanation</u>	0	UTC emitted at transmission of laser pulse - no corrections applied to range	1	A. S time at reception of laser pulse - refractivity correction given in columns 49-52 but not applied to range	2	A. S time at reception of laser pulse - refractivity correction applied to range	3	UTC time at the satellite (GOCC observations - refractivity correction applied to range)
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	58	Instrument description index										
		<table border="1"> <thead> <tr> <th><u>Code</u></th> <th><u>Explanation</u></th> </tr> </thead> <tbody> <tr> <td>8</td> <td>Laser observation</td> </tr> </tbody> </table>	<u>Code</u>	<u>Explanation</u>	8	Laser observation						
<u>Code</u>	<u>Explanation</u>											
8	Laser observation											
9	<u>59-64</u>	<u>Range correction</u> - pulse shape and size - codes 0 and 1 in column 57										
	59	sign										
	60	meters										

<u>Field</u>	<u>Column</u>	<u>Description</u>								
	61	10 centimeters								
	62	centimeters								
	63-64	blank								
11	<u>65-77</u>	<u>Pressure, humidity, temperature - uncorrected observations</u> only, code 0 in column 57								
	65-66	blank								
	67-70	barometric pressure in millibars								
	71-72	humidity in percent								
	73	sign of temperature								
	74-76	temperature to tenths of degrees Celsius								
	77	blank								
	<u>65-77</u>	<u>Conversion from the UT1 to the A. 1 time system</u> i. e., A. 1 - UT1 (actually A. S), code 2 in column 57								
	65	minus if A. S - UT1 is negative, or tens digit if positive and necessary								
	66	units digit of A. S - UT1 in seconds								
	67-72	decimal fraction of A. S - UT1								
	74-77	blank								
12	<u>78-80</u>	<u>Identification information</u>								
	78	blank								
	79	type of laser pass								
		<table border="1"> <thead> <tr> <th><u>Code</u></th> <th><u>Explanation</u></th> </tr> </thead> <tbody> <tr> <td>0</td> <td>night pass, satellite illuminated</td> </tr> <tr> <td>1</td> <td>night pass, satellite in shadow</td> </tr> <tr> <td>2</td> <td>daylight pass</td> </tr> </tbody> </table>	<u>Code</u>	<u>Explanation</u>	0	night pass, satellite illuminated	1	night pass, satellite in shadow	2	daylight pass
<u>Code</u>	<u>Explanation</u>									
0	night pass, satellite illuminated									
1	night pass, satellite in shadow									
2	daylight pass									
	80	blank								

## DATA VALIDATION

E. M. Gaposchkin and G. M. Mendes

Each observing station performs the following calibration exercises:

- A. Determination of the fundamental system delays to be applied to measured time intervals.
- B. Precision check of the fundamental time-interval measurement.
- C. Determination of the reliability of an individual measurement.
- D. Calculation of the reproducibility of an individual measurement.

Even with this elaborate procedure, further checking and information are necessary; two approaches are employed:

- A. Pass analysis, which depends on trends, noise, and consistency of the data.
- B. Comparison of observations with precision orbit computation.

This latter approach, the primary tool used, is the subject of this section.

Laser data have a precision of 1 m and an accuracy that is probably somewhat better. On the other hand, our current set of geodetic parameters is based primarily on 20- to 40-m camera data. From the determination of these geodetic results, Gaposchkin and Lambeck (1970, 1971) estimated the station coordinates to an accuracy of 5 to 10 m. The accuracy of our orbit computation is no better than 5 to 10 m for optimum satellites (Geos 1) and significantly worse for others (Peole). It is clear that we cannot obtain an unambiguous validation of 1-m laser data with a 10-m tool.

Orbits were computed with 4 days of data every 2 days, i. e., with 2 days of overlap. Four days was usually sufficient to compute a reliable orbit, yet short enough to minimize the effects of errors in the gravity field and station coordinates.

Owing to the paucity of laser data, we have also incorporated other tracking data, including minitrack and field-reduced Baker-Nunn. As the main computation center, SAO received laser data from CNES and GSFC to be used for predictions. We included these data in our reference orbits, although we recognized that such data did not have the benefit of refinement and validation by the originating agency. The supplementary laser data used were from stations 7050 (GSFC), 7060 (Guam), 7815 (moved to 7809, Haute Provence), and 7804 (San Fernando). The data from station 7820 (Dakar) were not included because of uncertainties in the coordinates and timing. It was not our purpose to validate the data from other agencies, but our success in validation indeed hinged on having these supplementary laser data.

Very bad data were easy to detect. The detection of poor data proved to be very difficult. Ultimately, three rules were applied:

A. The successive orbits had to be consistent. Nonuniform evolution of the mean orbital elements indicated poor data had been included in the orbit determination.

B. Orbital residuals had to be consistent (i. e., reproducible) in the two computed orbits. Using a conservative estimate of the orbital accuracy, observations that had residuals greater than 50 m (500 m for Peole) were rejected.

C. The run of residuals in a pass had to be smooth. A large variation in residuals ( $\geq 50$  m) from point to point (1 sec apart) must be an observational error as there is no unmodeled orbital perturbation of that magnitude. The run (trend or signature) of the residuals is a very powerful device and hinges on having more than 10 observations per pass. Many passes early in ISAGEX did not have sufficient data for this test to be applied, and these were consequently very difficult to validate.

The applications of these rules had varying degrees of success. The confidence we can put in the validated data varies considerably from satellite to satellite and from period to period. Some of the data in periods II and III are questionable. Such data are distributed so that only when they are combined with laser and precision-reduced camera data can a final evaluation be made.

A few data were analyzed on a pass-by-pass (short arc) basis. This involved use of the orbit-computation program as an interpolation device, determining the parameters  $I$ ,  $e$ ,  $M_0$ , and  $n$ . The other orbital elements were held fixed at the values computed from long-arc computation. This interpolation will reject the groww outliers and will provide a measure of the noise (i. e., the precision) of the data. We have found this noise to be 50 to 100 cm. Figure 1 gives the typical residuals from a short-arc orbital fit.

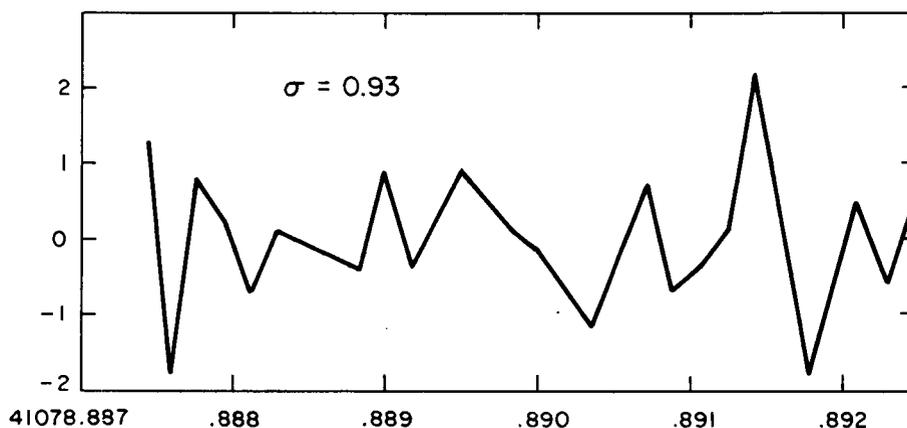


Figure 1. Residuals for short-arc fit of satellite 6508901 from station 7907.

It became apparent that the trend of residuals for 4-day arcs would provide the same information when plotted as the short-arc residuals do. Further, bad points not discarded by the short-arc procedure often resulted in poor short-arc fits. The longer arcs discarded these bad points. In addition, many passes had so few points (< 5) that short arcs were not possible. Finally, short-arc computation provided no estimate of the accuracy. We therefore abandoned this approach and proceeded to use 4-day arcs.

The SAO and the French laser data were input to our processing program in A.S time at reception of the laser pulse and without the refraction correction given. The refraction correction was computed following Tsiang and Lehr (this volume),

and time was converted to A. S at the satellite for orbital computation. The GSFC laser data were input in UTC time at the satellite and with refraction applied to range; again, time was converted to A. S at the satellite. We also used as inputs the tesseral harmonics determined in Standard Earth (II) (Gaposchkin and Lambeck, 1970), the station coordinates listed in Table 1 below, and the polar-motion values published by BIH. In Table 1, the station coordinates are given in megameters and the weight in arcseconds for all but the laser stations, which are in meters. Photoreduced weights where applicable are in parentheses.

Figure 2 gives the residuals of the same pass as in Figure 1 as it appears in a 4-day arc. The interpolation curve has been drawn to illustrate the short-arc fit. This particular pass was chosen because it was representative of SAO's ISAGEX laser data. It has a noise level of 0.93-m rms and an accuracy of 2.2-m rms.

Figure 3 gives the history of the semimajor axis for Geos 1 (6508901). The scatter appears to be less than 1 m, which indicates that the data are good, well-distributed, and well-understood. Geos 1 is a well-behaved satellite with an inclination of 59°. It is visible to many stations, thus giving rise to a nicely distributed data set. Its eccentricity of 0.071 presents no problems in the determination of the argument of perigee, and if we look at M2 as an indication of the magnitude of the modeled drag, we see that it is about  $5 \times 10^{-7}$  rev day<sup>-2</sup>, or 0.65 arcsec day<sup>-2</sup>.

Figure 4 gives the semimajor axis for Peole (7010901), with a scatter of less than 10 m. With an inclination of 15°, an eccentricity of 0.016, and M2 of  $3 \times 10^{-5}$  rev day<sup>-2</sup> (0.65 arcmin day<sup>-2</sup>), orbit computation for Peole is very challenging. Its inclination made it visible to very few stations (7907, 7929, 7060, 4492, 4800), four of which are in South America, giving rise to a badly distributed view of the satellite's orbit. Satellites with small inclinations make computation of both the argument of node and the argument of perigee difficult. The modeling of the drag and the solving for  $\dot{\omega}$  and  $\dot{\Omega}$  became impossible. We assumed values for  $\dot{\omega}$  and  $\dot{\Omega}$ , held them fixed, and solved for the drag modeling. This worked satisfactorily, but not so well as we had expected for laser data.

Table 1. Station coordinates used in the validation.

Station		COSPAR Number	X	Y	Z	Weight (arcsec)
Location						
<u>Baker-Nunn Stations</u>						
San Fernando, Spain	9004	5.105588	-0.555228	3.769667	34 (4)	
Naini Tal, India	9006	1.018203	5.471103	3.109623	34 (4)	
Maui, Hawaii	9012	-5.466053	-2.404282	2.242171	34 (4)	
Dakar, Senegal	9020	5.886264	-1.845649	1.615282	34 (4)	
Mt. Hopkins, Arizona	9021	-1.936782	-5.077704	3.331916	34 (4)	
Olifantsfontein, South Africa	9022	5.056125	2.716511	-2.775784	34 (4)	
Island Lagoon, Australia	9023	-3.977765	3.725101	-3.303034	34 (4)	
Dodaira, Japan	9025	-3.910438	3.376362	3.729219	34 (4)	
Arequipa, Peru	9027	1.943040	-5.804207	-1.796491	34 (4)	
Debre Zeit, Ethiopia	9028	4.903750	3.965201	0.963872	34 (4)	
Dionysos, Greece	9030	4.595200	2.039446	3.912606	34 (4)	
Natal, Brazil	9039	5.186461	-3.653856	-0.654325	34 (4)	
Rosamund, California (AF)	9113	-2.450011	-4.624421	3.635035	34 (4)	
Cold Lake, Canada (AF)	9114	-1.264838	-3.466884	5.185467	34 (4)	
Johnston Island (AF)	9117	-6.007402	-1.111859	1.825730	34 (4)	
Mt. John, New Zealand (AF)	9119	-4.533650	0.761590	-4.407772	34 (4)	
San Vito, Italy (AF)	9120	4.613757	1.485659	4.132293	34 (4)	
<u>Laser Stations</u>						(m)
GSFC, Maryland	7050	1.130673	-4.831368	3.994112	2	
Guam Island	7060	-5.068960	3.584106	1.458756	2	
Salisbury, Australia	7803	-3.939150	3.467040	-3.613265	2	
San Fernando, Spain	7804	5.105606	-0.555251	3.769633	2	
Haute Provence, France	7809	4.578352	0.457957	4.403160	2	
Haute Provence, France	7815	4.578371	0.457950	4.403134	2	
Dakar, Senegal	7820	5.886271	-1.845666	1.615250	2	
Olifantsfontein, South Africa	7902	5.056125	2.716511	-2.775784	2	
Arequipa, Peru	7907	1.942775	-5.804081	-1.796933	2	
Mt. Hopkins, Arizona	7921	-1.936781	-5.077701	3.331921	2	
Natal, Brazil	7929	5.186461	-3.653856	-0.654325	2	
Dionysos, Greece	7930	4.595207	2.039446	3.912595	2	

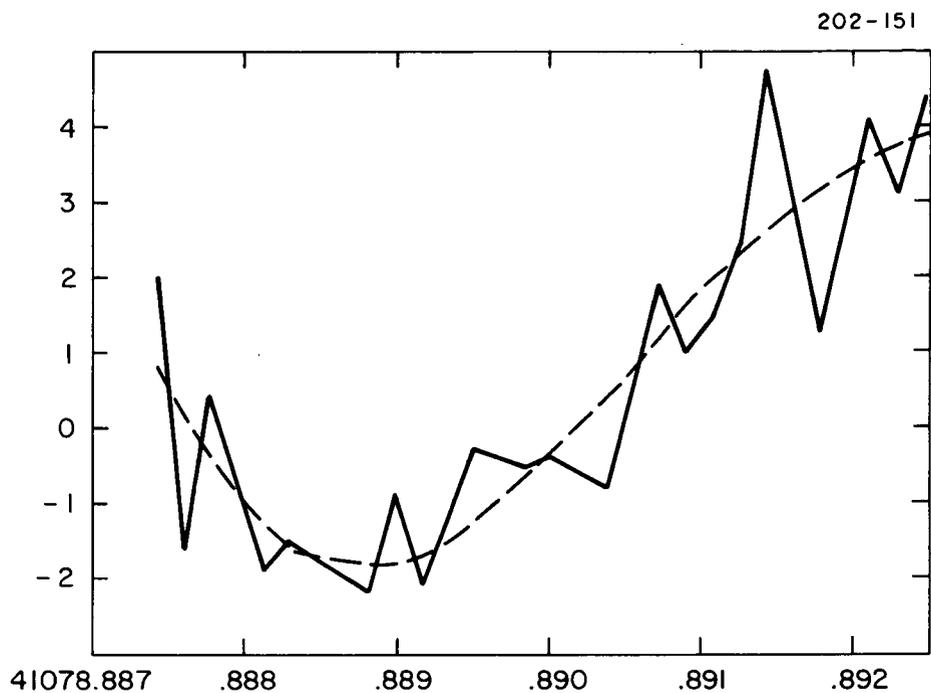


Figure 2. Residuals for 4-day fit for satellite 6508901 from station 7907.

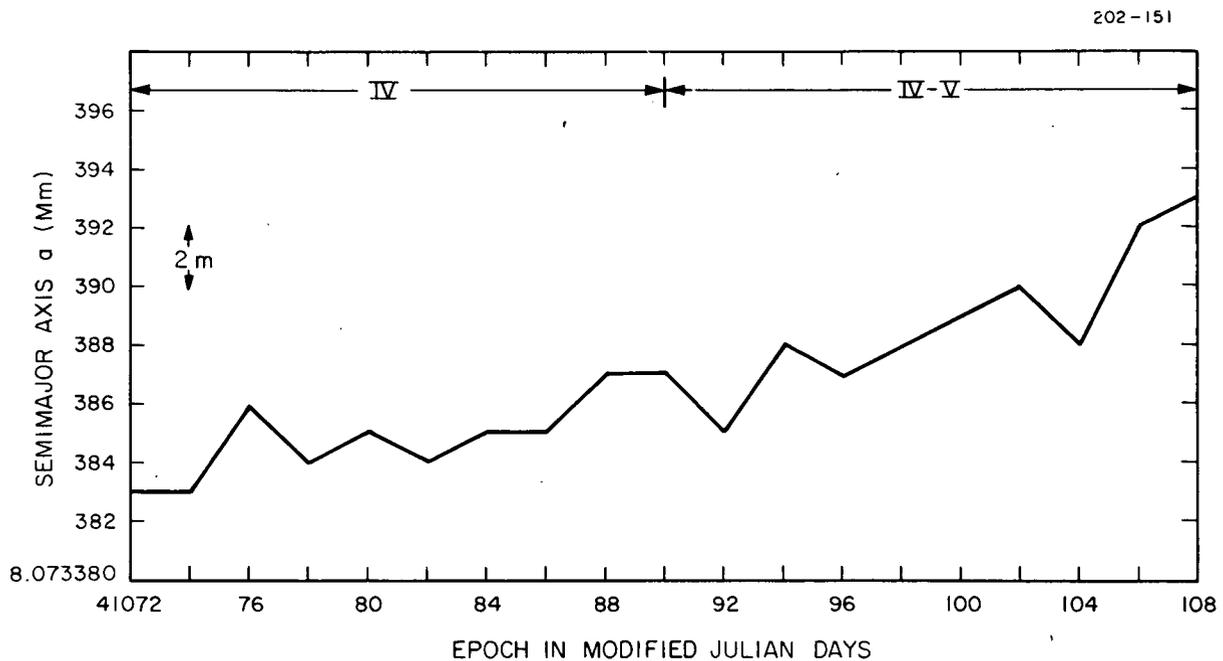


Figure 3. Evolution of semimajor axis for satellite 6508901 during ISAGEX periods IV and V.

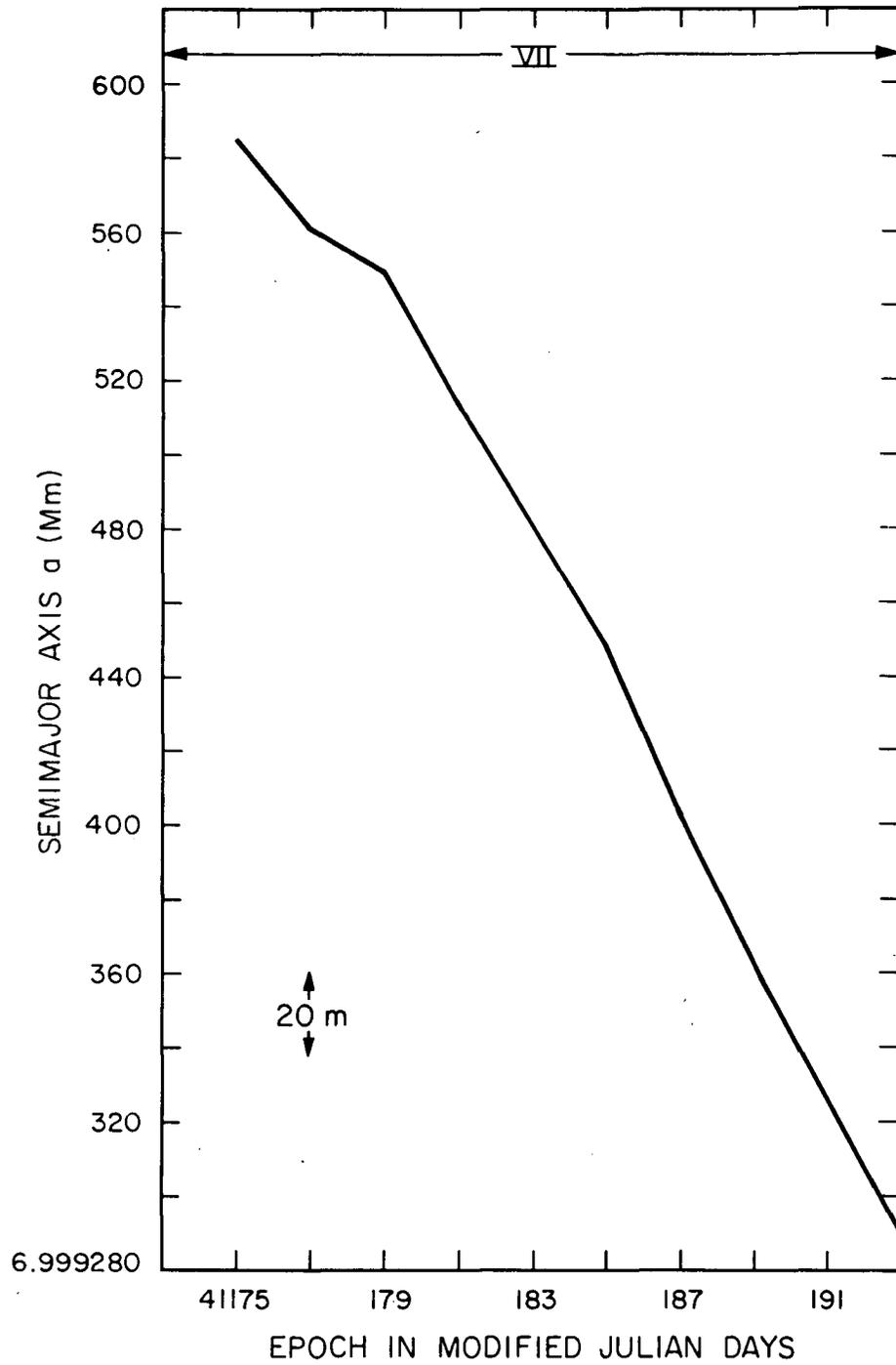


Figure 4. Evolution of semimajor axis for satellite 7010901 during ISAGEX period VII.

The scatter for the remaining satellites was about 3 m, except for BE-B (6406401), which had so few data that its scatter was about 7 m.

The number of validated SAO laser points broken down by station is given in Table 2. The ISAGEX periods covered by these data are listed in Table 3, where an X indicates a validated time period (Note: in-between time periods were also validated where possible).

Table 2. Number of validated SAO laser points.

Satellite	Station					Total
	7902	7907	7921	7929	7930	
6406401	-	12	13	5	-	30
6503201	3	162	3	41	33	242
6508901	3960	1408	164	528	106	6166
6701101	-	-	9	3	88	100
6701401	-	235	412	43	109	799
6800201	937	466	137	276	42	1858
7010901	-	35	-	230	-	265
Total	4900	2318	738	1126	378	9460

Table 3. Validated ISAGEX periods.

Satellite	II Feb. 15-Mar. 8 MJD 40997-41019		III Mar. 15-Apr. 15 41035-41057		IV Apr. 29-May 20 41070-41092		V June 5-26 41107-41129		VI July 14-31 41146-41164		VII Aug. 11-31 41174-41195
6406401			X								
6503201	X										X
6508901	X	X	X	X	X	X	X	X	X	X	X
6701101							X				
6701401					X			X			
6800201			X	X	X	X	X	X	X	X	X
7010901	X		X		X		X				X

As a result of our validation process, we can conclude the following:

A. The use of error signatures can be successfully employed to validate data.

By error signatures, we mean:

- 1) Large scatter in the data.
- 2) Inconsistent mean orbital elements.
- 3) Lack of systematic trends in the residuals.

To observe these signatures, we require more than 10 observations per pass.

B. The overall accuracy of the data is at least 1 m, and the noise level, 60 cm.

C. The overall accuracy of our geodesy is 10 m for Geos-type satellites, as can be seen from the standard error of unit weight for orbital fits as reported by Gaposchkin and Mendes (this volume). This accuracy confirms the evaluation of Gaposchkin and Lambeck (1970, 1971). The weight for laser data was taken as 2 m.

D. With routine validation by use of Geos 1 data, a monitoring of a station's reliability to the 10- to 20-m level is possible.

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**ORBITAL ELEMENTS FROM ISAGEX DATA**

E. M. Gaposchkin and G. M. Mendes

The process of data validation hinges on a consistent evolution of orbital elements derived from the data. These elements are also useful for other analyses – e. g., of zonal harmonics and earth tides – and are given here. Only those ISAGEX data available at SAO for validation were used; the orbital elements will be revised when the complete set of data has been processed. However, these orbits are an improvement over previous ones, especially for Peole (7010901), the first geodetic satellite with such a low inclination. This catalog of satellite data is similar to those previously published by SAO (see, e. g., Miller, 1968). The orbital elements are mean elements in the sense that the effects of the short-period perturbations due to the earth's gravity field have been eliminated.

The SAO mean elements have been computed from observations covering several days and are given in the form of a table. The successive sets of elements are essentially independent of each other. Only entries that are considered satisfactory are given. A missing epoch is due to insufficient data.

The times of epoch in the mean elements are reckoned in Julian Days. For convenience, the number 2400000.5 has been subtracted to provide an abbreviated rotation, which we call "Modified Julian Days" or MJD.

The units of the orbital elements are degrees for angular quantities, megameters for metric quantities, and revolutions for the mean anomaly.

The tabulated values of SAO mean elements are as follows:

line 1 Satellite designation, epoch, first and last dates, standard error of unit weight, number of observations, and date the orbit was computed.

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- line 2 Epoch.
- line 3  $\omega$ , argument of perigee and secular rate.
- line 4  $\Omega$ , right ascension of the ascending node and secular rate.
- line 5  $I$ , inclination.
- line 6  $e$ , eccentricity.
- line 7  $M$ , mean anomaly;  $n$ , mean motion; and higher polynomial term(s) as appropriate.

These elements include the long-period perturbations evaluated with the zonal harmonics as tabulated in Gaposchkin and Lambeck (1970, 1971); the short-period perturbations due to the geopotential computed with the numerical values given in Gaposchkin and Lambeck (1970, 1971); and the lunar perturbations with period  $2\lambda_{\zeta}$  computed as given by Gaposchkin (1966). The fundamental constants  $GM$ ,  $ae$ , and the velocity of light are as follows:

$$GM = 3.986013 \times 10^{20} \text{ cm}^3 \text{ sec}^{-2} \quad ,$$

$$ae = 6.378155 \times 10^8 \text{ cm} \quad ,$$

$$c = 2.997925 \times 10^{10} \text{ cm sec}^{-1} \quad .$$

The station coordinates used are given in Gaposchkin and Mendes (this volume).

The reference system adopted results in the inclination and the argument of perigee referred to the true equator of date. The right ascension of the ascending node is reckoned from the mean equinox of 1950.0 along the corresponding mean equator to the intersection with the moving true equator of date, and then along the true equator of date. To transform the right ascension of the node to the mean equinox of date, the following formula is used:

$$\Omega^{\circ} = \Omega^{\circ}(\text{SAO}) + 3:508 \times 10^{-5} (\text{MJD} - 33281) \quad .$$

The orbital theory used defines the mean elements. The orbit-computation program employed here is based on a set of formulas due to Aksnes (1970) for the short-period oblateness perturbations. The relationship between the mean elements from Von Zeipel's method as previously published and those by the Lie transform method is given by Aksnes (this volume).

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 1.500926 -16.303339 59.373286 .0716949 .9394274 11.9679466 8.0733855

ORBIT  
 1965 89 A 41047 41045.0 41049.0 1.3105 116 01/14/72  
 P111 P111 P111 P111 P31111  
 41047.000000 0.000000  
 1.8022521874 .6544407289  
 -20.8344327794 -2.2464421038  
 59.3732588603 -.0001168305  
 .0716553576 -.0000016078  
 .8781184221 11.9679474396 -5.0045E-09 2.3893E-07  
 41047.00000 1.3105402102  
 2.808097 -20.796253 59.373099 .0717209 .8753285 11.9679474 8.0733852

ORBIT  
 1965 89 A 41049 41047.0 41051.0 1.3758 137 01/14/72  
 P111 P111 P111 P111 P31111  
 41049.000000 0.000000  
 3.1102858564 .6537292700  
 -25.3274808035 -2.2465865986  
 59.3729093006 -.0002029073  
 .0716530377 -.0000009515  
 .8140170231 11.9679506693 1.1555E-07 -5.5883E-08  
 41049.00000 1.3758233697  
 4.114163 -25.289383 59.372680 .0717469 .8112325 11.9679507 8.0733837

ORBIT  
 1965 89 A 41051 41049.0 41053.0 2.5164 181 01/14/72  
 P111 P111 P111 P111 P31111  
 41051.000000 0.000000  
 4.4185444988 .6545542931  
 -29.8209310423 -2.2465537816  
 59.3719990431 -.0003747888  
 .0716512958 -.0000011978  
 .7499160411 11.9679476819 -1.1135E-07 8.4336E-08  
 41051.00000 2.5164165706  
 5.419873 -29.782943 59.371701 .0717733 .7471385 11.9679477 8.0733851

ORBIT  
 1965 89 A 41053 41051.0 41055.0 2.8445 184 01/14/72  
 P111 P111 P111 P111 P31111  
 41053.000000 0.000000  
 5.7248590068 .6541535045  
 -34.3136488904 -2.2466386061  
 59.3715222055 .0000381330  
 .0716513511 .0000008556  
 .6858211382 11.9679493798 1.1799E-07 -1.2221E-08  
 41053.00000 2.8445114693  
 6.723158 -34.275782 59.371156 .0718014 .6830510 11.9679494 8.0733843

ORBIT  
 1965 89 A 41055 41053.0 41057.0 1.5462 128 02/03/72  
 P111 P111 P111 P111 P31111  
 41055.000000 0.000000  
 7.0323120538 .6513063594  
 -38.8077941344 -2.2474430032  
 59.3685899714 -.0028630838  
 .0716531573 .0000017032  
 .6217230540 11.9679580643 -2.7396E-06 -8.3978E-07  
 \* 41055.00000 1.5461555473  
 8.026787 -38.770104 59.368154 .0718311 .6189645 11.9679587 8.0733804

ORBIT  
 1965 89 A 41057 41055.0 41059.0 .9513 67 02/03/72  
 P111 P111 P111 P111 P31111  
 41057.000000 0.000000  
 8.3283991521 .6515593066  
 -43.3002054204 -2.2446026509  
 59.3709814110 .0008316603  
 .0716729482 .0000147795  
 .5576521687 11.9679531328 5.1111E-06 1.8762E-06  
 \* 41057.00000 .9512889911  
 9.318916 -43.262620 59.370479 .0718786 .5549045 11.9679531 8.0733826

ORBIT  
 1965 89 A 41059 41057.0 41061.0 1.4005 117 02/03/72  
 P111 P111 P111 P111 P31111  
 41059.000000 0.000000  
 9.6497622792 .6499537325  
 -47.7944213904 -2.2466050068  
 59.3712110456 .0000456979  
 .0716455243 .0000004630  
 .4935176639 11.9679643397 -1.3494E-06 -8.1152E-07  
 \* 41059.00000 1.4004607545  
 10.636142 -47.757021 59.370640 .0718791 .4907815 11.9679643 8.0733776

ORBIT

1965 89 A	41061	41059.0	41063.0	1.5804	229 02/03/72
P111	P111	P111	P111	P31111	
41061.000000	0.000000				
10.9516000199	.6532533062				
-52.2880000010	-2.2468834093				
59.3707874372	-.0003636167				
.0716474174	.0000013151				
.4294358839	11.9679519170	2.6347E-07	1.2176E-07		

\* 41061.00000 1.5803995452

11.932945	-52.255795	59.370150	.0719084	.4267136	11.9679519	8.0733832
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ORBIT

1965 89 A	41063	41061.0	41065.0	2.4938	301 01/13/72
P111	P111	P111	P111	P31111	
41063.000000	0.000000				
12.2580141751	.6532163847				
-56.7817772531	-2.2468137335				
59.3703470704	-.0002194360				
.0716479078	-.0000091385				
.3653455553	11.9679526149	3.1806E-07	2.9954E-08		

41063.00000 2.4937655317

13.233839	-56.744787	59.369643	.0719363	.3626335	11.9679526	8.0733828
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ORBIT

1965 89 A	41065	41063.0	41067.0	2.1592	215 01/13/72
P111	P111	P111	P111	P31111	
41065.000000	0.000000				
13.5647201021	.6535308477				
-61.2756336349	-2.2470417341				
59.3701072907	-.0001218400				
.0716472498	-.00000025621				
.3012461438	11.9679524492	1.9302E-07	1.3355E-07		

41065.00000 2.1592126239

14.534573	-61.238875	59.369336	.0719628	.2985557	11.9679524	8.0733829
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ORBIT

1965 89 A	41067	41065.0	41069.0	1.6930	176 01/13/72
P111	P111	P111	P111	P31111	
41067.000000	0.000000				
14.8703756721	.6528661654				
-65.7696467107	-2.2470194382				
59.3699522960	-.0001967583				
.0716471307	.0000005021				
.2371544523	11.9679530346	-7.1119E-07	-6.1456E-08		

41067.00000 1.6930247634

15.833782	-65.733137	59.369116	.0719896	.2344818	11.9679530	8.0733827
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ORBIT

1965 89 A	41069	41067.0	41071.0	1.5885	293 01/13/72
P111	P111	P111	P111	P31111	
41069.000000	0.000000				
16.1760049841	.6525611643				
-70.2634158769	-2.2467396576				
59.3699302415	.0002597156				
.0716480476	.0000005095				
.1730589439	11.9679535819	1.5105E-07	-4.1236E-07		

41069.00000 1.5884815349

17.132485	-70.227170	59.369028	.0720172	.1704055	11.9679536	8.0733824
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ORBIT  
 1965 89 A 41071 41069.0 41073.0 1.3460 267 01/13/72  
 P111 P111 P01 P111 P31111  
 41071.000000 0.000000  
 17.4802449117 .6514510519  
 -74.7568908140 -2.2467096125  
 59.3704541214 .0005146499  
 .0716476416 -.0000017717  
 1.089666663 11.9679554372 1.2620E-06 7.4587E-07  
 41071.00000 1.3460419742  
 18.429408 -74.720919 59.369487 .0720433 .1063335 11.9679554 8.0733816

ORBIT  
 1965 89 A 41072 41070.0 41074.0 1.7923 276 12/04/71  
 P111 P111 P01 P111 P31111  
 41072.000000 0.000000  
 18.1340131645 .6528033447  
 282.9961156364 -2.2468871205  
 59.3698485969  
 .0716486714 -.0000001535  
 -.9230828252 11.9679523686 -2.1282E-07 -1.2003E-07

ORBIT  
 1965 89 A 41074 41072.0 41076.0 3.2673 139 12/04/71  
 P111 P111 P01 P111 P31111  
 41074.000000 0.000000  
 19.4389257007 .6522341453  
 278.5024353593 -2.2467493296  
 59.3699330823  
 .0716489502 .0000005135  
 .0128216132 11.9679517823 -1.1232E-07 1.7892E-07

ORBIT  
 1965 89 A 41076 41074.0 41078.0 1.8693 112 12/04/71  
 P111 P111 P01 P111 P31111  
 41076.000000 0.000000  
 20.7438759902 .6544555919  
 274.0089544126 -2.2466444494  
 59.3687177855  
 .0716501966 .0000028260  
 .9487249920 11.9679454462 -6.0260E-07 3.4567E-07

ORBIT  
 1965 89 A 41078 41076.0 41080.0 1.8894 232 12/04/71  
 P111 P111 P01 P111 P31111  
 41078.000000 0.000000  
 22.0492804199 .6529662419  
 269.5153612269 -2.2467910904  
 59.3700270576  
 .0716529050 .0000011703  
 .8846244713 11.9679493851 -1.0785E-07 -1.2480E-07

ORBIT  
 1965 89 A 41080 41078.0 41082.0 1.5273 241 12/04/71  
 P111 P111 P01 P111 P31111  
 41080.000000 0.000000  
 23.3551298339 .6528969802  
 265.0214562996 -2.2469722895  
 59.3698420366  
 .0716545634 .0000004951  
 .8205223404 11.9679480598 -3.2003E-07 8.2048E-08

ORBIT  
 1965 89 A 41082 41080.0 41084.0 .6775 168 12/04/71  
 P111 P111 P01 P111 P31111  
 41082.000000 0.000000  
 24.6595716954 .6514449953  
 260.5281397457 -2.2464301936  
 59.3696739148  
 .0716561002 .0000009250  
 .7564205973 11.9679503075 -4.1341E-07 9.5439E-09

ORBIT  
 1965 89 A 41084 41082.0 41086.0 1.6938 161 12/04/71  
 P111 P111 P01 P111 P31111  
 41084.000000 0.000000  
 25.9627467192 .6518248171  
 256.0351456372 -2.2465632918  
 59.3697614764  
 .0716577085 .0000005157  
 .6923188405 11.9679480421 -2.2778E-07 2.6027E-08

ORBIT  
 1965 89 A 41086 41084.0 41088.0 2.7325 126 12/04/71  
 P111 P111 P01 P111 P31111  
 41086.000000 0.000000  
 27.2698496966 .6514883038  
 251.5416652779 -2.2467241126  
 59.3704780175  
 .0716623352 .0000013975  
 .6282068086 11.9679484826 -4.5207E-07 1.1681E-07

ORBIT  
 1965 89 A 41088 41086.0 41090.0 1.8988 212 12/04/71  
 P111 P111 P01 P111 P31111  
 41088.000000 0.000000  
 28.5736916427 .6525234849  
 247.0483161918 -2.2467327303  
 59.3701739774  
 .0716643320 .0000020351  
 .5640998034 11.9679444686 -1.3559E-07 1.5556E-07

ORBIT  
 1965 89 A 41090 41088.0 41092.0 3.3913 372 12/04/71  
 P111 P111 P01 P111 P31111  
 41090.000000 0.000000  
 29.8788634541 .6525824216  
 242.5551298762 -2.2467558694  
 59.3697851942  
 .0716671300 .0000013158  
 .4999880537 11.9679443466 -4.6324E-08 -3.8353E-08

ORBIT  
 1965 89 A 41092 41090.0 41094.0 2.9677 279 12/11/71  
 P111 P111 P01 P111 P31111  
 41092.000000 0.000000  
 31.1825429218 .6511335911  
 238.0619261293 -2.2466669466  
 59.3696229375  
 .0716680296 .0000005668  
 .4358793353 11.9679469274 -3.6853E-07 -6.9064E-08

ORBIT  
 1965 89 A 41094 41092.0 41096.0 1.5351 177 12/11/71  
 P111 P111 P111 P111 P31111  
 41094.000000 0.000000  
 32.4882574705 .6532968541  
 233.5686819714 -2.2466708331  
 59.3694547130  
 .0716726653 .0000031099  
 .3717639919 11.9679400753 -5.4928E-07 1.5125E-07

ORBIT  
 1965 89 A 41096 41094.0 41098.0 1.4632 186 12/11/71  
 P111 P111 P111 P111 P31111  
 41096.000000 0.000000  
 33.7939155159 .6510990114  
 229.0757864449 -2.2465207879  
 59.3689237371  
 .0716784382 .0000010631  
 .3076442180 11.9679441562 3.4595E-07 -5.8870E-08

ORBIT  
 1965 89 A 41098 41096.0 41100.0 2.0047 218 12/11/71  
 P111 P111 P111 P111 P31111  
 41098.000000 0.000000  
 35.0978313954 .6522600093  
 224.5824696711 -2.2466852857  
 59.3688360163  
 .0716790845 .0000002841  
 .2435286961 11.9679415872 2.3227E-08 1.9063E-09

ORBIT  
 1965 89 A 41100 41098.0 41102.0 1.5152 173 12/11/71  
 P111 P111 P111 P111 P31111  
 41100.000000 0.000000  
 36.4032852544 .6536146337  
 220.0890595557 -2.2467686225  
 59.3686979838  
 .0716807809 .0000019611  
 .1794099015 11.9679383698 -3.7661E-07 -1.2845E-07

ORBIT  
 1965 89 A 41102 41100.0 41104.0 1.4889 200 12/11/71  
 P111 P111 P111 P111 P31111  
 41102.000000 0.000000  
 37.7088618846 .6534962328  
 215.5956562132 -2.2467840063  
 59.3682542945  
 .0716831044 .0000026975  
 .1152877896 11.9679366801 -3.4426E-07 6.4888E-08

ORBIT  
 1965 89 A 41104 41102.0 41106.0 1.6654 243 12/11/71  
 P111 P111 P111 P111 P31111  
 41104.000000 0.000000  
 39.0141344410 .6510600197  
 211.1024844213 -2.2465778436  
 59.3680476368  
 .0716861824 .0000003859  
 .0511633641 11.9679409199 7.3626E-09 1.1548E-07

ORBIT  
 1965 89 A 41106 41104.0 41108.0 3.6015 217 12/11/71  
 P111 P111 P111 P111 P31111  
 41106.000000 0.000000  
 40.3199970510 .6548664610  
 206.6091408276 -2.2467684207  
 59.3675953821  
 .0716884667 .0000014473  
 .9870377467 11.9679327079 -1.2908E-07 -3.7051E-08

ORBIT  
 1965 89 A 41108 41106.0 41110.0 2.1188 161 12/11/71  
 P111 P111 P111 P111 P31111  
 41108.000000 0.000000  
 41.6288031957 .6552610940  
 202.1159029239 -2.2464361853  
 59.3677591374  
 .0716913891 .0000015234  
 .9229037916 11.9679305985 3.7834E-07 3.1247E-07

ORBIT  
 1965 89 A 41109 41107.0 41111.0 .8999 154 12/11/71  
 P111 P111 P111 P111 P31111  
 41109.000000 0.000000  
 42.2798147313 .6516002760  
 199.8688715845 -2.2467538426  
 59.3673896797 -.0004961533  
 .0716902517 .0000006251  
 .8908459222 11.9679392837 -5.9330E-07 7.1863E-08

ORBIT  
 1965 89 A 41111 41109.0 41113.0 2.9877 202 12/11/71  
 P111 P111 P111 P111 P31111  
 41111.000000 0.000000  
 43.5842705512 .6507330261  
 195.3768863853 -2.2458666066  
 59.3680075786 .0009415216  
 .0716956542 .0000031057  
 .8267181762 11.9679399440 6.4951E-07 4.9738E-07

ORBIT  
 1965 89 A 41113 41111.0 41115.0 4.0083 150 12/11/71  
 P111 P111 P111 P111 P31111  
 41113.000000 0.000000  
 44.8859114197 .6512970124  
 190.8848588342 -2.2459227906  
 59.3691965267 -.0005920801  
 .0717021850 .0000032261  
 .7626037203 11.9679447677 -5.0988E-07 -9.4970E-07

ORBIT  
 1965 89 A 41115 41113.0 41117.0 1.5773 109 12/11/71  
 P111 P111 P111 P111 P31111  
 41115.000000 0.000000  
 46.1986949799 .6534143624  
 186.3877988692 -2.2467646364  
 59.3667548131 .0000302444  
 .0716984646 .0000013563  
 .6984623734 11.9679360157 4.7661E-08 7.0393E-08

ORBIT  
 1965 89 A 41117 41115.0 41119.0 3.6876 167 12/11/71  
 P111 P111 P111 P111 P31111  
 41117.000000 0.000000  
 47.5048353469 .6525417738  
 181.8943745239 -2.2467327786  
 59.3665121071 -.0004502082  
 .0717000742 .0000006053  
 .6343356553 11.9679364493 -6.2644E-07 3.9011E-07

ORBIT  
 1965 89 A 41119 41117.0 41121.0 3.6235 279 12/16/71  
 P111 P111 P111 P111 P31111  
 41119.000000 0.000000  
 48.8098507214 .5526562048  
 177.4009105583 -2.2467790053  
 59.3661471949 -.0000132545  
 .0717012477 .0000010703  
 .5702090532 11.9679365698 1.3765E-08 1.7954E-07

ORBIT  
 1965 89 A 41121 41119.0 41123.0 2.9878 266 12/16/71  
 P111 P111 P111 P111 P31111  
 41121.000000 0.000000  
 50.1158620319 .6535159213  
 172.9071902892 -2.2469387595  
 59.3659725324 -.0002100756  
 .0717036078 .0000012324  
 .5060820029 11.9679362915 2.9944E-07 -1.3911E-07

ORBIT  
 1965 89 A 41123 41121.0 41125.0 2.0173 247 12/16/71  
 P111 P111 P111 P111 P31111  
 41123.000000 0.000000  
 51.4238054578 .6544772316  
 168.4133743931 -2.2468239918  
 59.3661355540 .0007967926  
 .0717065840 .0000029015  
 .4419519762 11.9679320772 -6.5165E-07 2.2641E-07

ORBIT  
 1965 89 A 41125 41123.0 41127.0 .6945 266 12/17/71  
 P111 P111 P111 P111 P31111  
 41125.000000 0.000000  
 52.7306498980 .6529132386  
 163.9198475398 -2.2469780646  
 59.3667319858 -.0000213629  
 .0717097224 .0000013247  
 .3778215474 11.9679362855 9.0392E-08 1.3741E-07

ORBIT  
 1965 89 A 41127 41125.0 41129.0 1.3255 281 12/17/71  
 P111 P111 P111 P111 P31111  
 41127.000000 0.000000  
 54.0365759785 .6526401891  
 159.4259757686 -2.2467550427  
 59.3666913108 .0000521010  
 .0717121391 .0000010249  
 .3136948081 11.9679377458 -2.7955E-08 -8.9658E-08

ORBIT  
 1965 89 A 41129 41127.0 41131.0 1.0456 230 12/17/71  
 P111 P111 P111 P111 P31111  
 41129.000000 0.000000  
 55.3442644787 .6537675374  
 154.9328557852 -2.2467255330  
 59.3668851608 -.0000397414  
 .0717138423 .0000004855  
 .2495629177 11.9679352404 3.6513E-07 3.3235E-08

ORBIT  
 1965 89 A 41131 41129.0 41133.0 1.0251 282 12/17/71  
 P111 P111 P111 P111 P31111  
 41131.000000 0.000000  
 56.6515121893 .6528568132  
 150.4396282656 -2.2467345643  
 59.3672522123 .0002088604  
 .0717155549 .0000026897  
 .1854332033 11.9679364401 -2.4447E-07 -5.7245E-08

ORBIT  
 1965 89 A 41133 41131.0 41135.0 .6963 276 12/17/71  
 P111 P111 P111 P111 P31111  
 41133.000000 0.000000  
 57.9554987035 .6512891775  
 145.9461495527 -2.2468798474  
 59.3683123624 -.0000967632  
 .0717200085 -.0000000529  
 .1213113955 11.9679422901 4.0673E-07 7.9023E-08

ORBIT  
 1965 89 A 41135 41133.0 41137.0 .5596 309 12/17/71  
 P111 P111 P111 P111 P31111  
 41135.000000 0.000000  
 59.2643674231 .6543798420  
 141.4523630073 -2.2468232884  
 59.3679207926 .0000359774  
 .0717242371 .0000023437  
 .0571770801 11.9679338630 3.9573E-07 -5.4746E-08

ORBIT  
 1965 89 A 41137 41135.0 41139.0 .7398 280 12/17/71  
 P111 P111 P111 P111 P31111  
 41137.000000 0.000000  
 60.5729333489 .6541329076  
 136.9587181135 -2.2468234369  
 59.3680432144 .0000785516  
 .0717287383 .0000020823  
 .9930464621 11.9679353259 -5.5362E-08 -6.0557E-08

ORBIT  
 1965 89 A 41139 41137.0 41141.0 1.0888 175 12/17/71  
 P111 P111 P111 P111 P31111  
 41139.000000 0.000000  
 61.8789242812 .6524536304  
 132.4659932253 -2.2466035295  
 59.3685064215 .0000935001  
 .0717317157 .0000022328  
 .9289211778 11.9679396736 8.6926E-07 3.0424E-07

ORBIT						
1965 89 A		41141	41139.0	41143.0	.8429	186 12/17/71
P111	P111	P111	P111	P31111		
41141.000000		0.000000				
63.1835766155		.6526719629				
127.9726817755		-2.2467579845				
59.3693403156		-.0003393011				
.0717329997		.0000001910				
.8648030555	11.9679398813		7.9964E-07		5.2324E-07	

ORBIT						
1965 89 A		41143	41141.0	41145.0	1.7455	320 12/17/71
P111	P111	P111	P111	P31111		
41143.000000		0.000000				
64.4923300448		.6545133949				
123.4789138610		-2.2466703441				
59.3691297683		.0000518560				
.0717401517		.0000026463				
.8006771492	11.9679366524		-3.9050E-07		1.7784E-07	

ORBIT						
1965 89 A		41145	41143.0	41147.0	2.5000	404 12/17/71
P111	P111	P111	P111	P31111		
41145.000000		0.000000				
65.8003951835		.6535609976				
118.9856468452		-2.2466456253				
59.3694009128		.0001471941				
.0717449166		.0000022003				
.7365533567	11.9679394325		-1.5930E-07		1.3793E-07	

ORBIT						
1965 89 A		41147	41145.0	41149.0	1.9375	263 12/17/71
P111	P111	P111	P111	P31111		
41147.006000		0.000000				
67.1082721201		.6543551189				
114.4925316912		-2.2463930455				
59.3697121733		.0002039850				
.0717508788		.0000046416				
.6724296545	11.9679367713		5.2797E-07		2.0303E-07	

ORBIT						
1965 89 A		41148	41146.0	41150.0	1.6072	326 12/10/71
P111	P111	P111	P111	P31111		
41148.000000		0.000000				
67.7598510484		.6534045377				
112.2456287296		-2.2466140515				
59.3697315564		.0001489188				
.0717510644		.0000025345				
.6403779587	11.9679419386		4.1135E-07		3.0145E-08	

ORBIT						
1965 89 A		41150	41148.0	41152.0	1.0417	337 12/10/71
P111	P111	P111	P111	P31111		
41150.000000		0.000000				
69.0683209024		.6547442488				
107.7521544933		-2.2467307141				
59.3697290506		.0000094462				
.0717573047		.0000036445				
.5762588331	11.9679392226		2.0719E-08		2.6181E-08	

ORBIT  
 1965 89 A 41152 41150.0 41154.0 1.1569 323 12/10/71  
 P111 P111 P111 P111 P31111  
 41152.000000 0.000000  
 70.3777546489 .6533160138  
 103.2589076680 -2.2465899879  
 59.3699466534 .0001869836  
 .0717642944 .0000014419  
 .5121375171 11.9679441828 -8.6169E-08 2.2560E-08

ORBIT  
 1965 89 A 41154 41152.0 41156.0 1.4729 306 12/10/71  
 P111 P111 P111 P111 P31111  
 41154.000000 0.000000  
 71.6840368334 .6528971412  
 98.7659103009 -2.2464245343  
 59.3703393843 .0002333871  
 .0717671133 .0000017002  
 .4489265161 11.9679456973 5.6532E-07 6.6601E-08

ORBIT  
 1965 89 A 41156 41154.0 41158.0 2.1665 299 12/11/71  
 P111 P111 P111 P111 P31111  
 41156.000000 0.000000  
 72.9923446285 .6542728183  
 94.2721419226 -2.2469353902  
 59.3698508275 -.0001872203  
 .0717717667 .0000022464  
 .3839137922 11.9679446459 5.3899E-07 -4.5837E-08

ORBIT  
 1965 89 A 41158 41156.0 41160.0 2.9329 274 12/11/71  
 P111 P111 P111 P111 P31111  
 41158.000000 0.000000  
 74.3001224450 .6543001763  
 89.7783148753 -2.2469897871  
 59.3696244490 -.0001441708  
 .0717747144 .0000022825  
 .3198075602 11.9679458281 -4.1306E-08 -5.7599E-08

ORBIT  
 1965 89 A 41160 41158.0 41162.0 1.8840 178 12/11/71  
 P111 P111 P111 P111 P31111  
 41160.000000 0.000000  
 75.6072029973 .6519773751  
 85.2850894517 -2.2464553540  
 59.3702768002 .0005723497  
 .0717758987 -.0000017291  
 .2557022928 11.9679541040 1.0629E-06 -6.0841E-07

ORBIT  
 1965 89 A 41162 41160.0 41164.0 .5269 152 01/12/72  
 P111 P111 P111 P111 P31111  
 41162.000000 0.000000  
 76.9123987425 .6574393102  
 80.7885007980 -2.2456337223  
 59.3669265137 .0013908160  
 .0717631575 .0000079177  
 .1916129659 11.9679338781 3.3441E-07 -5.6575E-08  
 41162.000000 .5263948739  
 77.132124 80.796815 59.363999 .0729594 .1910031 11.9679330 8.0733913

ORBIT  
 1965 89 A 41164 41162.0 41166.0 1.0622 262 01/12/72  
 P111 P111 P111 P111 P31111  
 41164.00000 0.000000  
 78.2269279977 .6555709826  
 76.2972194295 -2.2469894626  
 59.3695117318 -.0002021650  
 .0717856317 .0000036465  
 .1274835931 11.9679447545 2.3658E-07 -2.3194E-07  
 41164.00000 1.0621823137  
 78.424876 76.304720 59.366568 .0729880 .1269342 11.9679448 8.0733864

ORBIT  
 1965 89 A 41166 41164.0 41168.0 1.1114 263 12/16/71  
 P111 P111 P111 P111 P31111  
 41166.00000 0.000000  
 79.5330930917 .6568957114  
 71.8030798001 -2.2468583134  
 59.3692884982 .0002842296  
 .0717846601 .0000060571  
 .0633922563 11.9679385875 -8.5354E-07 6.0239E-07

ORBIT  
 1965 89 A 41168 41166.0 41170.0 2.6150 346 12/16/71  
 P111 P111 P111 P111 P31111  
 41168.00000 0.000000  
 80.8446910161 .6530626218  
 67.3097219338 -2.2469030135  
 59.3694501227 -.0001651678  
 .0717936569 .0000002311  
 .9992775282 11.9679532054 2.4843E-07 9.2726E-08

ORBIT  
 1965 89 A 41170 41168.0 41172.0 1.5592 344 12/16/71  
 P111 P111 P111 P111 P31111  
 41170.00000 0.000000  
 82.1526860781 .6546334782  
 62.8155273819 -2.2472956051  
 59.3692594287 .0000213397  
 .0717947501 .0000009672  
 .9351802972 11.9679504751 3.5427E-07 -4.6053E-08

ORBIT  
 1965 89 A 41172 41170.0 41174.0 .8129 209 12/16/71  
 P111 P111 P111 P111 P31111  
 41172.00000 0.000000  
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 58.3208698359 -2.2473523947  
 59.3692017530 -.0000117250  
 .0717966458 .0000007382  
 .8710811613 11.9679494997 -1.6663E-06 -6.7173E-07

ORBIT  
 1965 89 A 41174 41172.0 41176.0 1.1012 156 12/16/71  
 P111 P111 P111 P111 P31111  
 41174.00000 0.000000  
 84.7751162512 .6511062586  
 53.8275268353 -2.2478848892  
 59.3680063325 .0017589276  
 .0717975599 .0000013217  
 .8069714200 11.9679623898 -1.6665E-06 2.6992E-07

ORBIT  
 1965 89 A 41176 41174.0 41178.0 4.2435 351 12/10/71  
 P111 P111 P111 P111 P31111  
 41176.000000 0.000000  
 86.0793682969 .6534585898  
 49.3332532007 -2.2469122406  
 59.3702872307 -.0001147772  
 .0717987292 .0000005386  
 .7428873555 11.9679525143 4.3335E-07 -1.9621E-08

ORBIT  
 1965 89 A 41178 41176.0 41180.0 3.4796 564 12/10/71  
 P111 P111 P111 P111 P31111  
 41178.000000 0.000000  
 87.3860339024 .6546237936  
 44.8393743263 -2.2470368755  
 59.3701305327 .0001238538  
 .0718000771 .0000004955  
 .6787916623 11.9679504923 1.0305E-07 -1.7853E-08

ORBIT  
 1965 89 A 41180 41178.0 41182.0 2.2467 621 12/10/71  
 P111 P111 P111 P111 P31111  
 41180.000000 0.000000  
 88.6952361979 .6548831572  
 40.3452219188 -2.2471406814  
 59.3703852350 .0002313166  
 .0718001763 -.0000006122  
 .6146931557 11.9679499664 -1.0443E-07 -7.1395E-08

ORBIT  
 1965 89 A 41182 41180.0 41184.0 2.0390 446 12/10/71  
 P111 P111 P111 P111 P31111  
 41182.000000 0.000000  
 90.0030894671 .6537994533  
 35.8511403792 -2.2470105685  
 59.3708672561 .0002200939  
 .0717987706 -.0000006479  
 .5505969917 11.9679517504 -1.5334E-07 5.1091E-08

ORBIT  
 1965 89 A 41184 41182.0 41186.0 3.2032 433 12/10/71  
 P111 P111 P111 P111 P31111  
 41184.000000 0.000000  
 91.3110874994 .6537109940  
 31.3574488706 -2.2467744257  
 59.3711394147 .0000579032  
 .0717970395 .0000003707  
 .4864989502 11.9679524682 -6.4152E-08 -1.2695E-07

ORBIT  
 1965 89 A 41186 41184.0 41188.0 1.5891 371 12/16/71  
 P111 P111 P111 P111 P31111  
 41186.000000 0.000000  
 92.6183244882 .6532713471  
 26.8638747976 -2.2469288542  
 59.3713256234 .0002844896  
 .0717971082 -.00000033438  
 .4224035008 11.9679532342 -2.6162E-07 -2.8709E-07

ORBIT  
 1965 89 A 41188 41186.0 41190.0 1.6454 294 12/16/71  
 P111 P111 P111 P111 P31111  
 41188.000000 0.000000  
 93.9255450243 .6538565998  
 22.3700253458 -2.2469536302  
 59.3720213968 .0004088878  
 .0717954205 -.0000011700  
 .3583056597 11.9679496475 -1.9109E-07 6.1989E-08

ORBIT  
 1965 89 A 41190 41188.0 41192.0 2.7197 255 12/16/71  
 P111 P111 P111 P111 P31111  
 41190.000000 0.000000  
 95.2327930926 .6532854269  
 17.8763672720 -2.2464857385  
 59.3726182285 -.0000893820  
 .0717928747 -.0000010151  
 .2942052946 11.9679499137 -2.4174E-07 1.9352E-08

ORBIT  
 1965 89 A 41192 41190.0 41194.0 1.7918 266 12/16/71  
 P111 P111 P111 P111 P31111  
 41192.000000 0.000000  
 96.5389365013 .6529594193  
 13.3835433564 -2.2467624809  
 59.3722183844 .0002053864  
 .0717910332 -.0000008978  
 .2301054638 11.9679510670 -6.0498E-08 -1.4803E-07

ORBIT  
 1965 89 A 41194 41192.0 41196.0 1.3598 260 12/16/71  
 P111 P111 P111 P111 P31111  
 41194.000000 0.000000  
 97.8452294152 .6532221871  
 8.8896669160 -2.2472261476  
 59.3729312424 .0008847676  
 .0717887777 -.0000019952  
 .1660054211 11.9679493345 -4.9995E-07 -2.6585E-07

ORBIT  
 1967 II A 41111 41109.0 41113.0 3.3796 51 12/30/71  
 P111 P111 P01 P111 P31111  
 41111.000000 0.000000  
 219.1412813552 5.9834012503  
 231.3521619735 -4.7384303320  
 39.9621862732  
 .0512050626 -.0000025413  
 -.9179076443 13.8602244502 5.4541E-06 -3.9829E-07

ORBIT  
 1967 II A 41113 41111.0 41115.0 5.2336 90 12/30/71  
 P111 P111 P01 P111 P31111  
 41113.000000 0.000000  
 231.0998791644 5.9875854473  
 221.8648803545 -4.7415544215  
 39.9593428911  
 .0512063728 -.0000031031  
 .8026103373 13.8602401276 6.4319E-06 1.4176E-07

ORBIT  
 1967 II A 41115 41113.0 41117.0 2.6217 106 12/30/71  
 P111 P111 P01 P111 P31111  
 41115.000000 0.000000  
 243.0869880137 5.9824316854  
 212.3797858405 -4.7440671888  
 39.9582628423  
 .0512002971 .0000023592  
 .5230798244 13.8602809476 4.6073E-06 4.9439E-07

ORBIT  
 1967 II A 41117 41115.0 41119.0 2.3661 125 12/30/71  
 P111 P111 P01 P111 P31111  
 41117.000000 0.000000  
 255.0535338676 5.9840096233  
 202.8927200475 -4.7429100453  
 39.9583904001  
 .0512029935 .0000003330  
 .2436557435 13.8602953203 4.2929E-06 -9.6002E-07

ORBIT  
 1967 II A 41119 41117.0 41121.0 2.4694 262 02/07/72  
 P111 P111 P01 P111 P31111  
 41119.000000 0.000000  
 267.0185666717 5.9828510563  
 193.4075623389 -4.7433000459  
 39.9582685020  
 .0512033952 .0000000220  
 .9642636773 13.8603175790 4.8333E-06 -8.3451E-07  
 \* 41119.00000 2.4693908979  
 266.988775 193.407417 39.960144 .0506690 .9643469 13.8603175 7.3207118

ORBIT  
 1967 II A 41121 41119.0 41123.0 2.6791 288 02/07/72  
 P111 P111 P01 P111 P31111  
 41121.000000 0.000000  
 278.9842804257 5.9838915964  
 183.9209141956 -4.7432149815  
 39.9582167241  
 .0512030195 -.0000000742  
 .6849119189 13.8603263311 4.0381E-06 6.7345E-07

\* 41121.00000 2.6790791398  
 279.073796 183.921352 39.960074 .0506740 .6846618 13.8603263 7.3207087

ORBIT  
 1967 11 A 41123 41121.0 41125.0 2.2058 198 02/07/72  
 P111 P111 P11 P11 P31111  
 41123.000000 0.000000  
 290.9557243466 5.9854022425  
 174.4347652172 -4.7427801671  
 39.9582377538  
 .0512031898 -.0000002239  
 .4055737068 13.8603383858 3.6940E-06 -7.3818E-07  
 \* 41123.00000 2.2057914277  
 291.161503 174.435770 39.960903 .0507003 .4049987 13.8603384 7.3207045

ORBIT  
 1967 11 A 41125 41123.0 41127.0 2.1089 189 02/07/72  
 P111 P111 P11 P11 P31111  
 41125.000000 0.000000  
 302.9307628393 5.9919821858  
 164.9491135219 -4.7428761273  
 39.9581501280  
 .0512022339 -.0000007460  
 .1262545565 13.8603369708 6.5994E-06 7.1384E-07  
 \* 41125.00000 2.1089314531  
 303.245598 164.950646 39.959752 .0507459 .1253749 13.8603370 7.3207050

ORBIT  
 1967 11 A 41127 41125.0 41129.0 2.5996 179 02/07/72  
 P111 P111 P11 P11 P31111  
 41127.000000 0.000000  
 314.9010183468 5.9852055295  
 155.4589862729 -4.7437775564  
 39.9593869082  
 .0511998217 -.0000016276  
 .8469983191 13.8613787505 8.0150E-06 1.5864E-06  
 \* 41127.00000 2.5896001786  
 315.313500 155.460988 39.960759 .0508089 .8458460 13.8603788 7.3206903

ORBIT  
 6701401 41072 41070.0 41074.0 3.0511 221 01/04/72  
 P111 P111 P01 P111 P31111  
 41072.000000 0.000000  
 151.5922726442 5.4102751498  
 168.2220309987 -4.2226935477  
 39.4323738573  
 .0840063953 -.0000017948  
 -.6557712736 13.0935086672 1.7557E-06 3.2077E-07

ORBIT  
 6701401 41074 41072.0 41076.0 2.9861 256 01/04/72  
 P111 P111 P01 P111 P31111  
 41074.000000 0.000000  
 162.4127744247 5.4103556512  
 159.7766831038 -4.2226137172  
 39.4323254836  
 .08400630351 -.0000016851  
 .5312559170 13.0935194587 3.5281E-06 1.4367E-07

ORBIT  
 6701401 41076 41074.0 41078.0 4.5980 218 01/04/72  
 P111 P111 P01 P111 P31111  
 41076.000000 0.000000  
 173.2356992522 5.4115259269  
 151.3312189519 -4.2227118985  
 39.4321122184  
 .0839981327 -.0000022750  
 .7183040149 13.0935321334 4.1558E-06 -2.9122E-08

ORBIT  
 6701401 41078 41076.0 41080.0 2.7782 263 01/04/72  
 P111 P111 P01 P111 P31111  
 41078.000000 0.000000  
 184.0597274343 5.4125156843  
 142.8859204657 -4.2226366699  
 39.4322276296  
 .0839958510 -.0000020238  
 .9053834529 13.0935484783 5.0901E-06 -1.8311E-07

ORBIT  
 6701401 41080 41078.0 41082.0 1.9301 262 01/04/72  
 P111 P111 P01 P111 P31111  
 41080.000000 0.000000  
 194.8851157284 5.4122663622  
 134.4405633183 -4.2227517022  
 39.4322913959  
 .0839925474 -.0000013064  
 .0924979783 13.0935637248 2.2402E-06 -6.0045E-07

ORBIT  
 6701401 41082 41080.0 41084.0 3.7237 265 01/04/72  
 P111 P111 P01 P111 P31111  
 41082.000000 0.000000  
 205.7108830798 5.4123688883  
 125.9948451196 -4.2228431891  
 39.4321397588  
 .0839893669 -.0000013424  
 .2796317090 13.0935693383 1.4881E-06 4.8919E-07

ORBIT  
 6701401 41084 41082.0 41086.0 3.9848 242 01/04/72  
 P111 P111 P111 P111 P31111  
 41084.000000 0.000000  
 216.5325465823 5.4105059598  
 117.5495556974 -4.2226265066  
 39.4320777365  
 .0839856797 -.0000014595  
 .4667774204 13.0935745152 1.0572E-06 5.4817E-07

ORBIT  
 6701401 41086 41084.0 41088.0 2.6045 118 01/04/72  
 P111 P111 P111 P111 P31111  
 41086.000000 0.000000  
 227.3529050653 5.4104304581  
 109.1042450757 -4.2224448856  
 39.4324515688  
 .0839817229 -.0000011586  
 .6539352525 13.0935821170 2.9104E-06 4.7381E-07

ORBIT  
 6701401 41088 41086.0 41090.0 2.9464 81 01/04/72  
 P111 P111 P111 P111 P31111  
 41088.000000 0.000000  
 238.1727459342 5.4093113816  
 100.6605732508 -4.2219375412  
 39.4322378316  
 .0839798125 -.0000015482  
 .8411145165 13.0935968753 3.9027E-06 8.3131E-07

ORBIT  
 6701401 41090 41088.0 41092.0 2.9730 105 01/04/72  
 P111 P111 P111 P111 P31111  
 41090.000000 0.000000  
 248.9958839267 5.4107968115  
 92.2139570525 -4.2219769856  
 39.4336273981  
 .0839791284 -.0000015346  
 .0283224446 13.0936139487 4.2632E-06 -3.9050E-07

ORBIT  
 6701401 41148 41146.0 41150.0 1.73 223 12/28/71  
 P111 P111 P111 P111 P2111  
 41148.000000 0.000000  
 202.8333448540 5.4103091846  
 207.3126284934 -4.2231986041  
 39.4314439178  
 .0838908996 .0000018741  
 .4627741899 13.0937969155 9.3962E-07

ORBIT  
 6701401 41150 41148.0 41152.0 1.12 200 12/28/71  
 P111 P111 P111 P111 P2111  
 41150.000000 0.000000  
 213.6549020564 5.4111086027  
 198.8677183541 -4.2220575244  
 39.4315115389  
 .0838937164 -.0000050702  
 .6503600765 13.0937978914 4.0251E-06

ORBIT

6701401		41152	41150.0	41154.0	1.93	140 12/28/71
P111	P111	P11	P111	P2111		
41152.000000		0.000000				
224.4740511289		5.4584984115				
190.4221617227		-4.2258281968				
39.4311231517						
.0838878469		.0000084194				
.8379885632		13.0938240538		9.7049E-06		

ORBIT

6701401		41154	41152.0	41156.0	1.75	101 12/28/71
P111	P111	P11	P111	P2111		
41154.000000		0.000000				
235.2943647136		5.4093728258				
181.9801108984		-4.2211591133				
39.4311001482						
.0838892281		.0000017979				
.0256235282		13.0938257147		2.8133E-06		

ORBIT  
 68 002 01 41037 41035.0 41039.0 2.6290 121 01/13/72  
 P111 P111 P111 P111 P31111  
 41037.000000 0.000000  
 70.5557863676 -1.6229134021  
 241.6939204418 1.4221953764  
 105.8049278790 .0034039321  
 .0319771143 -.0000066691  
 .7005154170 12.8306052717 1.9850E-06 -1.7780E-06  
 41037.00000 2.6289534070  
 71.076043 241.692711 105.805355 .0327989 .6990705 12.8306053 7.7073317

ORBIT  
 68 002 01 41039 41037.0 41041.0 1.5576 165 01/13/72  
 P111 P111 P111 P111 P31111  
 41039.000000 0.000000  
 67.3342694598 -1.6251825597  
 244.4998491656 1.3999496948  
 105.8102362650 -.0021716896  
 .0319770677 -.0000036586  
 .3617722561 12.8306554249 -6.8598E-08 -9.8375E-09  
 41039.00000 1.5576300815  
 67.936364 244.498448 105.810654 .0327810 .3601000 12.8306654 7.7073076

ORBIT  
 68 002 01 41041 41039.0 41043.0 1.2975 183 01/13/72  
 P111 P111 P111 P111 P31111  
 41041.000000 0.000000  
 64.0824108109 -1.6261944378  
 247.2993328206 1.4000384736  
 105.8084640684 -.0000848689  
 .0319698798 -.0000001878  
 .0231064983 12.8306697521 5.9624E-07 1.6325E-07  
 41041.00000 1.2975326434  
 64.765472 247.297745 105.808871 .0327534 .0212094 12.8306698 7.7073059

ORBIT  
 68 002 01 41043 41041.0 41045.0 2.2158 147 01/13/72  
 P111 P111 P111 P111 P31111  
 41043.000000 0.000000  
 60.8214587189 -1.5321711422  
 250.1019213372 1.4014846411  
 105.8128717012 .0019313604  
 .0319558971 -.0000076201  
 .6844686197 12.8306892842 3.9275E-06 9.8511E-07  
 41043.00000 2.2157920317  
 61.583513 250.100150 105.813266 .0327163 .6823521 12.8306893 7.7072980

ORBIT  
 68 002 01 41045 41043.0 41047.0 1.1427 128 01/13/72  
 P111 P111 P111 P111 P31111  
 41045.000000 0.000000  
 57.5929597668 -1.6265393527  
 252.9026826268 1.4004494207  
 105.8100226195 .0006168129  
 .0319702633 -.0000008005  
 .3457484619 12.8306710563 -3.3434E-07 3.8313E-08  
 41045.00000 1.1427438684  
 58.430303 252.900735 105.810404 .0327055 .3434228 12.8306711 7.7073053

ORBIT  
 68 002 01 41047 41045.0 41049.0 1.3492 127 01/13/72  
 P111 P111 P111 P111 P31111  
 41047.000000 0.000000  
 54.354275333 -1.6254184041  
 255.7047331554 1.4014459745  
 105.8105538974 .0012168118  
 .0319759837 -.0000057659  
 .0070543353 12.8306674757 4.8006E-08 -4.8076E-08  
 41047.00000 1.349237359  
 55.254531 255.702620 105.810921 .0326835 .0045262 12.8306675 7.7073068

ORBIT  
 68 002 01 41049 41047.0 41051.0 1.6834 116 01/13/72  
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 105.8122440256 .0017080433  
 .0319634077 -.0000076945  
 .6683959745 12.8306669518 7.7014E-07 -7.0759E-08  
 41049.00000 1.6834144756  
 52.081254 258.502682 105.812596 .0326408 .6656712 12.8306670 7.7073070

ORBIT  
 68 002 01 41051 41049.0 41053.0 1.8307 157 02/03/72  
 P111 P111 P111 P111 P31111  
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 261.3078714605 1.4003926778  
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 .0319729815 .0000002990  
 .3297065582 12.8306570886 -2.3192E-06 1.2150E-06  
 \* 41051.00000 1.8306591010  
 48.907421 261.305433 105.812557 .0326183 .3267965 12.8306571 7.7073109

ORBIT  
 68 002 01 41053 41051.0 41055.0 1.9435 173 02/03/72  
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 .0319724281 -.0000007478  
 .9910233660 12.8306655854 1.7113E-07 -1.0947E-07  
 \* 41053.00000 1.9434511755  
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ORBIT  
 68 002 01 41055 41053.0 41057.0 2.4572 152 02/03/72  
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 105.8125728922 -.0004969453  
 .0319744125 -.0000030700  
 .6523323936 12.8306592303 4.5741E-06 2.4715E-06  
 \* 41055.00000 2.4571745047  
 42.548581 266.908602 105.812871 .0325493 .6490789 12.8306592 7.7073101

ORBIT  
 68 002 01 41057 41055.0 41059.0 .6505 134 02/03/72  
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 .0319885193 .0000091129  
 .3136095429 12.8306421620 4.0412E-06 1.4962E-06  
 \* 41057.00000 .6504556028  
 39.374337 269.709999 105.814178 .0325255 .3102019 12.8306422 7.7073169

ORBIT  
 68 002 01 41059 41057.0 41061.0 1.0207 149 02/03/72  
 P111 P111 P111 P111 P31111  
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 272.5153065552 1.4012963478  
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 .9749517794 12.8306484786 -7.2234E-07 8.4616E-07  
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ORBIT  
 68 002 01 41061 41059.0 41063.0 1.2253 154 02/03/72  
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 .6362973356 12.8306927849 -1.3690E-06 -2.2206E-07  
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ORBIT  
 68 002 01 41063 41061.0 41065.0 .9891 170 02/03/72  
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 .2976062257 12.8306589950 4.4788E-07 4.7841E-07  
 \* 41063.00000 .9891206337  
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ORBIT  
 68 002 01 41065 41063.0 41067.0 1.7648 139 01/13/72  
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 .9589281975 12.8306667317 -8.9646E-08 -6.0274E-08  
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ORBIT  
 68 002 01 41067 41065.0 41069.0 1.6022 135 01/13/72  
 P111 P111 P111 P111 P31111  
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 .0319854719 -.0000006278  
 .6202412598 12.8306568717 -7.3033E-07 -6.3234E-08  
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ORBIT  
 68 002 01 41069 41067.0 41071.0 1.8625 207 01/13/72  
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 .2815606009 12.8306607977 -1.7666E-07 9.1321E-07  
 41069.00000 1.8625005083  
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ORBIT  
 68 002 01 41071 41069.0 41073.0 1.7092 166 01/13/72  
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ORBIT  
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ORBIT  
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 P111 P111 P111 P111 P31111  
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ORBIT  
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 P111 P111 P111 P111 P31111  
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ORBIT  
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 P111 P111 P111 P111 P31111  
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ORBIT  
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 P111 P111 P111 P111 P31111  
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ORBIT  
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 P111 P111 P111 P111 P31111  
 41082.000000 0.000000  
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 304.7389251584 1.4008134678  
 105.8164132031  
 .0319935786 .0000010182  
 .0800865211 12.8306482554 1.6009E-06 -7.8328E-07

ORBIT  
 1968 02 A 41084 41082.0 41086.0 1.6841 174 12/11/71  
 P111 P111 P111 P111 P31111  
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 307.5399185082 1.4004521785  
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 .0319959802 .0000028155  
 .7414900515 12.8306654727 5.6716E-06 2.2177E-06

ORBIT  
 1968 02 A 41086 41084.0 41088.0 1.1253 106 12/11/71  
 P111 P111 P111 P111 P31111  
 41086.000000 0.000000  
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 .0319979320 -.0000051096  
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ORBIT  
 1968 02 A 41088 41086.0 41090.0 3.7206 195 12/16/71  
 P111 P111 P111 P111 P31111  
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ORBIT  
 1968 02 A 41090 41088.0 41092.0 5.0266 261 12/16/71  
 P111 P111 P111 P111 P31111  
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ORBIT  
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ORBIT  
 1968 02 A 41096 41094.0 41098.0 .3375 51 12/16/71  
 P111 P111 P111 P111 P31111  
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ORBIT  
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 P111 P111 P111 P111 P31111  
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ORBIT  
 68 02 A 41109 41107.0 41111.0 30.4275 64 12/10/71  
 P111 P111 P111 P111 P31111  
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 314.0001977436 -2.1297616666  
 342.5859900640 1.3998013846  
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ORBIT  
 68 02 A 41111 41109.0 41113.0 .8857 101 12/10/71  
 P111 P111 P111 P111 P31111  
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 345.3704388467 1.4012725915  
 105.8171196885  
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ORBIT  
 68 02 A 41113 41111.0 41115.0 2.1307 102 12/10/71  
 P111 P111 P111 P111 P31111  
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 348.1728690371 1.4010299522  
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ORBIT  
 68 02 A 41115 41113.0 41117.0 3.4264 67 12/10/71  
 P111 P111 P111 P111 P31111  
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 350.9753708467 1.4011612853  
 105.8175931052  
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 .4917443110 12.8306545855 -7.2814E-07 -4.1947E-07

ORBIT  
 68 02 A 41117 41115.0 41119.0 1.8140 236 12/10/71  
 P111 P111 P111 P111 P31111  
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 353.7770909792 1.4011317206  
 105.8171569622  
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 .1530784388 12.8306688790 -5.4310E-07 5.0270E-08

ORBIT  
 68 02 A 41119 41117.0 41121.0 4.0519 276 12/10/71  
 P111 P111 P111 P111 P31111  
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 356.5792988548 1.4010533924  
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 .0319801030 .0000001552  
 .8144122964 12.8306637208 1.0872E-07 5.1313E-08

ORBIT  
 68 02 A 41121 41119.0 41123.0 3.7784 141 12/10/71  
 P111 P111 P111 P111 P31111  
 41121.000000 0.000000  
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 359.3813587671 1.4009648645  
 105.8173346756  
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 .4757421622 12.8306608775 -1.5492E-07 -4.4726E-08

ORBIT  
 68 02 A 41123 41121.0 41125.0 1.5319 132 12/10/71  
 P111 P111 P111 P111 P31111  
 41123.000000 0.000000  
 291.2253927804 -1.6233676822  
 2.1836573916 1.4011124853  
 105.8171233723  
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 .1370677160 12.8306704905 -6.2835E-07 -1.0463E-06

68 02 A	41125	41123.0	41127.0	1.4492	145 12/16/71
P111	P111	P01	P111	P31111	
41125.000000	0.000000				
287.981649129	-1.6175453743				
4.9857162522	1.4011397874				
105.8167060712					
.0319766192	-.0000039955				
.7983943770	12.8306487748		1.2809E-06	1.3224E-06	

ORBIT

68 02 A	41127	41125.0	41129.0	1.0555	138 12/16/71
P111	P111	P01	P111	P31111	
41127.000000	0.000000				
284.7331132713	-1.6200516034				
7.7875031766	1.4010879128				
105.8166941424					
.0319746024	.0000007468				
.4597333204	12.8306588805		3.4639E-07	2.3825E-08	

ORBIT

68 02 A	41129	41127.0	41131.0	1.0779	121 12/16/71
P111	P111	P01	P111	P31111	
41129.000000	0.000000				
281.4932634516	-1.6141846779				
10.5908368436	1.4009210004				
105.8171033462					
.0319823069	-.0000000082				
.1210529830	12.8306376129		-2.1138E-07	2.2080E-06	

ORBIT

1968 02 A	41148	41146.0	41150.0	1.2823	141 12/10/71
P111	P111	P111	P111	P31111	
41148.000000	0.000000				
250.6758197652	-1.6247783691				
37.2099031685	1.4014147067				
105.8156102960	.0008249557				
.0319784024	.0000005664				
.9036728028	12.8306726306		-3.3693E-07	-2.9624E-09	

ORBIT

1968 02 A	41150	41148.0	41152.0	1.1658	130 12/10/71
P111	P111	P111	P111	P31111	
41150.000000	0.000000				
247.4248009747	-1.6253107916				
40.0131277658	1.4016171520				
105.8176545937	.0010575693				
.0319757106	-.0000015949				
.5650196825	12.8306726326		3.8982E-07	2.4922E-07	

ORBIT

1968 02 A	41152	41150.0	41154.0	.9790	97 12/10/71
P111	P111	P111	P111	P31111	
41152.000000	0.000000				
244.1911591274	-1.6284051224				
42.8143108132	1.4013649157				
105.8157926019	.0008003756				
.0319849244	.0000002918				
.2263225180	12.8306829955		-4.8174E-07	-5.7769E-07	

ORBIT  
 1968 02 A 41154 41152.0 41156.0 1.0207 132 12/10/71  
 P111 P111 P111 P111 P31111  
 41154.000000 0.000000  
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 45.6179561377 1.3979560605  
 105.8162621228 -.0039749646  
 .0319961227 .0000063997  
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ORBIT  
 1968 02 A 41156 41154.0 41158.0 1.3741 121 12/10/71  
 P111 P111 P111 P111 P31111  
 41156.000000 0.000000  
 237.7368920012 -1.6072366890  
 48.4185635757 1.4009657010  
 105.8152859670 -.0002896322  
 .0319868819 -.0000025116  
 .5488802710 12.8306167864 7.6309E-07 1.0995E-06

ORBIT  
 1968 02 A 41158 41156.0 41160.0 1.2467 92 12/10/71  
 P111 P111 P111 P111 P31111  
 41158.000000 0.000000  
 234.4855706435 -1.6131910581  
 51.2226658199 1.4027567286  
 105.8138352824 -.0021005369  
 .0319995906 .0000034515  
 .2102263727 12.8306346172 7.4098E-06 5.6897E-06

ORBIT  
 1968 02 A 41160 41158.0 41162.0 1.0978 123 12/10/71  
 P111 P111 P111 P111 P31111  
 41160.000000 0.000000  
 231.2250055446 -1.6086743907  
 54.0217701441 1.4002331545  
 105.8142698518 -.0009872071  
 .0319909938 .0000015503  
 .8715998492 12.8306278137 6.6465E-07 -3.0033E-07

ORBIT  
 1968 02 A 41162 41160.0 41164.0 .8289 114 12/10/71  
 P111 P111 P111 P111 P31111  
 41162.000000 0.000000  
 227.9844958583 -1.6075865045  
 56.8247287885 1.4007035780  
 105.8156889538 .0002933462  
 .0319910729 .0000046070  
 .5329193610 12.8306240386 -1.2694E-06 8.9561E-07

ORBIT  
 1968 02 A 41164 41162.0 41166.0 1.7609 131 12/10/71  
 P111 P111 P111 P111 P31111  
 41164.000000 0.000000  
 224.7670634809 -1.6118718502  
 59.6271437481 1.4014248403  
 105.8152527760 -.0003257848  
 .0320017272 .0000053146  
 .1941743423 12.8306181713 -4.3300E-06 1.0711E-05

ORBIT  
 1968 02 A 41166 41164.0 41168.0 1.3246 125 12/17/71  
 P111 P111 P111 P111 P31111  
 41166.000000 0.000000  
 221.5230627918 -1.6249846245  
 62.4283648228 1.4009351312  
 105.8154582164 .0015519410  
 .0319877816 .0000047009  
 .8554970137 12.8306738218 3.6309E-07 -8.4733E-07

ORBIT  
 1968 02 A 41168 41166.0 41170.0 1.4568 129 12/17/71  
 P111 P111 P111 P111 P31111  
 41168.000000 0.000000  
 218.2647903200 -1.6277198147  
 65.2312001450 1.3994246307  
 105.8137704249 .0003445700  
 .0319869990 -.0000052719  
 .5168666218 12.8306891560 -2.6527E-06 -5.4344E-06

ORBIT  
 1968 02 A 41170 41168.0 41172.0 1.0953 127 12/17/71  
 P111 P111 P111 P111 P31111  
 41170.000000 0.000000  
 215.0783671363 -1.5773316028  
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 105.8127909273 -.0033324271  
 .0319624978 -.0000285471  
 .1779973238 12.8305250684 1.9849E-05 7.7586E-06

ORBIT  
 1968 02 A 41172 41170.0 41174.0 .9700 113 12/17/71  
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 41172.000000 0.000000  
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 70.8341296928 1.4005183817  
 105.8147676219 -.0005294646  
 .0319933172 .0000035353  
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ORBIT  
 1968 02 A 41174 41172.0 41176.0 1.2141 140 12/17/71  
 P111 P111 P111 P111 P31111  
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 73.6357726523 1.4009895492  
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 .0319895985 .0000036173  
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ORBIT  
 1968 02 A 41176 41174.0 41178.0 .8422 147 12/11/71  
 P111 P111 P111 P111 P31111  
 41176.000000 0.000000  
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 .0319930418 -.0000011489  
 .1620965901 12.8306464157 -3.7554E-07 2.1328E-07

ORBIT  
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 P111 P111 P111 P111 P31111  
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 .0319915956 .0000015594  
 .8234078014 12.8306573919 -1.1583E-07 5.4302E-09

ORBIT  
 1968 02 A 41180 41178.0 41182.0 2.6784 225 12/16/71  
 P111 P111 P111 P111 P31111  
 41180.000000 0.000000  
 198.8272338977 -1.6198143937  
 82.0408040283 1.4009949188  
 105.8122472003 -.0001081578  
 .0319940089 .0000002110  
 .4847223538 12.8306553348 4.2228E-08 -7.7658E-08

ORBIT  
 1968 02 A 41182 41180.0 41184.0 3.6024 245 12/16/71  
 P111 P111 P111 P111 P31111  
 41182.000000 0.000000  
 195.5913041050 -1.6187206571  
 84.8424009475 1.4007197774  
 105.8112470942 -.0002406234  
 .0319924100 .0000006178  
 .1460226079 12.8306511153 -2.8686E-07 5.0067E-08

ORBIT  
 1968 02 A 41184 41182.0 41186.0 2.6971 229 12/16/71  
 P111 P111 P111 P111 P31111  
 41184.000000 0.000000  
 192.3691063922 -1.6062366303  
 87.6437401823 1.4008227242  
 105.8108983558 .0000885715  
 .0319962666 .0000019513  
 .8072847933 12.8306169714 4.8816E-09 1.5317E-06

ORBIT  
 1968 02 A 41188 41186.0 41190.0 3.1229 162 12/16/71  
 P111 P111 P111 P111 P31111  
 41188.000000 0.000000  
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 93.2467705959 1.4003027959  
 105.8107200093 -.0000662665  
 .0319909740 -.0000016763  
 .1299456230 12.8306419495 5.0357E-08 -1.8295E-06

ORBIT  
 1968 02 A 41190 41188.0 41192.0 1.7981 163 12/16/71  
 P111 P111 P111 P111 P31111  
 41190.000000 0.000000  
 182.6330696346 -1.6198838032  
 96.0479681931 1.4005618227  
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 .0319909079 .0000006527  
 .7912451380 12.8306528341 -3.8208E-08 2.0140E-08

ORBIT						
1968 02 A	41192	41190.0	41194.0	1.8506	254 12/16/71	
P111	P111	P111	P111	P31111		
41192.000000	0.000000					
179.3921345141	-1.6179499412					
98.8497994447	1.4010168533					
105.8103107128	.0003864645					
.0319913600	.0000000204					
.4525549167	12.8306478488		-5.9879E-07	2.7902E-07		

ORBIT						
1968 02 A	41194	41192.0	41196.0	1.8419	179 12/16/71	
P111	P111	P111	P111	P31111		
41194.000000	0.000000					
176.1557495179	-1.6144640776					
101.6515777746	1.4002010042					
105.8103976578	-.0004754438					
.0319913413	.0000022781					
-.1138510514	12.8306356725		1.3538E-06	2.9663E-06		

ORBIT  
 1970 109 A 41008 41006.0 41010.0 5.6095 89 01/27/72  
 P110 P110 P01 P111 P2111  
 41009.000000 -1.000000  
 113.9976101141 13.0699780000  
 40.1426252113 -6.9858900000  
 15.0044491333  
 .0166053083 -.0001003948  
 -.9252147843 14.8162110555 4.7799E-05  
 41008.00000 5.6095030529  
 113.685947 40.140942 15.003715 .0169121 .0756566 14.8162111 7.0023532

ORBIT  
 1970 109 A 41010 41008.0 41012.0 4.8217 98 01/27/72  
 P110 P110 P01 P111 P2111  
 41010.000000 .000000  
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 15.0037254664  
 .0163618383 -.0001211353  
 .7074090058 14.8163574633 2.9465E-05  
 41010.00000 4.8216520160  
 139.524257 26.166577 15.003213 .0165082 .7090864 14.8163575 7.0023071

ORBIT  
 1970 109 A 41012 41010.0 41014.0 2.2372 65 01/27/72  
 P110 P110 P01 P111 P2111  
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 15.0025986709  
 .0165052933 -.0000521316  
 .3387453979 14.8164992639 3.6126E-05  
 41012.00000 2.2372442199  
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ORBIT  
 1970 109 A 41014 41012.0 41016.0 .6192 41 01/27/72  
 P110 P110 P01 P111 P2111  
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 15.0017593658  
 .0164334004 -.0000355307  
 .9718856193 14.8166391045 3.0637E-05  
 41014.00000 .6191936770  
 192.135180 -1.698220 15.001933 .0163840 .9740347 14.8166391 7.0022183

ORBIT  
 1970 109 A 41016 41014.0 41018.0 1.3377 36 01/27/72  
 P110 P110 P01 P111 P2111  
 41016.000000 .000000  
 218.8455178157 13.0699780000  
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 41016.00000 1.3377433259  
 218.231240 -15.657947 14.991363 .0163704 .6074835 14.8167001 7.0021991

ORBIT  
 1970 109 A 41018 41016.0 41020.0 1.9996 30 01/27/72  
 P110 P110 P01 P111 P2111  
 41018.000000 .000000  
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 41018.00000 1.9995933979  
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ORBIT  
 1970 109 A 41020 41018.0 41022.0 2.0134 85 01/27/72  
 P110 P110 P01 P111 P2111  
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 41020.00000 2.0134058133  
 272.642926 -43.607101 15.008649 .0161931 .8689197 14.8168958 7.0021375

ORBIT  
 1970 109 A 41038 41036.0 41040.0 .8583 28 01/27/72  
 P110 P110 P01 P111 P2111  
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 -.4256554425 14.8176512168 3.3790E-05  
 41038.00000 .8582907454  
 150.017648 190.880564 14.992056 .0166141 .5762867 14.8176012 7.0019152

ORBIT  
 1970 109 A 41040 41038.0 41042.0 .7905 14 01/27/72  
 P110 P110 P01 P111 P2111  
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 41040.00000 .7804545015  
 176.176491 176.896531 14.996487 .0163855 .2125555 14.8177491 7.0018686

ORBIT  
 1970 109 A 41042 41040.0 41044.0 2.8939 50 01/27/72  
 P110 P110 P01 P111 P2111  
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 .8454033090 14.8178828633 4.4414E-05  
 41042.00000 2.8939663027  
 202.866874 162.973657 14.997768 .0162466 .8474397 14.8178829 7.0018265

ORBIT

1970 109 A	41044	41042.0	41046.0	1.0042	31 01/27/72
P110	P110	P01	P111	P2111	
41044.000000	.000000				
229.9226581558	13.2313000000				
149.043209805	-6.9791030000				
14.9958412161					
.0163243337	-.0000780984				
.4816972316	14.8180543923		3.2508E-05		
41044.00000	1.0041568058				
229.407910	149.040762	14.996450	.0161499	.4831353	14.8180544 7.0017725

ORBIT

1970 109 A	41046	41044.0	41048.0	1.3568	77 01/27/72
P110	P110	P01	P111	P2111	
41046.000000	.000000				
256.7004040101	13.2313000000				
135.0745054124	-6.9791030000				
14.9947531524					
.0163056117	.0000165588				
.1171541725	14.8182676546		4.9134E-05		
41046.00000	1.3560215758				
256.515836	135.073642	14.995530	.0160829	.1176698	14.8182677 7.0017053

ORBIT

1970 109 A	41048	41046.0	41050.0	1.3237	54 01/27/72
P110	P110	P01	P111	P2111	
41048.000000	.000000				
283.2032100914	13.2313000000				
121.1217506888	-6.9791030000				
14.9926911854					
.0162915214	-.0000119598				
.7537623959	14.8184709852		4.8760E-05		
41048.00000	1.3237049448				
283.386594	121.122657	14.993468	.0160687	.7532501	14.8184710 7.0016412

ORBIT

1970 109 A	41050	41048.0	41052.0	.5989	32 01/27/72
P110	P110	P01	P111	P2111	
41050.000000	.000000				
309.7159209709	13.2313000000				
107.1522774642	-6.9791030000				
15.0112820179					
.0162767959	.0000195378				
.3907837941	14.8186856512		6.4437E-05		
41050.00000	.5988923626				
310.228916	107.154697	15.011893	.0161012	.3893507	14.8186857 7.0015736

ORBIT

1970 109 A	41052	41050.0	41054.0	.6394	44 01/27/72
P110	P110	P01	P111	P2111	
41052.000000	.000000				
336.2479283687	13.2313000000				
93.1647466374	-6.9791030000				
15.0142015670					
.0162904078	.0000125002				
.0283002030	14.8189431450		6.4272E-05		
41052.00000	.6394289474				
336.978902	93.168269	15.014517	.0161997	.0262579	14.8189431 7.0014925

ORBIT  
 1970 109 A 41054 41052.0 41056.0 1.0063 84 01/27/72  
 P110 P110 P01 P111 P2111  
 41054.000000 .000000  
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 .6662522403 14.8191906766 6.3572E-05  
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ORBIT  
 1970 109 A 41056 41054.0 41058.0 1.2202 86 01/27/72  
 P110 P110 P01 P111 P2111  
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 .3048175320 14.8194815092 7.8292E-05  
 41056.00000 1.2202134497  
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ORBIT  
 1970 109 A 41078 41076.0 41080.0 1.2164 94 01/26/72  
 P110 P110 P01 P111 P2111  
 41078.000000 0.000000  
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 .3576925923 14.8212304127 3.2133E-05  
 41078.00000 1.2164003330  
 321.216815 271.772595 14.998149 .0160657 .3559574 14.8212304 7.0007722

ORBIT  
 1970 109 A 41080 41078.0 41082.0 1.7718 111 01/26/72  
 P110 P110 P01 P111 P2111  
 41080.000000 0.000000  
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 .0002864568 14.8213324141 1.2854E-05  
 41080.00000 1.7717995816  
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ORBIT  
 1970 109 A 41082 41080.0 41084.0 2.1962 80 01/26/72  
 P110 P110 P01 P111 P2111  
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 .6429352798 14.8214496016 5.4071E-05  
 41082.00000 2.1962220098  
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ORBIT  
 1970 109 A 41084 41082.0 41086.0 1.7420 65 01/26/72  
 P110 P110 P01 P111 P2111  
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 .2857767936 14.8214937713 2.1241E-05  
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ORBIT  
 1970 109 A 41086 41084.0 41088.0 .8018 76 01/26/72  
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 .9286924360 14.8215920256 2.1360E-05  
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ORBIT  
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 P110 P110 P01 P111 P2111  
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 .5719566099 14.8216961565 3.3062E-05  
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ORBIT  
 1970 109 A 41090 41088.0 41092.0 1.5620 90 01/26/72  
 P110 P110 P01 P111 P2111  
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 .2162916908 14.8218355316 3.1293E-05  
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ORBIT  
 1970 109 A 41109 41107.0 41111.0 .7052 76 01/27/72  
 P110 P110 P01 P111 P2111  
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 11.931971 55.364535 14.994110 .0162449 .8339935 14.8226019 7.0003406

1970 109 A 41111 41109.0 41113.0 .7817 82 01/27/72  
 P110 P110 P01 P111 P2111  
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 38.241582 41.406734 14.993301 .0163905 .4797253 14.8227128 7.0003054

ORBIT  
 1970 109 A 41113 41111.0 41115.0 .7937 78 01/27/72  
 P110 P110 P01 P111 P2111  
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ORBIT  
 1970 109 A 41115 41113.0 41117.0 2.1062 93 01/27/72  
 P110 P110 P01 P111 P2111  
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 41115.00000 2.1062156613  
 90.455404 13.474905 14.998905 .0164883 .7730461 14.8229311 7.0002367

ORBIT  
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 P110 P110 P01 P111 P2111  
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 41117.00000 12.8867483437  
 116.541315 -.487216 14.999100 .0164509 .4200650 14.8230331 7.0002046

ORBIT  
 1970 109 A 41119 41117.0 41121.0 1.4955 118 01/27/72  
 P110 P110 P01 P111 P2111  
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 41119.00000 1.4955271590  
 142.743686 -14.455768 14.998372 .0163911 .0669788 14.8231333 7.0001730

ORBIT  
 1970 109 A 41121 41119.0 41123.0 1.4929 97 01/27/72  
 P110 P110 P01 P111 P2111  
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ORBIT  
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ORBIT  
 1970 109 A 41175 41173.0 41177.0 1.7166 57 01/27/72  
 P110 P110 P01 P111 P2111  
 41175.000000 -1.000000  
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 14.9957520698  
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ORBIT  
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 P110 P110 P01 P111 P2111  
 41177.000000 .000000  
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 .0161470748 -.0000305658  
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 41177.00000 2.3690513880  
 190.905303 300.545841 14.997201 .0161025 .8662310 14.8250775 6.9995610

ORBIT  
 1970 109 A 41179 41177.0 41181.0 2.2346 62 01/27/72  
 P110 P110 P01 P111 P2111  
 41179.000000 .000000  
 218.2623337740 13.2313000000  
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 14.9955314669  
 .0161679425 .0000007012  
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 41179.00000 2.2346122867  
 217.629957 286.574580 14.996017 .0160275 .5157094 14.8251155 6.9995490

ORBIT  
 1970 109 A 41181 41179.0 41183.0 1.5320 42 01/27/72  
 P110 P110 P01 P111 P2111  
 41181.000000 .000000  
 244.6891722027 13.2313000000  
 272.6105980294 -6.9843190000  
 15.0030392811  
 .0160325894 -.0001026940  
 .1644126216 14.8252301318 2.9862E-05  
 41181.00000 1.5320326982  
 244.340515 272.609013 15.003748 .0158259 .1653865 14.8252301 6.9995130

ORBIT  
 1970 109 A 41183 41181.0 41185.0 2.3427 36 01/27/72  
 P110 P110 P01 P111 P2111  
 41183.000000 .000000  
 271.0531340676 13.2313000000  
 258.6292653723 -6.9843190000  
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 .0161951745 -.0000721248  
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 41183.00000 2.3426983513  
 271.067998 258.629334 15.013695 .0159660 .8152298 14.8253324 6.9994808

ORBIT  
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 41185.000000 .000000  
 297.8733021328 13.2313000000  
 244.6692965364 -6.9843190000  
 15.0050750695  
 .0160777369 -.0000369165  
 .4650331142 14.8254363985 3.1718E-05  
 41185.00000 .7959201566  
 298.253454 244.671036 15.005770 .0158757 .4639713 14.8254364 6.9994480

ORBIT  
 1970 109 A 41187 41185.0 41189.0 .4898 72 01/27/72  
 P110 P110 P01 P111 P2111  
 41187.000000 .000000  
 324.4608090375 13.2313000000  
 230.6941616555 -6.9843190000  
 15.0120282462  
 .0160638066 .0000099147  
 .1157142991 14.8255816758 3.6436E-05  
 41187.00000 .4897865172  
 325.120986 230.697227 15.012481 .0159319 .1138701 14.8255817 6.9994023

ORBIT  
 1970 109 A 41189 41187.0 41191.0 1.1768 60 01/27/72  
 P110 P110 P01 P111 P2111  
 41189.000000 .000000  
 351.0431980811 13.2313000000  
 216.7308527916 -6.9843190000  
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 .0160920548 .0000471117  
 .7666769253 14.8257669377 1.9887E-05  
 41189.00000 1.1768441673  
 351.837659 216.734635 14.996890 .0160582 .7644573 14.8257069 6.9993629

ORBIT  
 1970 109 A 41191 41189.0 41193.0 1.1241 48 01/27/72  
 P110 P110 P01 P111 P2111  
 41191.000000 .000000  
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 .0161546547 -.0000172657  
 .4180872685 14.8258272354 4.4253E-05  
 41191.00000 1.1241308961  
 18.261910 202.806746 15.000103 .0162245 .4159664 14.8258272 6.9993250

ORBIT  
 1970 109 A 41193 41191.0 41195.0 .8775 66 01/27/72  
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 41193.00000 .8775494227  
 44.597739 188.821874 14.997573 .0163461 .0681486 14.8259427 6.9992889

A NOTE ON THE RELATIONSHIP AND AGREEMENT BETWEEN  
TWO SATELLITE THEORIES

K. Aksnes

We shall refer to the two theories under consideration as theory A (Gaposchkin, Cherniack, Briggs, and Benima, 1971; Kozai, 1962) and theory B (Aksnes, 1970). The main purpose here is to introduce some simple theoretical relations by means of which, when the elements of one theory are given, we can predict those of the other.

Theory B differs from theory A in that the former makes use of 1) a reference orbit that is a rotating ellipse (intermediate orbit) instead of a fixed ellipse, 2) Hill variables instead of Delaunay variables, and 3) Hori's method in Lie series rather than Von Zeipel's method in Taylor series.

In both theories, the periodic perturbations are expressed as deviations from a mean orbit whose elements are semimajor axis  $a$ , eccentricity  $e$ , mean anomaly  $M$ , inclination  $i$ , argument of perigee  $\omega$ , and right ascension of the ascending node  $\Omega$ . To distinguish between the two sets of elements, we shall attach a subscript "0" to those of theory A. Although, strictly speaking, the formulas introduced below are valid only if the mean elements are constants or linear functions of time, they should also be sufficiently good approximations if the mean elements are allowed to contain long-period terms, as is the case with the SAO mean elements.

For convenience,  $M$ ,  $\omega$ , and  $\Omega$  have been so defined that

$$M = M_0, \quad \omega = \omega_0, \quad \Omega = \Omega_0 \quad . \quad (1)$$

However, internally, theory B utilizes a set of elements  $g$  and  $h$  that differ from  $\omega$  and  $\Omega$  by the amount of rotation that the intermediate orbit is undergoing, viz. ,

$$\omega = g + g_{21}(g + M) \quad , \quad \Omega = h + g_{32}(\omega + M) \quad , \quad (2)$$

where

$$g_{21} = -\frac{3}{4} \gamma (1 - 5c^2) - \frac{1}{64} \gamma^2 (41 + 30c^2 - 135c^4) + 0(\gamma^3) ,$$

$$g_{32} = -\frac{3}{16} c [8\gamma + \gamma^2(7 - 33c^2)] + 0(\gamma^3) . \quad (3)$$

Here and in the following, we have used the notation

$$c = \cos i , \quad n = \sqrt{\mu/a^3} , \quad \eta = \sqrt{1 - e^2} , \quad G = \eta \sqrt{\mu a} ,$$

$$\gamma = \frac{J_2}{a^2 \eta^4} , \quad \gamma_4 = \frac{J_4}{J_2^2} ,$$

where  $a$  is measured in earth radii. The mean rates of change of  $M_0$ ,  $\omega_0$ , and  $\Omega_0$  are given by

$$\begin{aligned} \dot{M}_0 = n_0 - \frac{3}{4} n_0 \gamma_0 \eta_0 \left\{ 1 - 3c_0^2 - \frac{1}{32} \gamma_0 [10(1 - 6c_0^2 + 13c_0^4) \right. \\ \left. - 5(5 - 18c_0^2 + 5c_0^4) e_0^2 + 16\eta_0 (1 - 6c_0^2 + 9c_0^4) \right. \\ \left. - 15\gamma_4 (3 - 30c_0^2 + 35c_0^4) e_0^2] \right\} + 0(\gamma^3) , \end{aligned} \quad (4)$$

$$\begin{aligned} \dot{\omega}_0 = -\frac{3}{4} n_0 \gamma_0 \left\{ 1 - 5c_0^2 + \frac{1}{32} \gamma_0 [2(5 + 43c_0^2)(1 - 5c_0^2) \right. \\ \left. + (25 - 126c_0^2 + 45c_0^4) e_0^2 - 24\eta_0 (1 - 8c_0^2 + 15c_0^4) \right. \\ \left. + 20\gamma_4 (3 - 36c_0^2 + 49c_0^4) + 45\gamma_4 (1 - 14c_0^2 + 21c_0^4) e_0^2] \right\} + 0(\gamma^3) , \end{aligned} \quad (5)$$

and

$$\begin{aligned} \dot{\Omega}_0 = -\frac{3}{2} n_0 \gamma_0 c_0 \left\{ 1 - \frac{1}{16} \gamma_0 [4 - 40c_0^2 - (9 - 5c_0^2) e_0^2 \right. \\ \left. + 12\eta_0 (1 - 3c_0^2) - 5\gamma_4 (3 - 7c_0^2)(2 + 3e_0^2)] \right\} + 0(\gamma^3) . \end{aligned} \quad (6)$$

These are invariant rates, and although expressed differently, they must be equal individually to  $\dot{M}$ ,  $\dot{\omega}$ , and  $\dot{\Omega}$ , which are given by

$$\dot{M} = n + \frac{3}{128} n \gamma^2 \eta \left[ 8(1 - 6c^2 + 5c^4) - 5(5 - 18c^2 + 5c^4) e^2 - 15\gamma_4(3 - 30c^2 + 35c^4) e^2 \right] + 0(\gamma^3) , \quad (7)$$

$$\left. \begin{aligned} \dot{\omega} &= \dot{g} + g_{21}(\dot{g} + \dot{M}) , \\ \dot{g} &= -\frac{1}{128} n \gamma^2 \left[ 44 - 300c^4 + (75 - 378c^2 + 135c^4) e^2 + 60\gamma_4(3 - 36c^2 + 49c^4) + 135\gamma_4(1 - 14c^2 + 21c^4) e^2 \right] + 0(\gamma^3) , \end{aligned} \right\} (8)$$

and

$$\left. \begin{aligned} \dot{\Omega} &= \dot{h} + g_{32}(\dot{\omega} + \dot{M}) , \\ \dot{h} &= \frac{3}{32} n c \gamma^2 [2 - 10c^2 - (9 - 5c^2) e^2 - 5\gamma_4(3 - 7c^2)(2 + 3e^2)] + 0(\gamma^3) . \end{aligned} \right\} (9)$$

Aksnes (1970) has shown that the elements  $a$ ,  $e$ , and  $i$  relate to  $a_0$ ,  $e_0$ , and  $i_0$  through the following equations:

$$\frac{1}{a} = \frac{1}{a_0} \left\{ 1 - \frac{1}{2} \eta_0 \gamma_0 (1 - 3c_0^2) + \frac{1}{32} \eta_0 \gamma_0^2 \left[ 1 + 6\eta_0 - (6 + 36\eta_0) c_0^2 + (45 + 54\eta_0) c_0^4 \right] \right\} + 0(\gamma^3) , \quad (10)$$

$$G = G_0 \left[ 1 + \frac{1}{4} \gamma_0 (1 - 3c_0^2) \right] + 0(\gamma^2) , \quad (11)$$

and

$$c = c_0 \left[ 1 + \frac{3}{4} \gamma_0 (1 - c_0^2) \right] + 0(\gamma^2) . \quad (12)$$

While the third-order parts of equations (3) to (9) are available in the cited literature, the terms beyond the second order in equation (10) and the first order in equations (11) and (12) are not known.

For conversion between the two sets of mean elements, equations (1) and (10) to (12) will suffice. In view of the importance of these equations, we have tested them on the orbits of three actual satellites (Table 1). The first two sets of elements have been derived by fitting theory A (line 1) and theory B (line 2) to a series of mostly very accurate laser observations in the manner described by Gaposchkin and Mendes ("Orbital Elements from ISAGEX Data," this volume). The third line shows the elements predicted for theory B by means of the above-mentioned equations and the elements on line 1. The agreement between the last two sets of elements is very good and well within standard errors. The two theories agree on the computed ranges to a few tenths of a meter.

Table 1. Comparison of mean elements for three satellites at epoch 41080.0.

Satellite	Theory	a	e	i	M	$\omega$	$\Omega$
6701401, 261 observations	A	1.1925238	0.0838280	39° 44949	33° 64344	194° 54493	134° 43624
	B	1.1921616	0.0838340	39.43323	33.64344	194.54491	134.43625
	B (Pred.)	1.1921616	0.0838329	39.43321	33.64344	194.54493	134.43624
6508901, 241 observations	A	1.2656913	0.0721680	59.38145	294.47892	24.26520	265.05593
	B	1.2657869	0.0721670	59.36859	294.47892	24.26519	265.05594
	B (Pred.)	1.2657869	0.0721664	59.36859	294.47892	24.26520	265.05593
6800201, 138 observations	A	1.2080431	0.0320070	105.80768	149.20308	2.42648	301.93324
	B	1.2083920	0.0320040	105.81605	149.20308	2.42648	301.93324
	B (Pred.)	1.2083920	0.0320045	105.81605	149.20308	2.42648	301.93324

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**ABBREVIATIONS USED IN THIS REPORT**

BIH	Bureau International de l'Heure
CNES	Centre National d'Etudes Spatiales
EPSOC	Earth Physics Satellite Observation Campaign
GOCC	Geodetic Operations Control Center
GSFC	Goddard Space Flight Center
IGY	International Geophysical Year
ISAGEX	International Satellite Geodesy Experiment
NASA	National Aeronautics and Space Administration
NTU	National Technical University, Athens, Greece
PMT	Photomultiplier Tube
SAO	Smithsonian Astrophysical Observatory
USNO	United States Naval Observatory

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